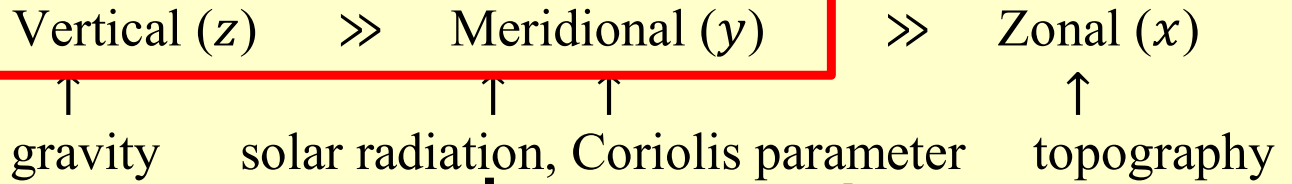


4. Mean zonal-meridional circulations: 4.1-a. Differential heating

Variability of the atmosphere:



stronger diurnal ← → weaker annual

Meridional structure (geographic):

tropics, temperate, frigid ↓
0 ← → stronger

(Eulerian-) zonal mean and disturbance: (λ : longitude, $a = 6.37 \times 10^6$ m: Earth's radius)

$$\overline{(\quad)} \equiv \frac{1}{2\pi} \int_0^{2\pi} (\quad) d\lambda = \frac{1}{2\pi a} \int_0^{2\pi a} (\quad) dx, \quad (\quad)' \equiv (\quad) - \overline{(\quad)} \quad (22)$$

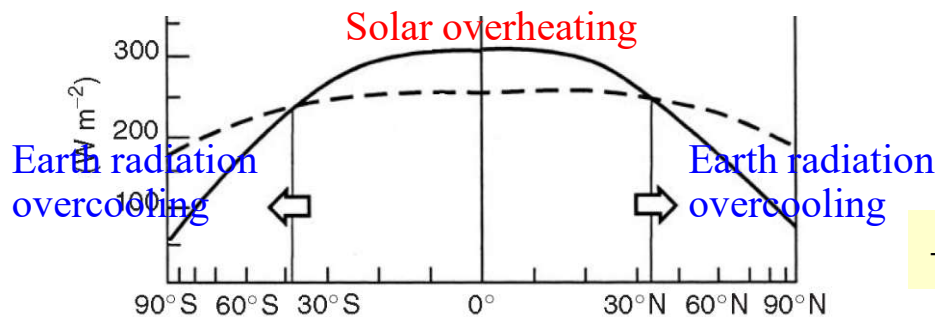


Figure 7 Annual-mean net incoming solar radiation (solid line) and outgoing terrestrial radiation (dashed line) as a function of latitude, expressed in units of $W m^{-2}$. Distance on the latitude scale is proportional to area on the Earth's surface. (Diagram provided by Socorro Medina.)

→ Poleward atmospheric and oceanic heat transport

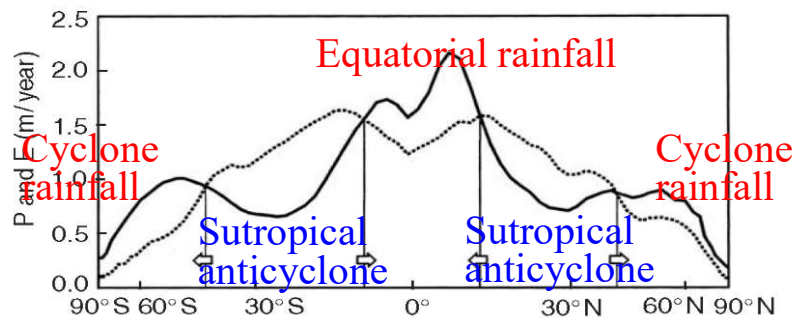
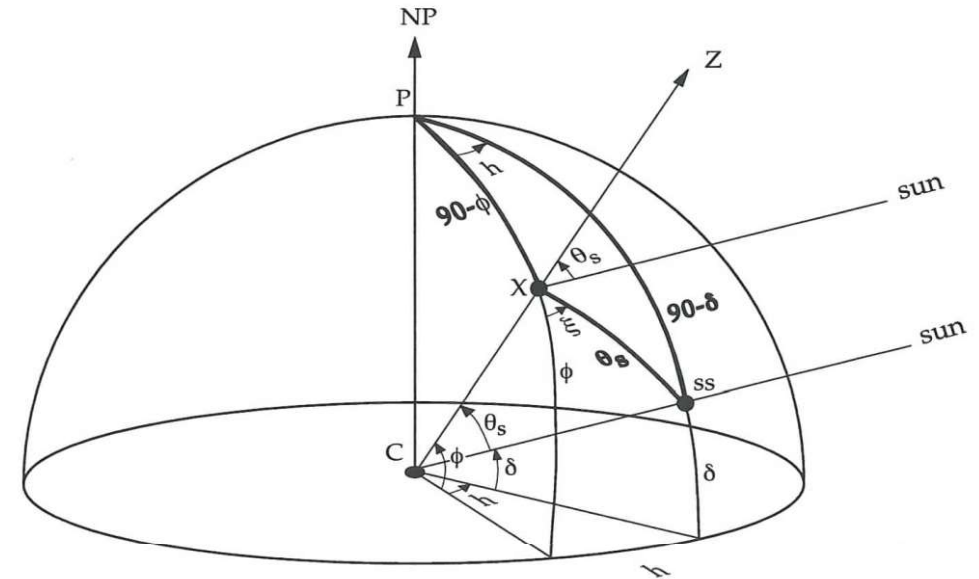
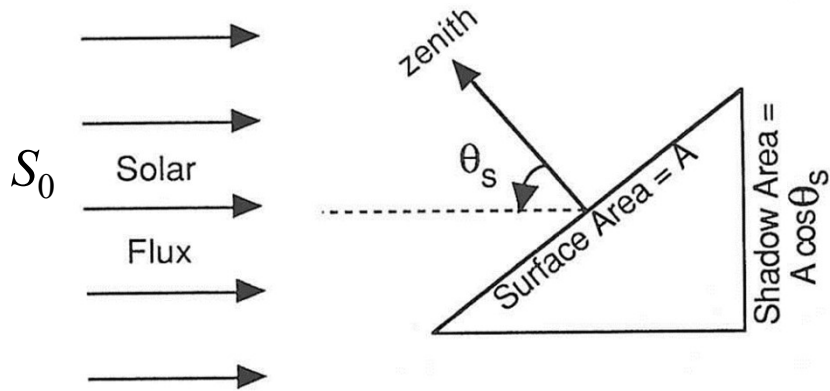


Figure 10 Annual-mean precipitation (solid) and evaporation (dashed) as a function of latitude, expressed in units of meters per year. Distance on the latitude scale is proportional to area on the Earth's surface. Based on NCEP/NCAR Reanalyses for the period 1958–1997. (Diagram provided by Socorro Medina.)

→ Equatorward water vapor transport

2-dimensional climate: Meridional-seasonal



$$Q = S_0 \left(\frac{\bar{d}}{d} \right)^2 \cos \theta_s$$

(d : Solar distance, given by Kepler's equation)

$$\cos \theta_s = \cos(90 - \phi) \cos(90 - \delta) + \sin(90 - \phi) \sin(90 - \delta) \cos h$$

$$\cos \theta_s = \sin \phi \sin \delta + \cos \phi \cos \delta \cos h$$

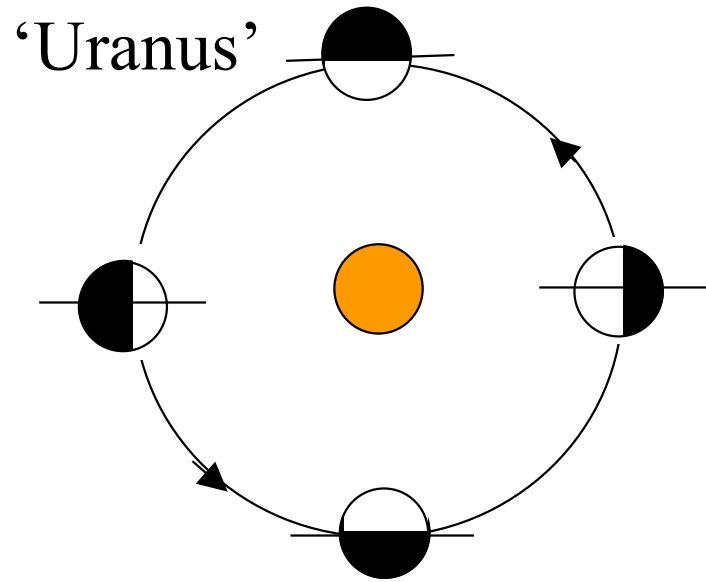
$$\frac{\sin(180 - \xi)}{\sin(90 - \delta)} = \frac{\sin h}{\sin \theta_s}$$

$$\cos h_0 = -\tan \phi \tan \delta$$

$$\sin \xi = \frac{\cos \delta \sin h}{\sin \theta_s}$$

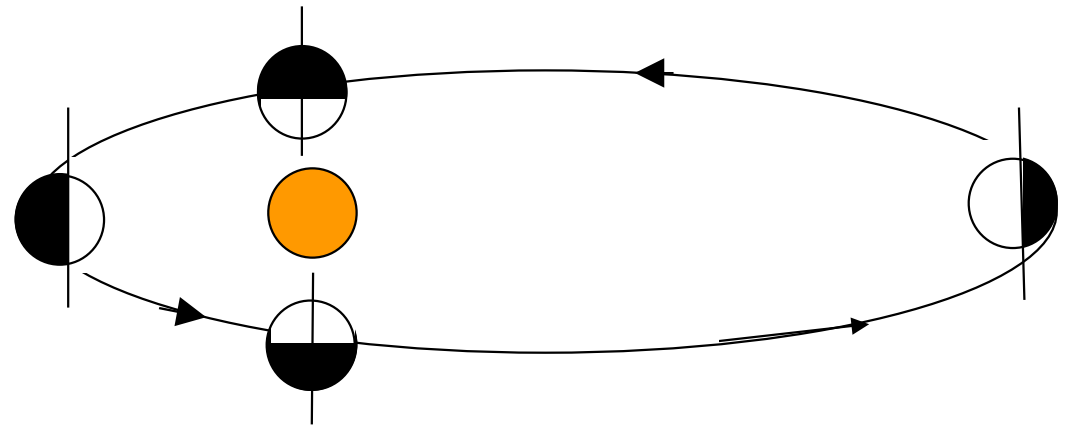
$$\bar{Q}^{\text{day}} = \frac{S_0}{\pi} \left(\frac{\bar{d}}{d} \right)^2 \left[h_0 \sin \phi \sin \delta + \cos \phi \cos \delta \sin h_0 \right]$$

Two extreme cases of seasonal cycle forcing

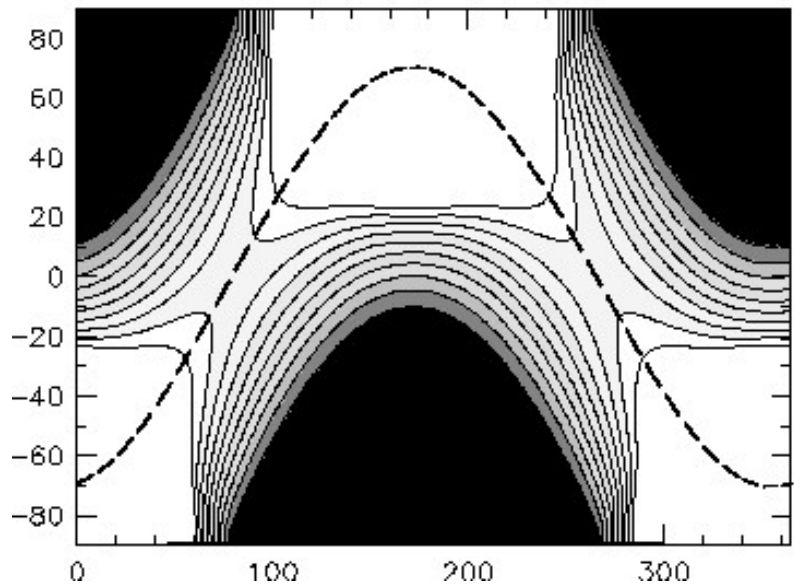


Rotation-axis obliquity

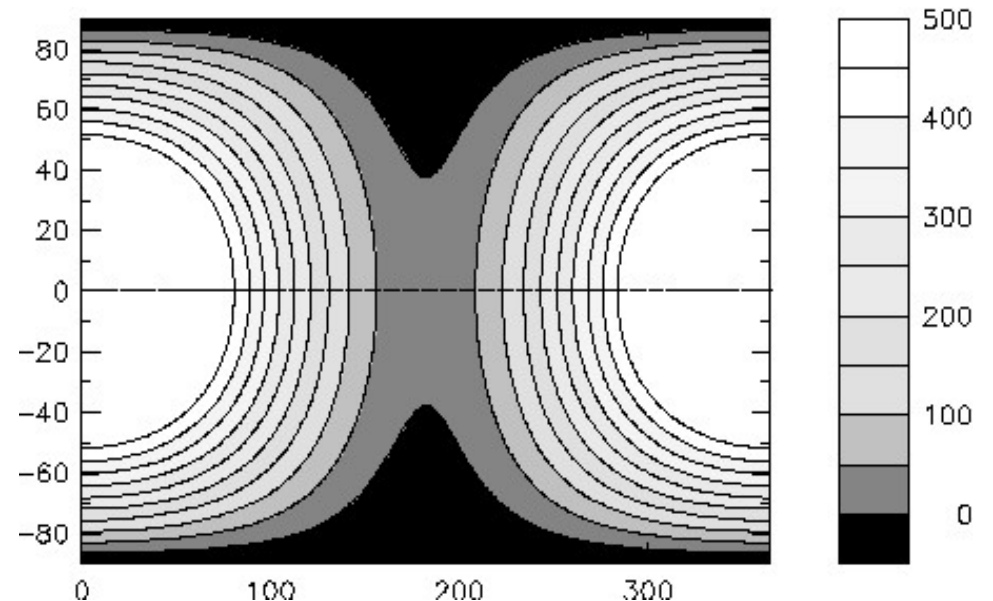
‘a type of extra-solar planet’



Orbital eccentricity

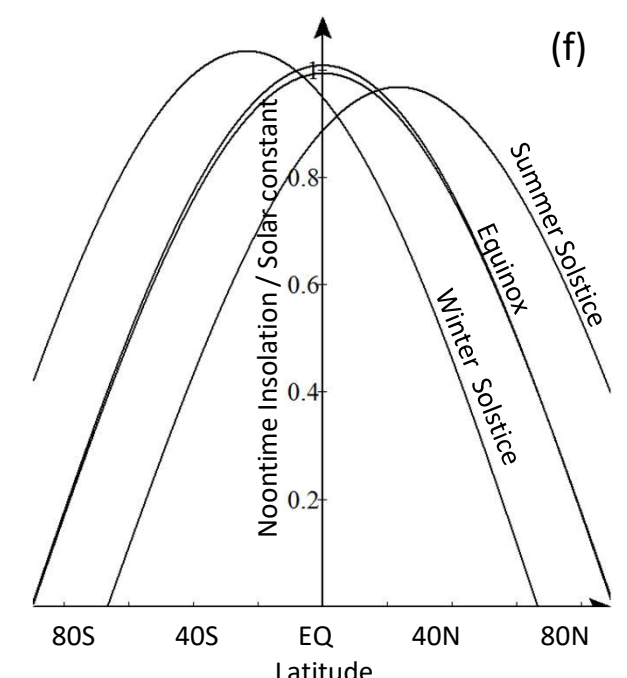
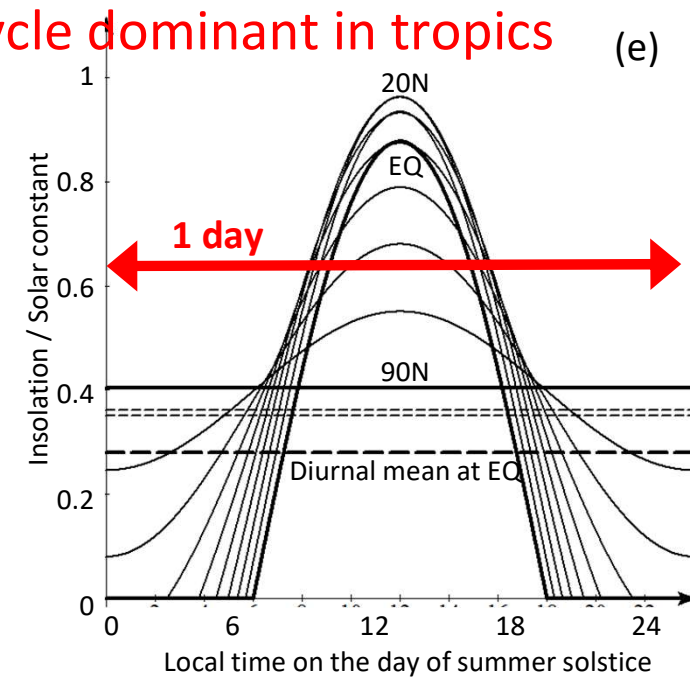
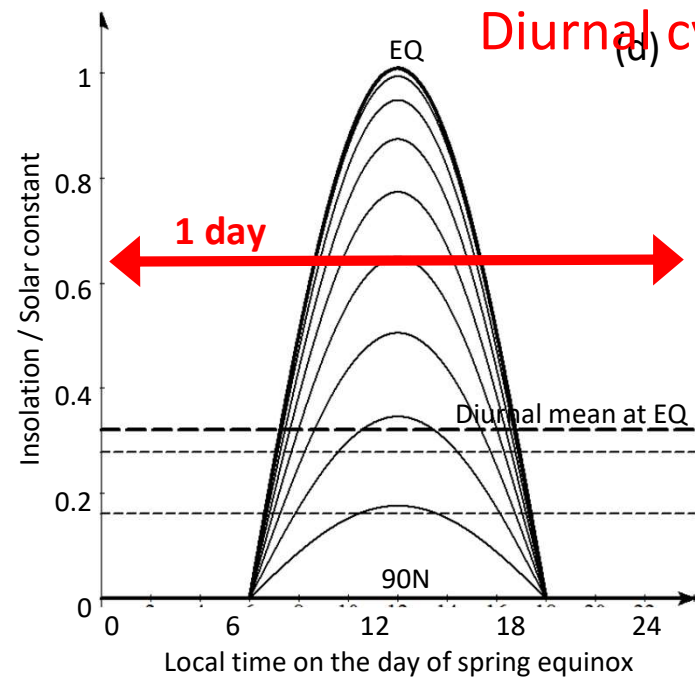
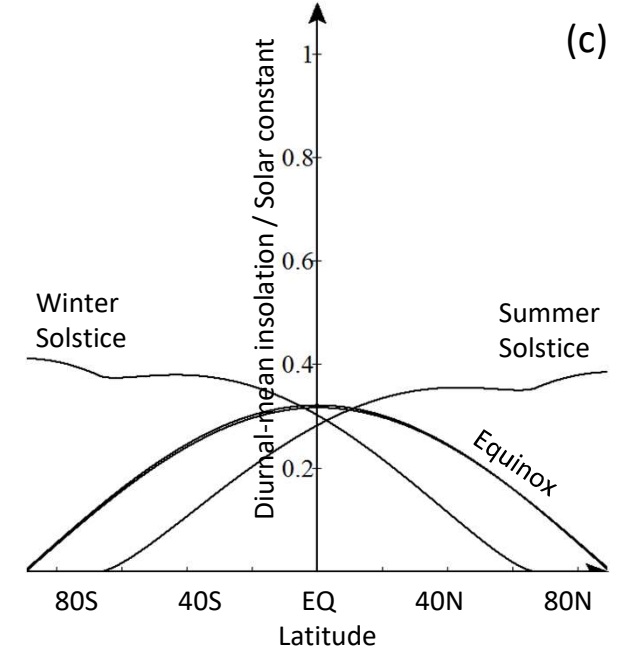
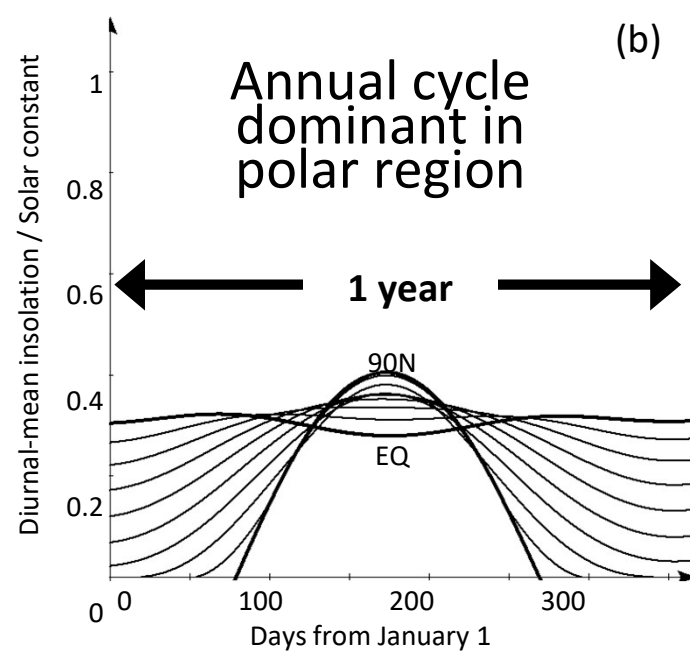
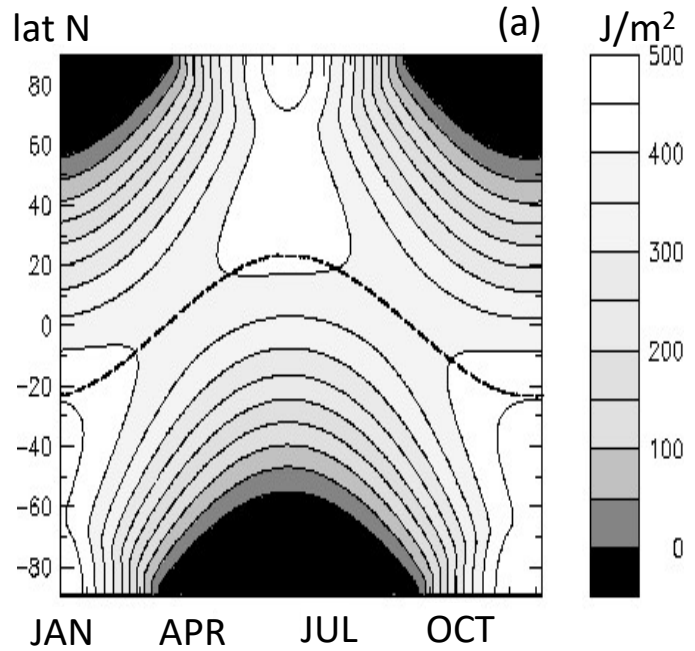


Hemispherically anti-phase



Hemispherically in-phase

Solar heating on earth with revolution and rotation



Why tropics is tropics?

● High latitudes: Low solar angle

- Large reflection (strong albedo)

- Large refraction (long optical depth)

→ Low temperature → Cryosphere/Clouds

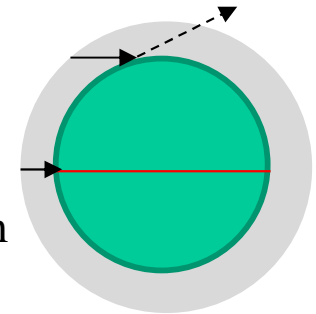
● Rapid rotation &

Slow circular revolution

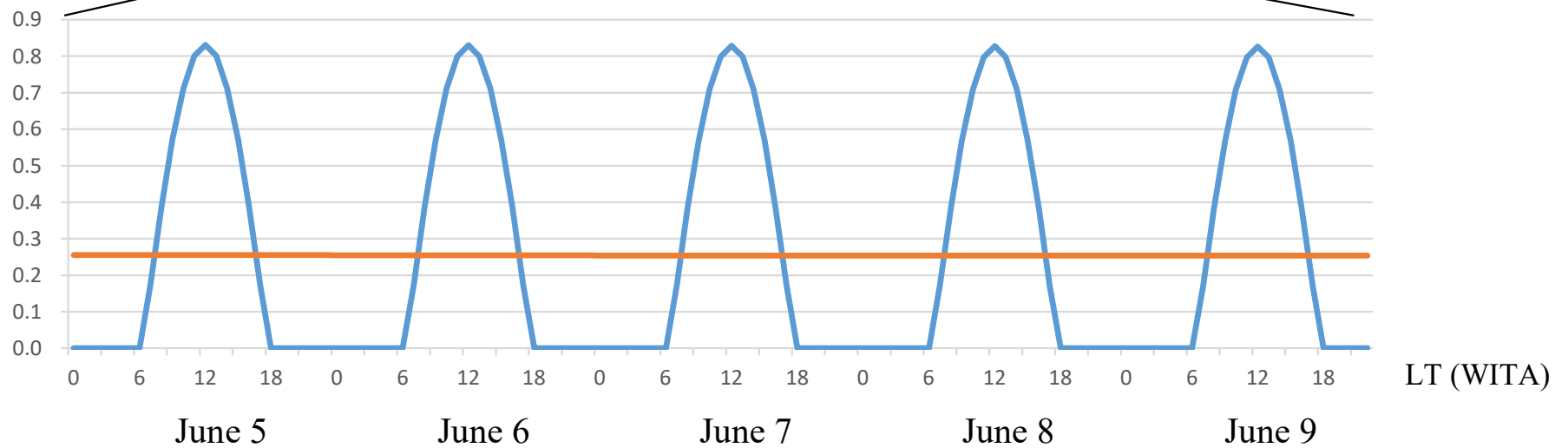
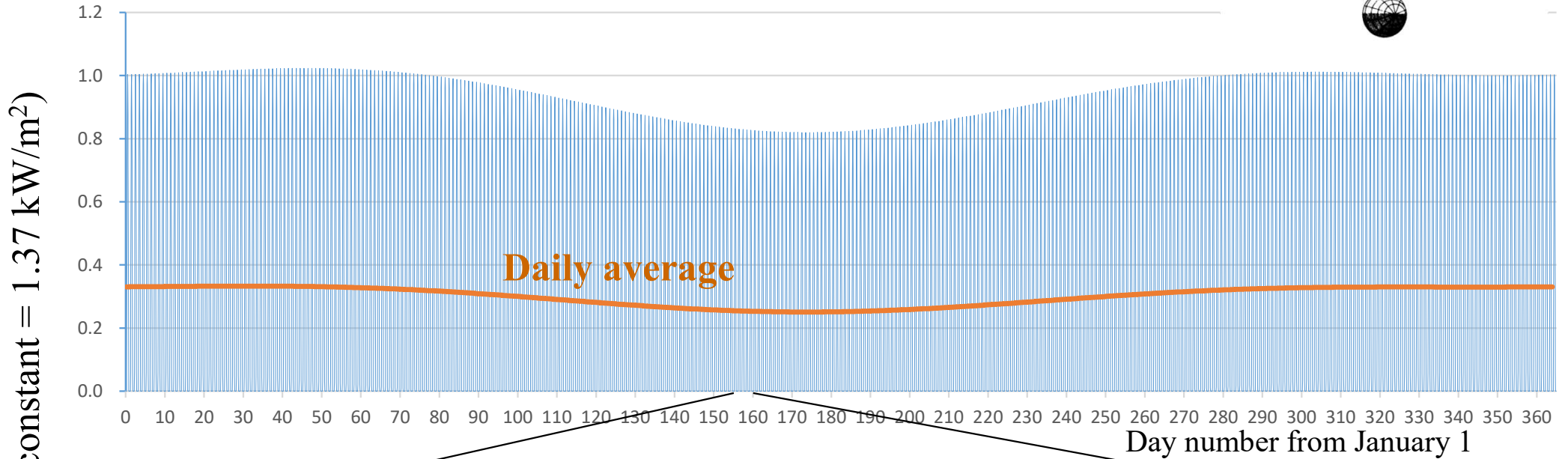
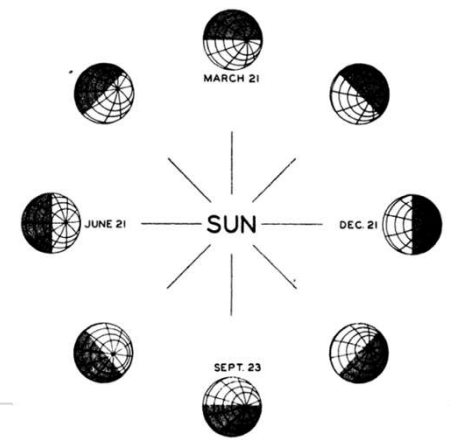
- Zonal homogeneity

- Weak hemispheric anti-symmetry

Low latitude:
High solar angle
Short path length

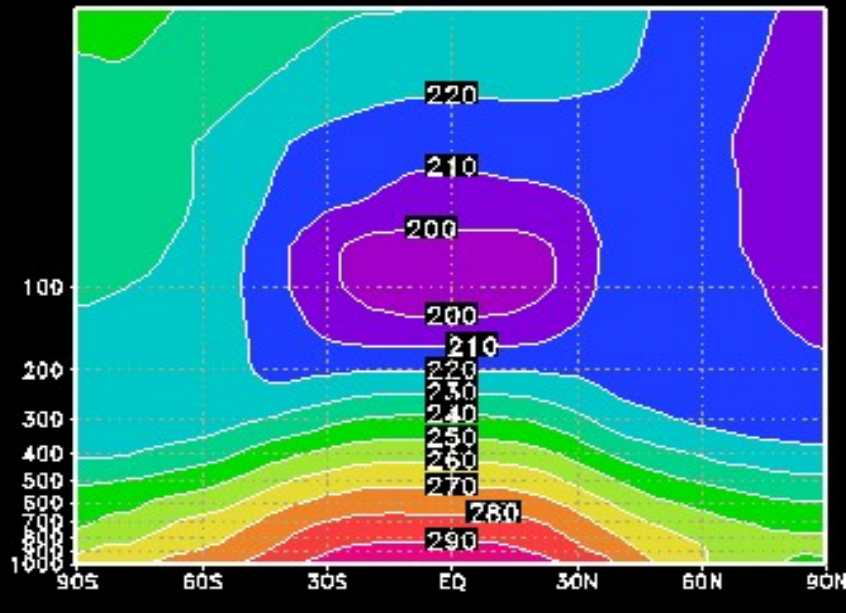


Solar radiation heating at the atmosphere top over Denpasar (8°39'S)

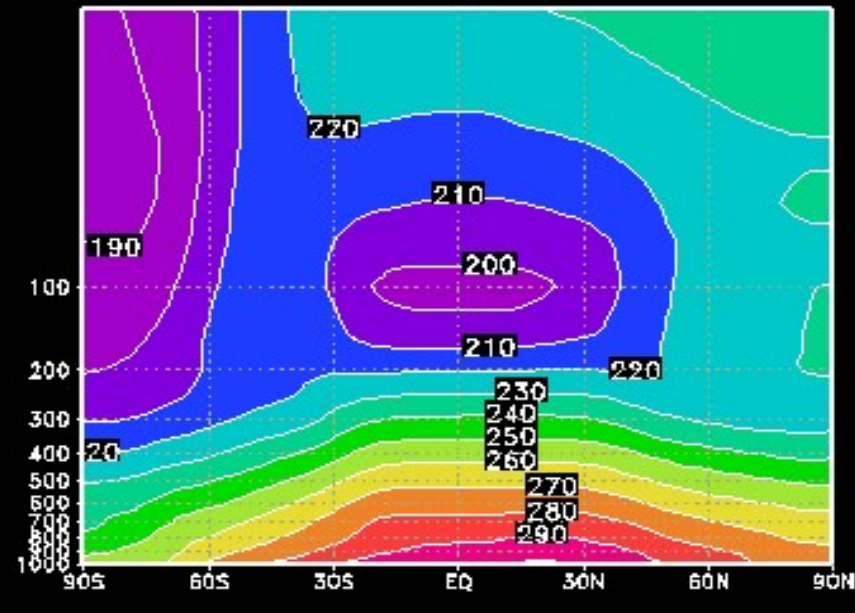


4.1-c Trade wind (Equatorial easterly & mid-latitude westerly)

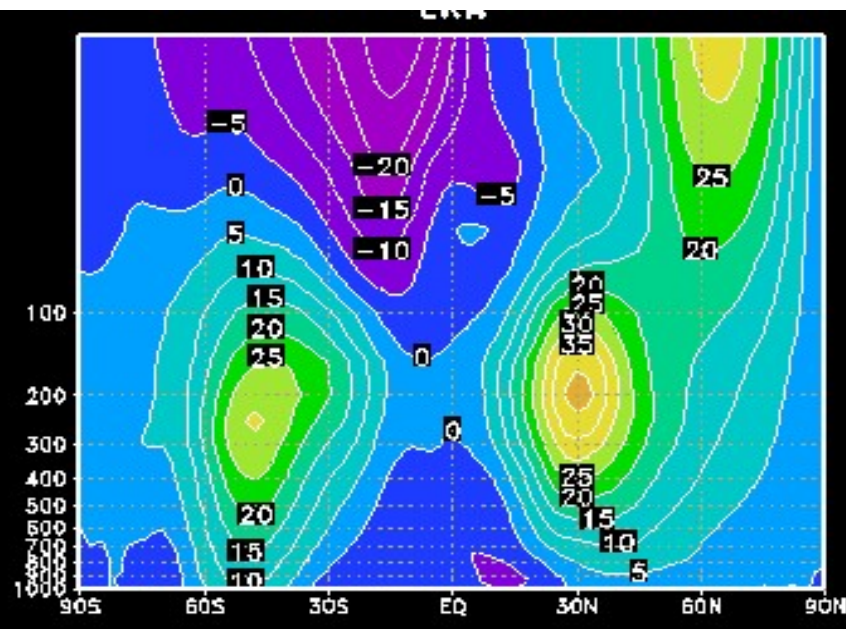
DJF
 \bar{T}



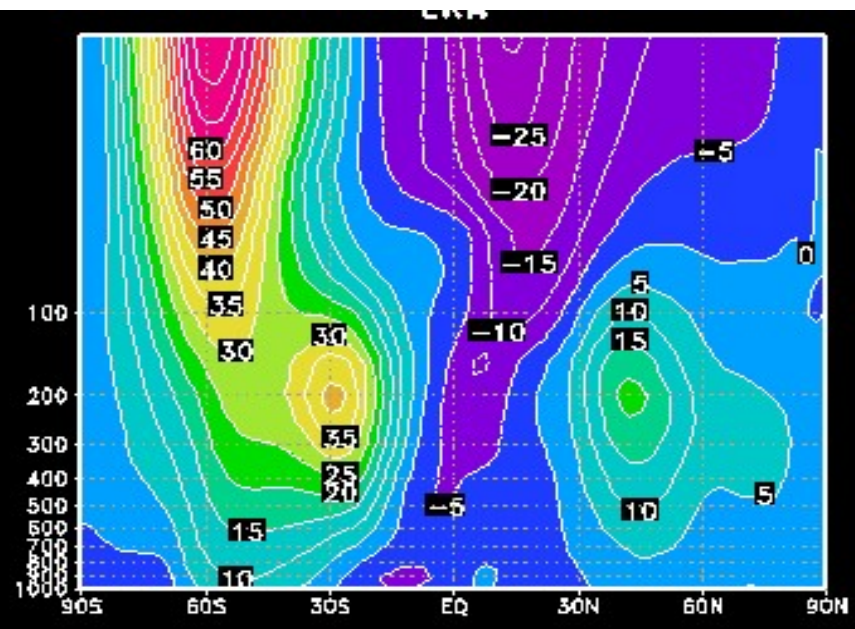
JJA
 \bar{T}



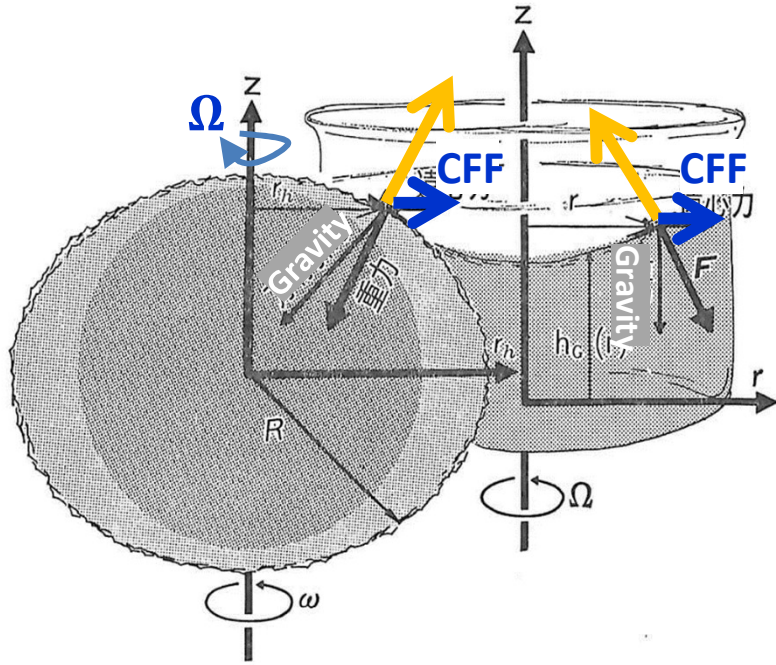
DJF
 \bar{u}



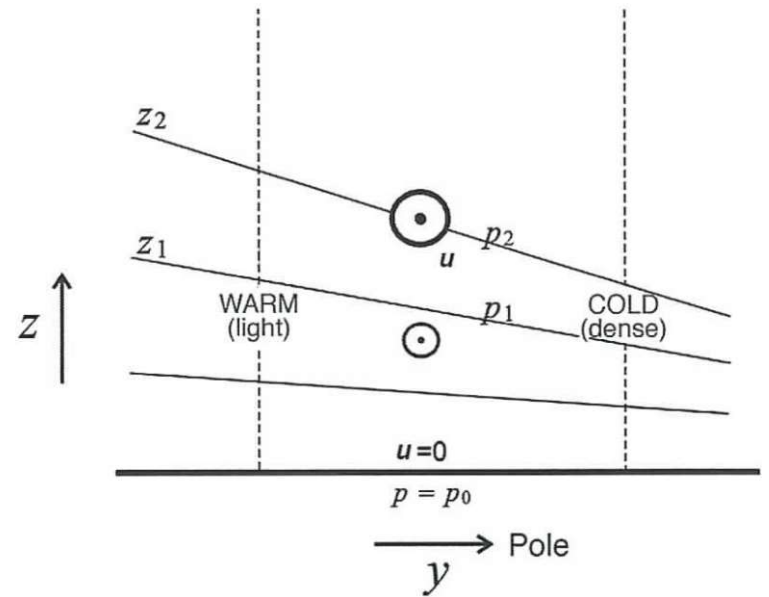
JJA
 \bar{u}



Rotating fluid and centrifugal/Coriolis forces



“Thermal wind” equilibrium



Angular momentum conservation and westerly generation

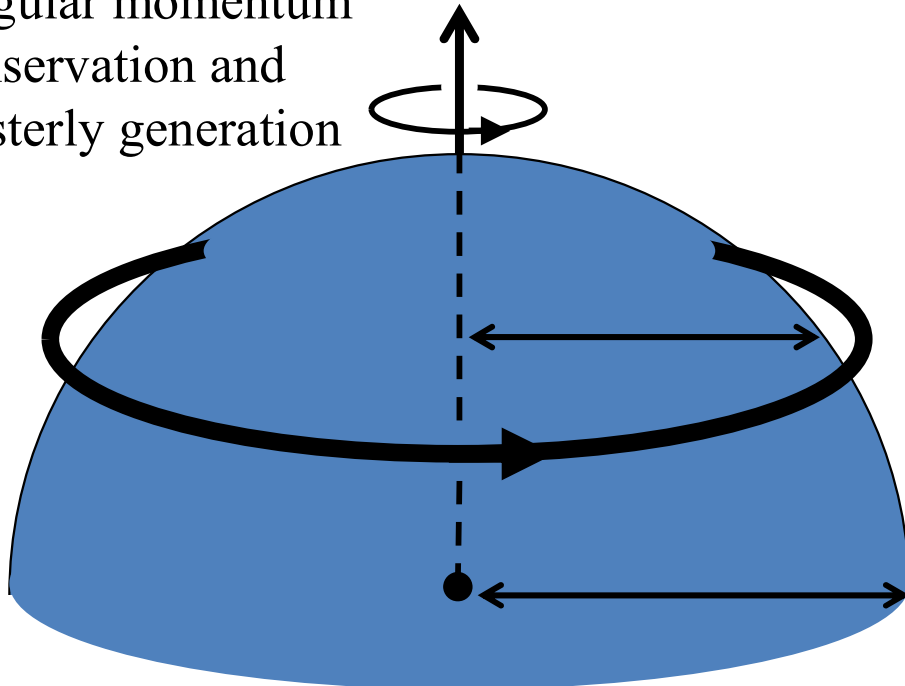


FIGURE 5.14. Warm columns of air expand, cold columns contract, leading to a tilt of pressure surfaces, a tilt which typically increases with height in the troposphere. In Section 7.3, we will see that the corresponding winds are out of the paper, as marked by \odot in the figure.

(Marshall & Plumb, 2009)

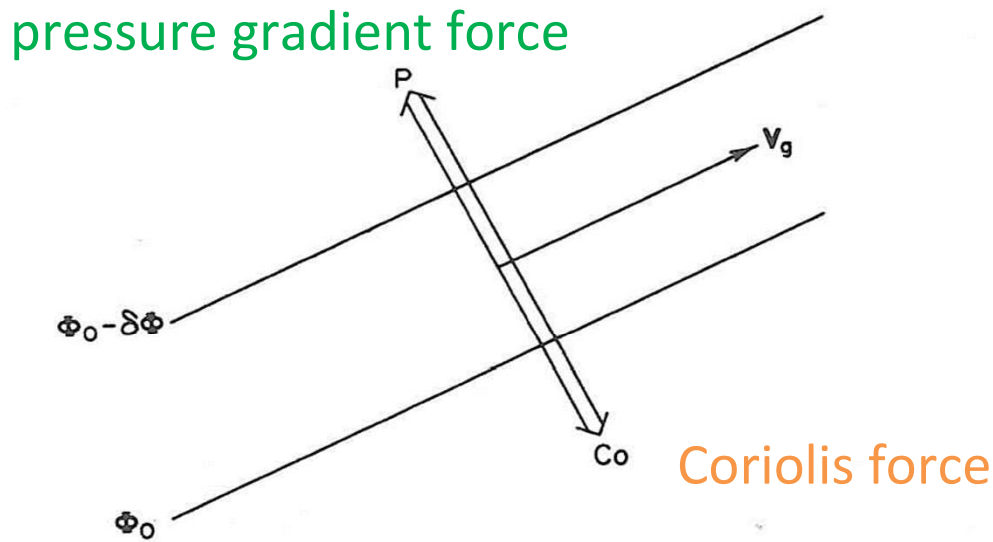


Fig. 3.2 Balance of forces for geostrophic equilibrium. The pressure gradient force is designated by P and the Coriolis force by Co .

(Holton, 1972)

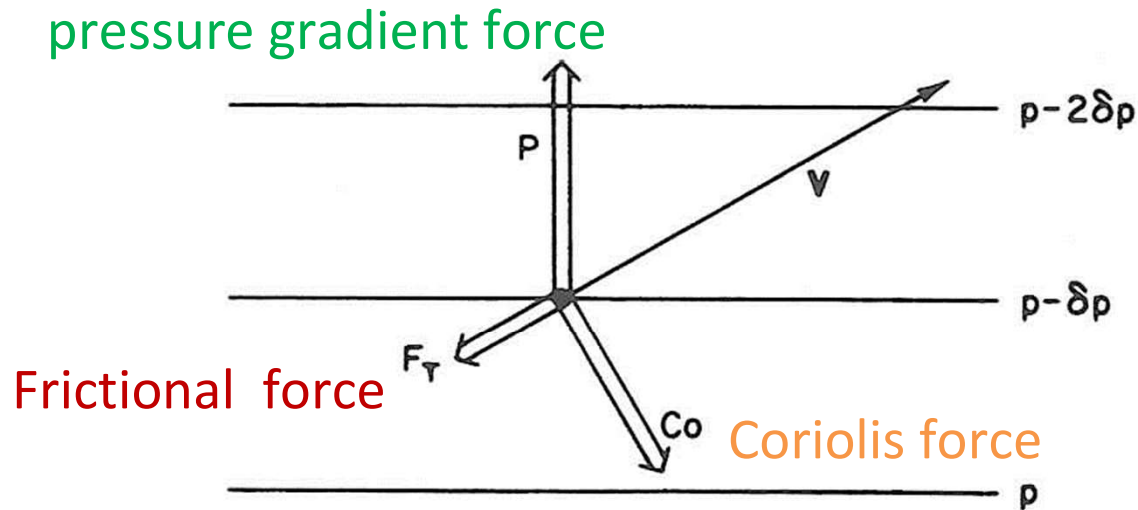


Fig. 5.3 Balance of forces in the well-mixed planetary boundary layer: P designates gradient force, Co the Coriolis force, and F_T the turbulent drag.

Upper ocean
(Marshall & Plumb, 2008)

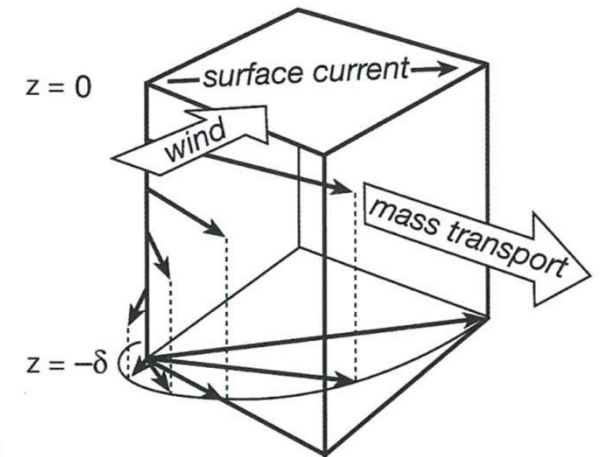
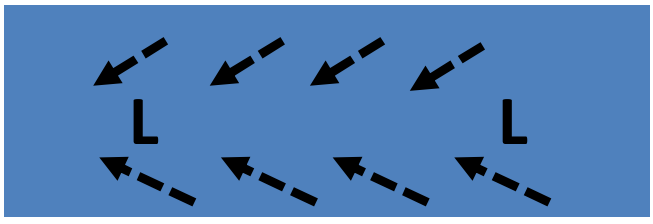
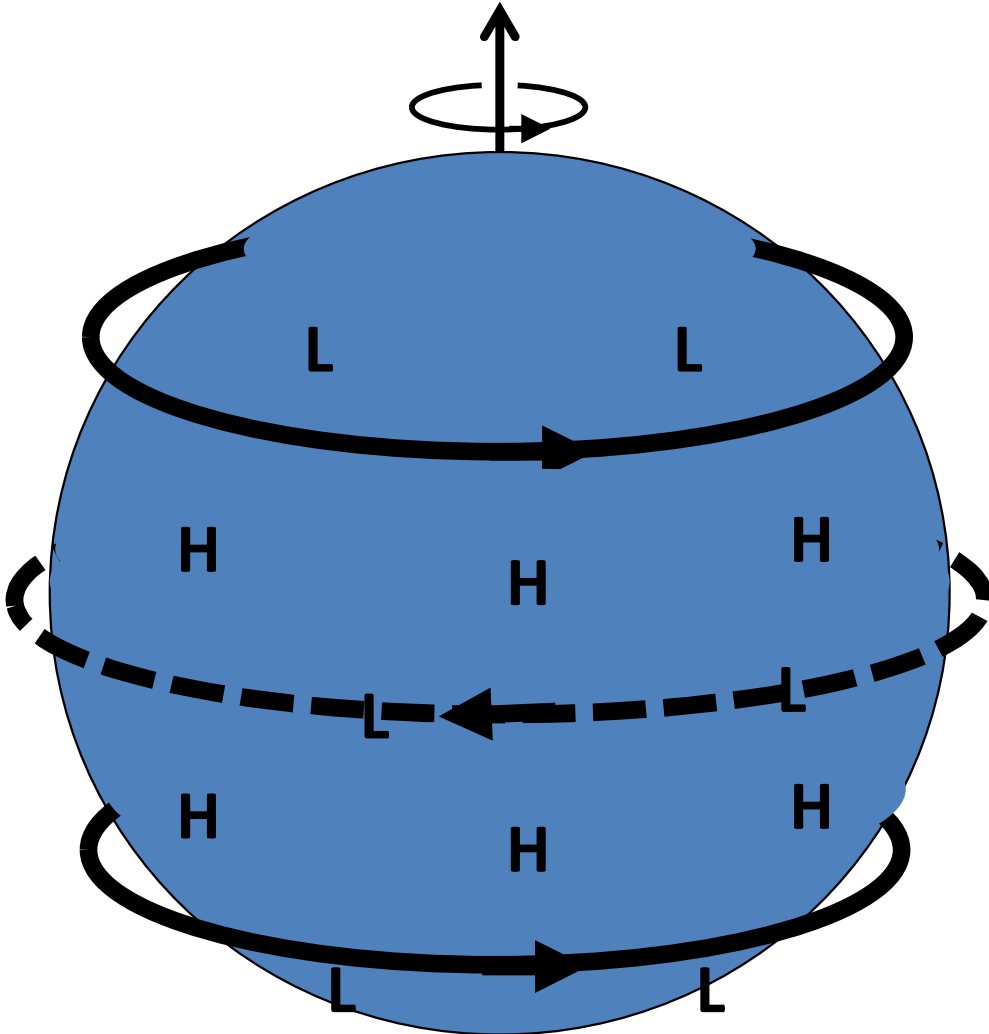


FIGURE 10.5. The mass transport of the Ekman layer is directed to the right of the wind in the northern hemisphere (see Eq. 10-5). Theory suggests that horizontal currents, u_{ag} , within the Ekman layer spiral with depth as shown.

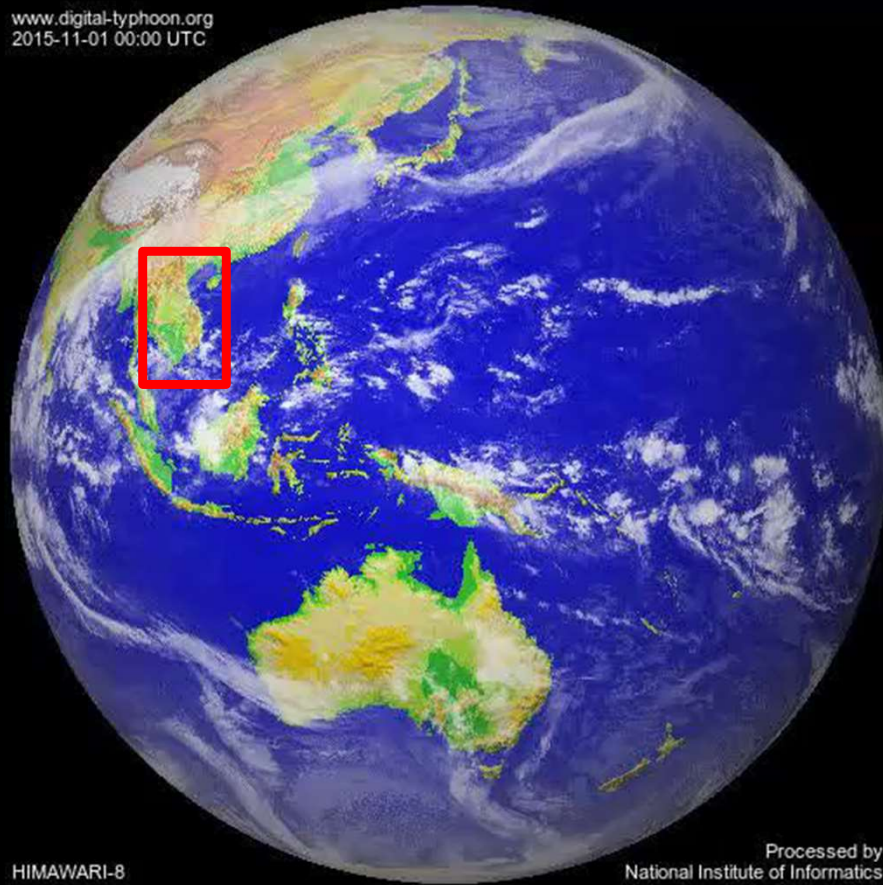
Geostrophic and surface (frictional) flow



“convergence”

Earth \neq “Aqua-Planet”

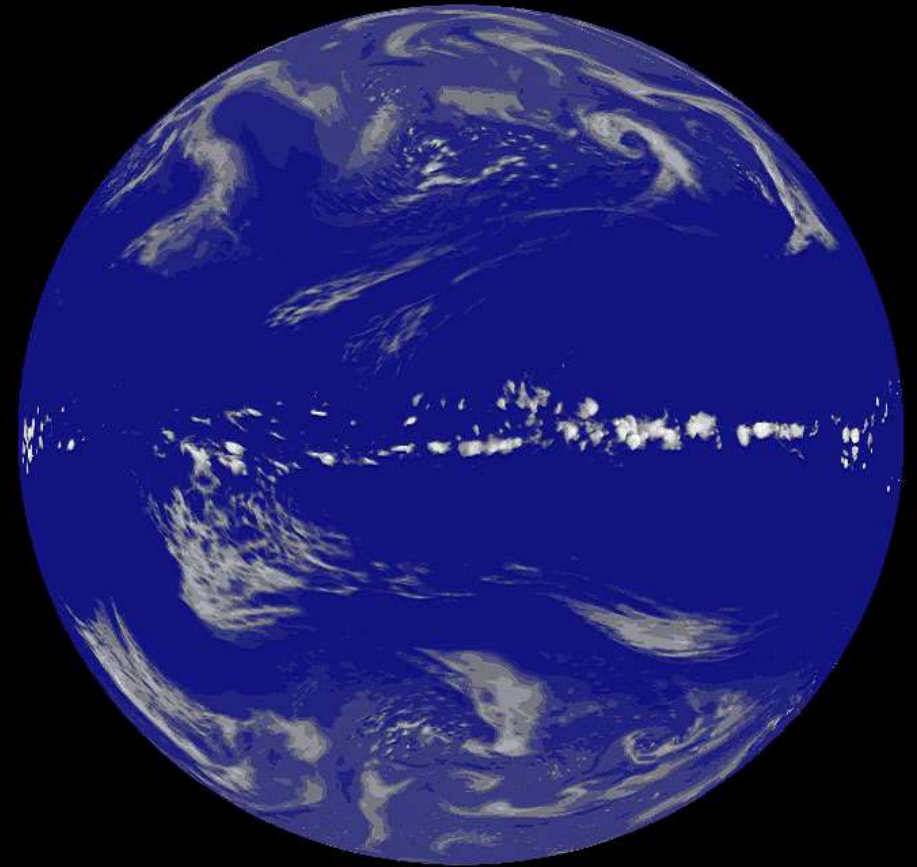
www.digital-typhoon.org
2015-11-01 00:00 UTC



HIMAWARI-8

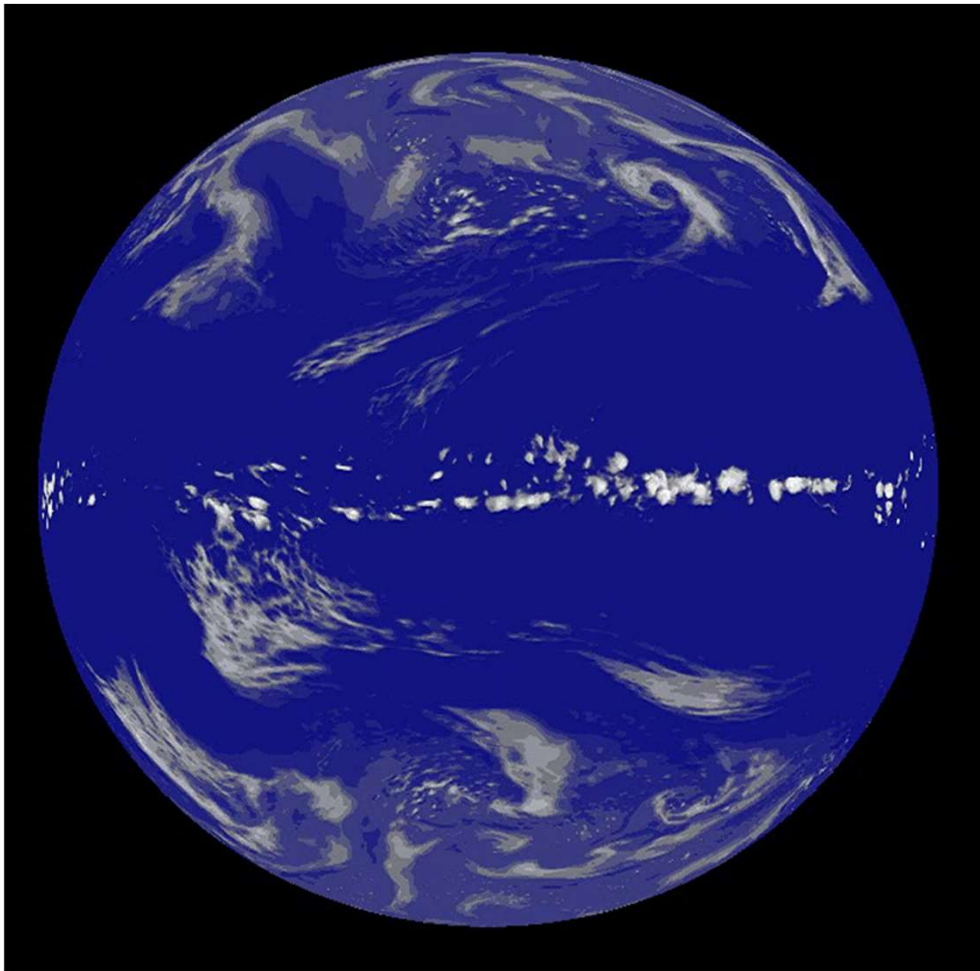
Processed by
National Institute of Informatics

Himawari-8/JMA (1-30 Nov 2015)



NICAM/JAMSTEC (Satoh et al., 2008)

“Virtual Earth” (Aqua-Planet) by the Earth Simulator

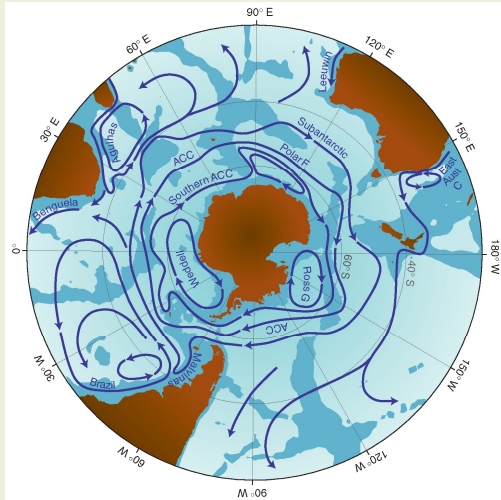


- Imagine the Earth is a complete sphere, then the liquid water covers the whole Earth, that is so-called an “aqua-planet”.
- In such a case the oceanic circulation should be almost zonal (as appeared in stripe patterns of dense atmospheres of Jupiter and Saturn).
- The atmosphere circulation must have meridional circulations dependent on differential solar heating
- Updraft and cloud are generated along the equator, which is observed over open oceans of the actual Earth and called inter-tropical convergence zone (ITCZ).
- More striking feature is induced by ocean-atmosphere interactions, which is so-called intraseasonal variations (ISVs) or super cloud clusters moving eastward with periods of 1-2 months and zonal scales of a few thousand kilometers as observed actually over Indian and Pacific Oceans.

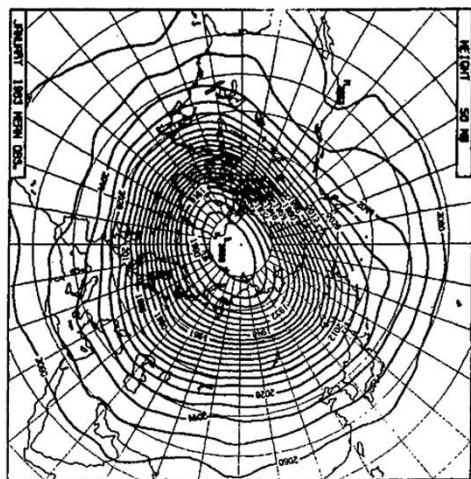
“stripes”
of deep
atmosphere
of Jupiter



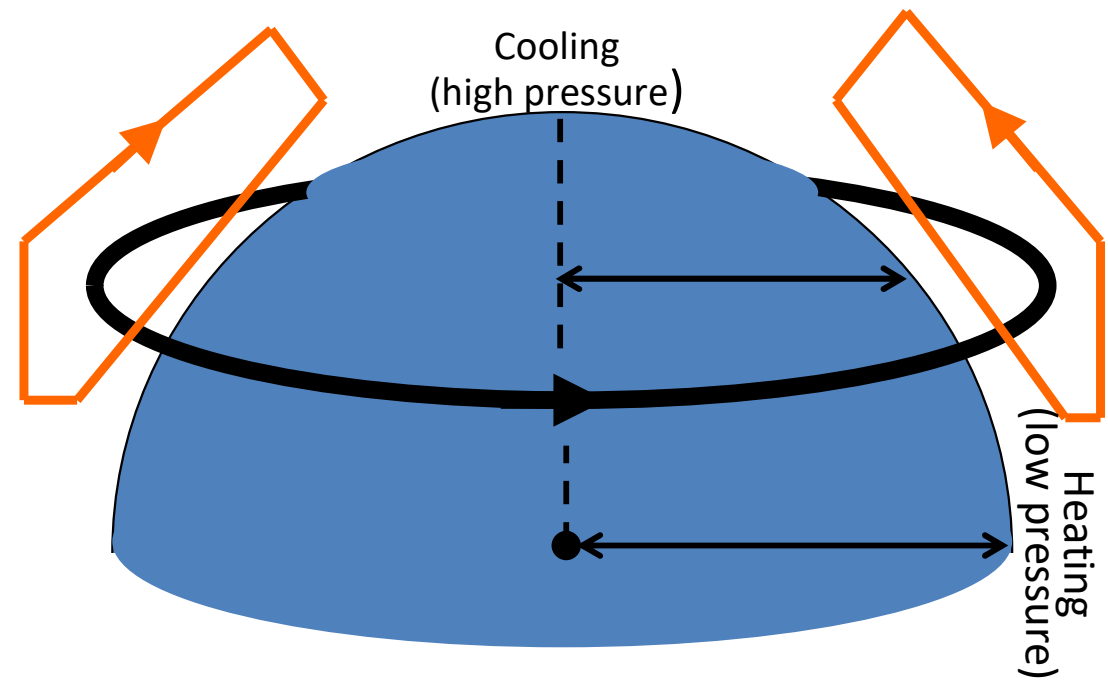
circumpolar
ocean
current
around
Antarctica



Arctic
anticyclone
in summer
stratosphere



“Aqua-planet” zonal vortex with differential heating



Zonal flow (u) \Leftrightarrow Meridional (y) pressure gradient
(y -momentum eq. (geostrophic) : $f u = -\rho^{-1} \partial p / \partial y$)

Meridional flow (v) \Leftrightarrow Zonal (x) friction/drag
(x -momentum eq. (Ekman) : $-f v = -\alpha u$)

Vertical flow (w) \Leftrightarrow Radiative/Latent heating
(thermodynamic eq. (pseudoadiabatic) : $\Gamma w = \alpha' \Delta T$)

Hermann Ludwig Ferdinand von Helmholtz (1821 – 1894)



Quasi-2D (horizontal) flow velocity:

$$\mathbf{u} = \nabla \times (0, 0, \phi) - \nabla \chi$$

(ϕ : stream function; χ : velocity potential)

$$\nabla \cdot (\phi_y, -\phi_x, 0) = 0$$

non-divergent
(solenoidal)

“vortex”

on (quasi-)horizontal plane

↓ $\phi \sim$ geopotential
(geostrophic)

Weather maps

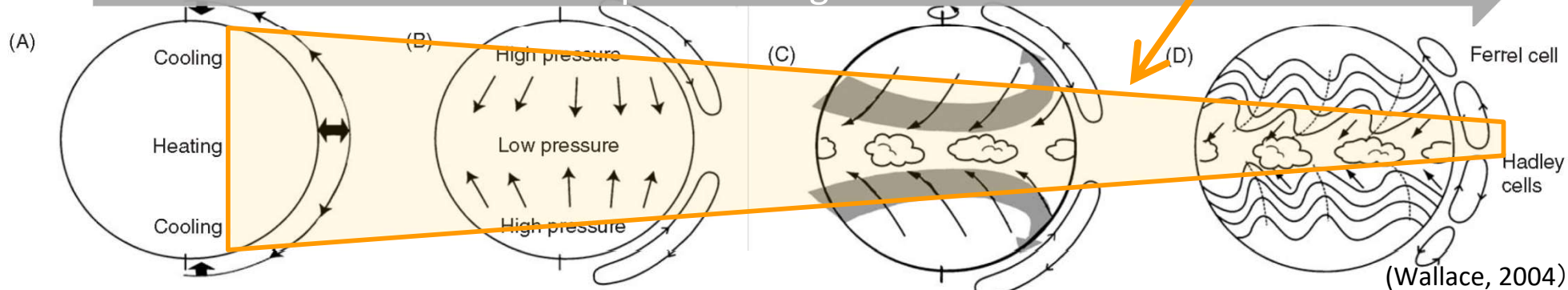
$$\nabla \times (-\nabla \chi) = 0$$

irrotational

“convection”

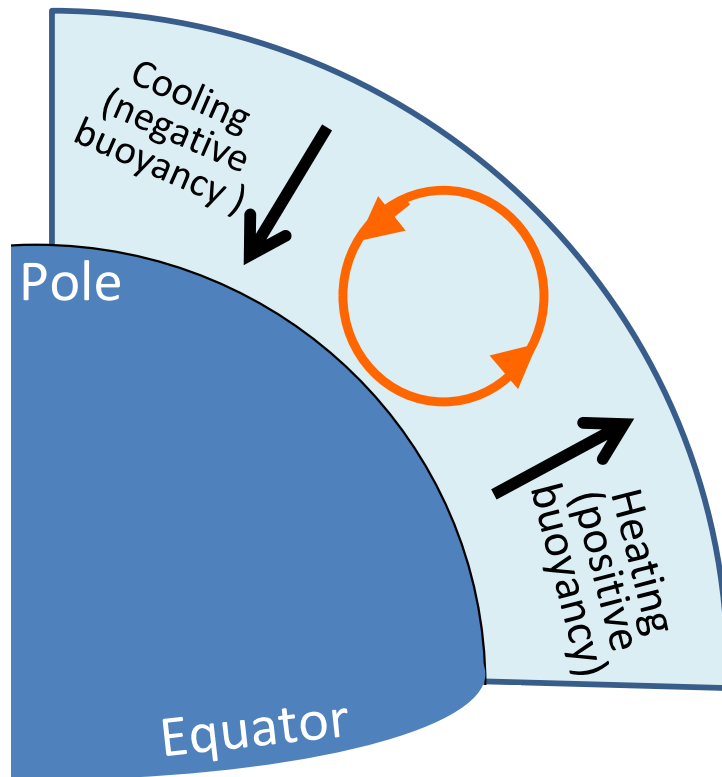
on vertical plane

Rotation makes narrower equatorial region and baroclinic mid-latitude zone.

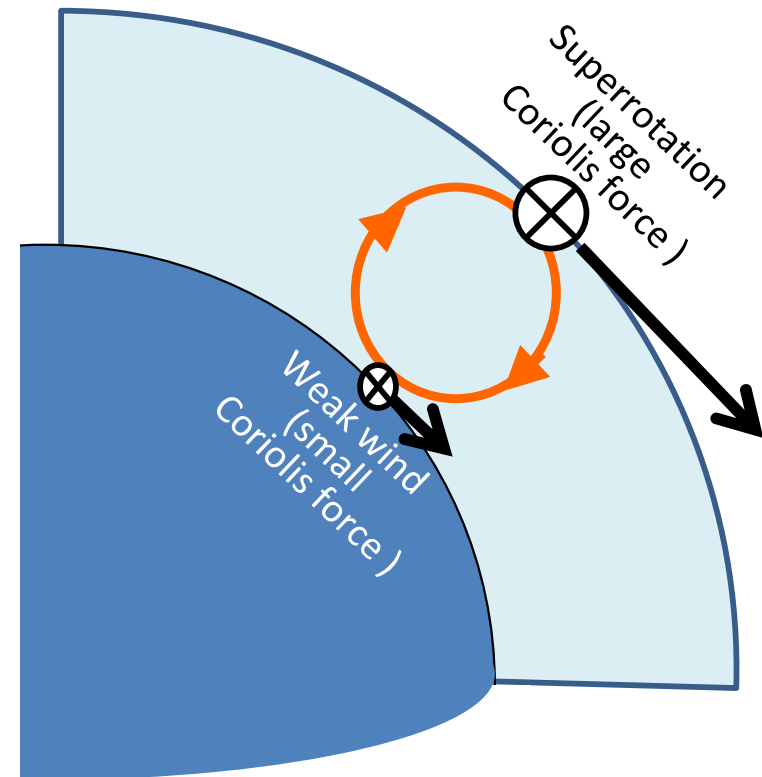


Direct / indirect circulations due to differential heating / shear satisfying by meridional temperature gradient

(a) Buoyancy torque



(b) Centrifugal/Coriolis torque



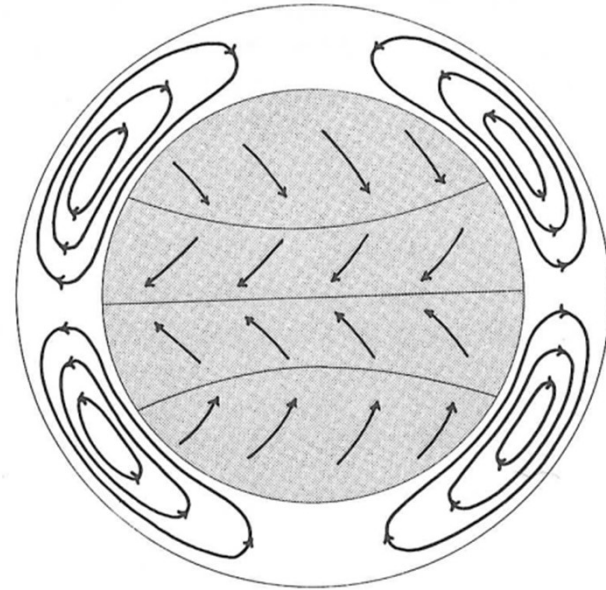
(adapted from Matsuda & Yoden, 1985)

Edmond Halley (1656 – 1742)

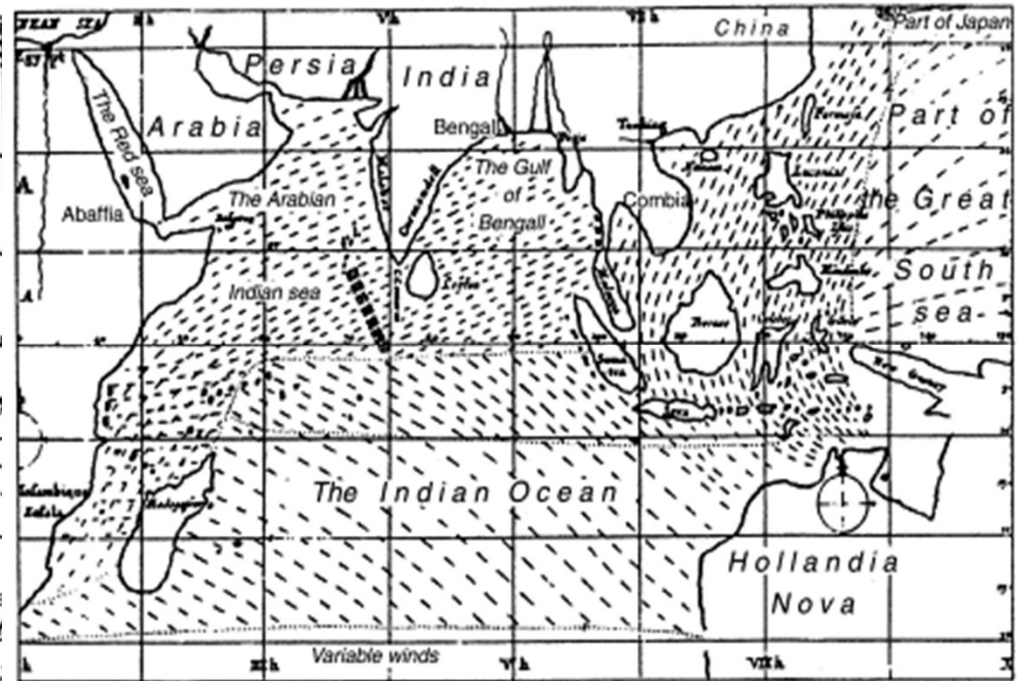
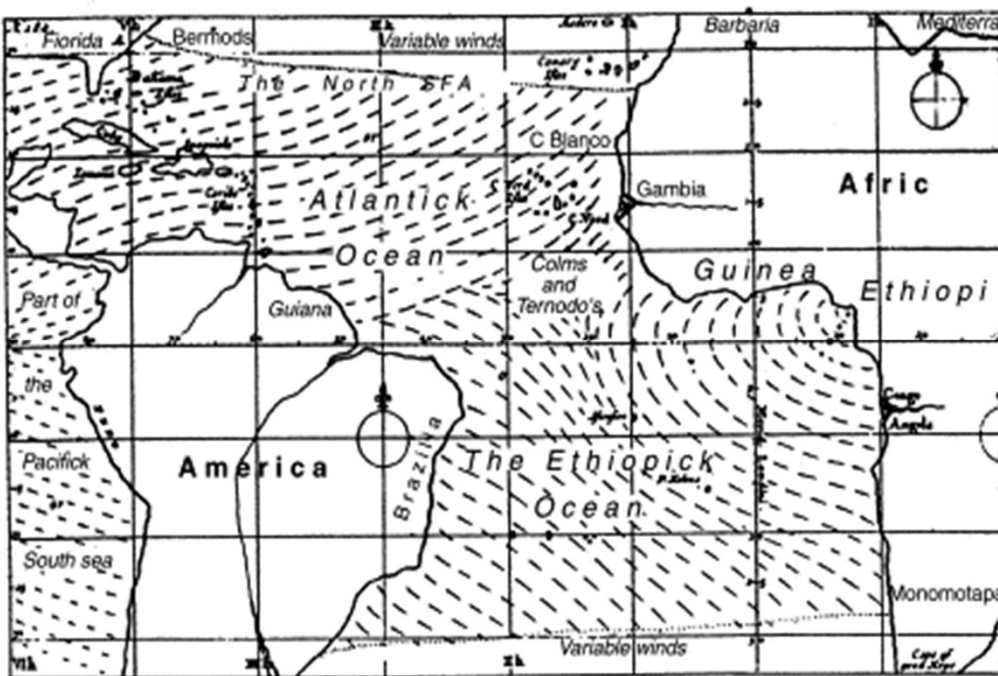


(<http://www.staff.science.uu.nl/~gent0113/astrology/newton.htm>)

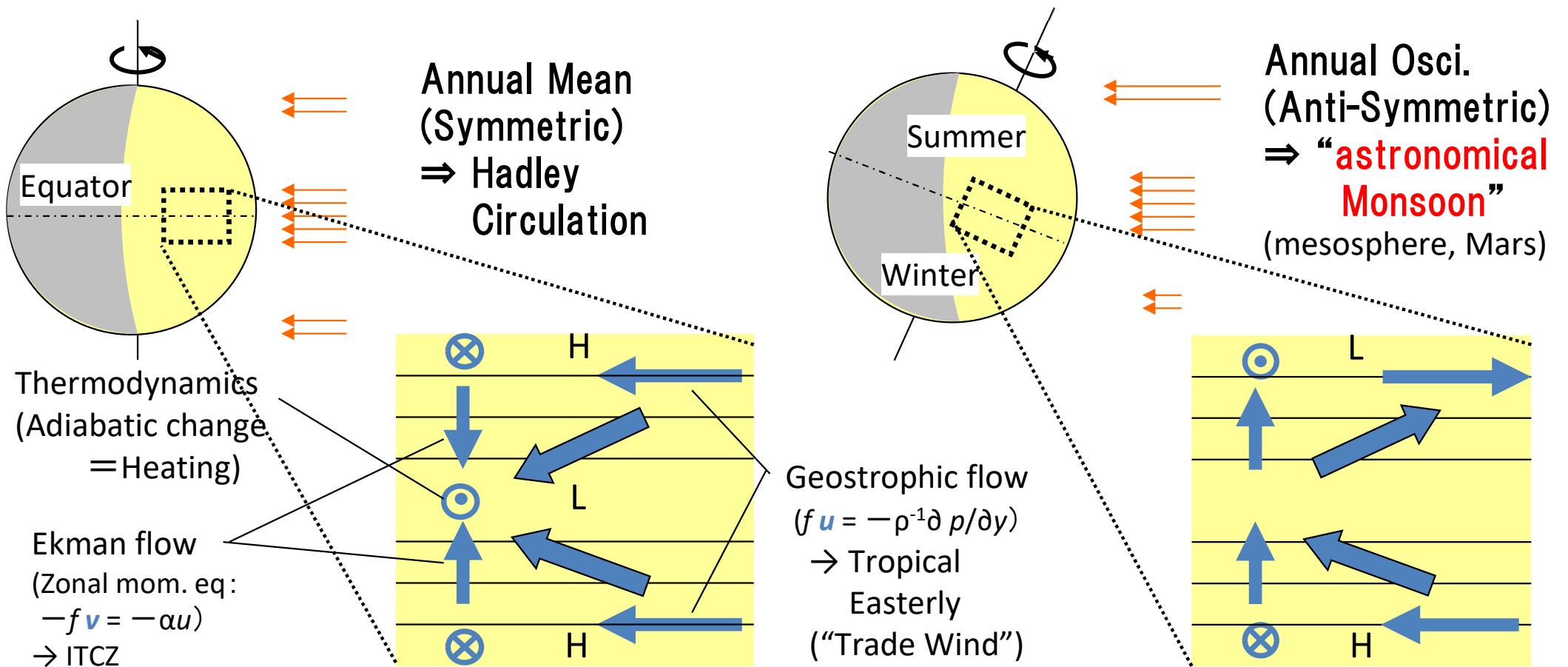
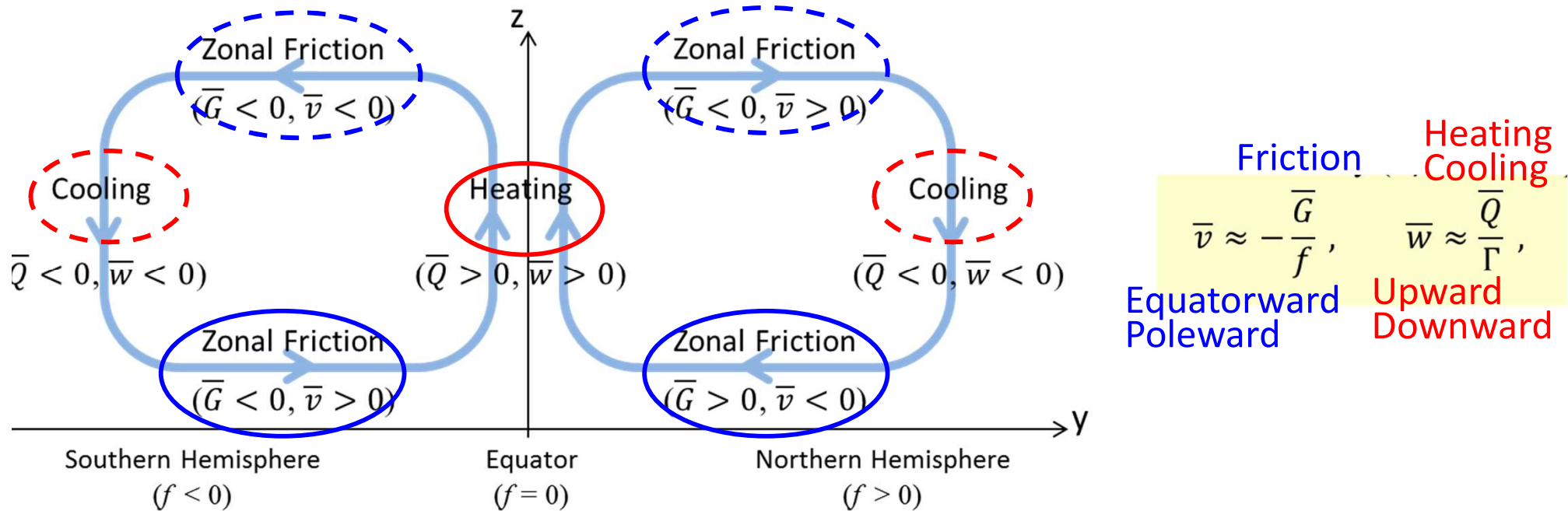
George Hadley (1685 – 1768)



(Hadley, 1735; reproduced by Lorenz, 1967)



(Halley, 1686)



Solar differential heating and monsoon circulation

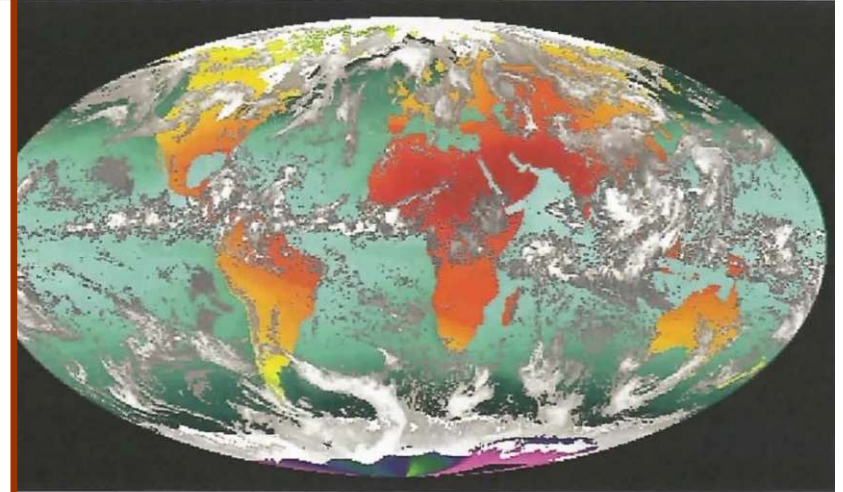
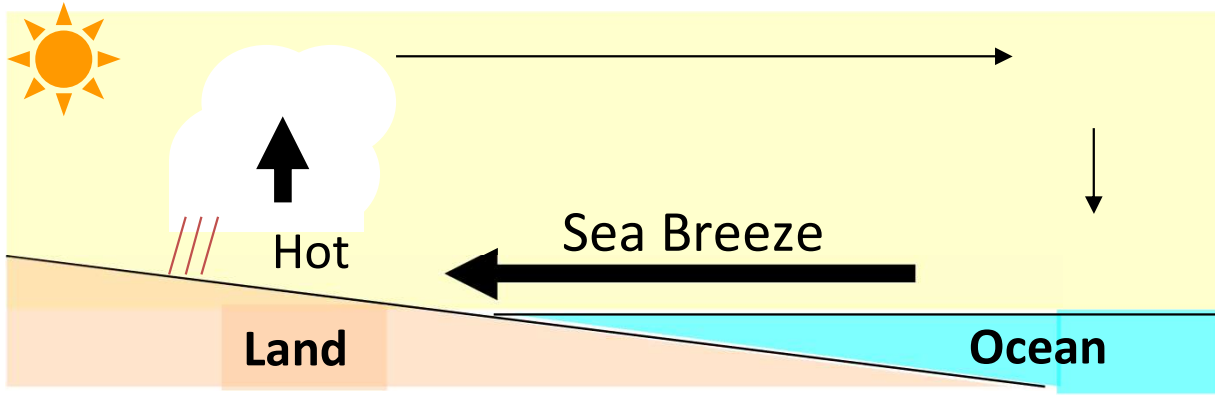
	Global differential heating	Local heat capacity contrast
Rotation (Diurnal)	Day-Night circulation (Tides)	Sea-Land (Mountain-valley) breeze
Revolution (Annual)	Summer-winter circulation (hemispheric anti phase) Perihelion-Aphelion (in phase)	Ocean-Continent (Monsoon)

Monsoon is a seasonal cycle of wind, generated by the solar radiation through the following two reasons:

- (i) Astronomical (planetological): Hemispheric differential heating, driving a global meridional circulation. The dominant component of solar heating is equatorially symmetric, which drives the Hadley circulation. Seasonal variation of solar radiation on a planet is generated by the following two reasons:
- Eccentricity of the planetary orbit: globally in-phase, which does not contribute to monsoon.
 - Inclination of the planetary rotation axis: anti-phase between the hemispheres, contributing to monsoon.
- Diurnal cycle of (day-night) hemispheric differential heating of solar radiation may generate thermal tides.
- (ii) Geographical (terrestrial) : Continent-ocean differential heating, driving a flow crossing the coastline. Land (solid) – sea (liquid) heat capacity contrast, generating a diurnal cycle of sea-land breeze circulation. Integration (residue) of diurnal cycle generates the seasonal cycle and thus the monsoon. Wind from ocean to continent (in summer hemisphere) brings moisture, and therefore generates rainy season. The cloud-precipitation water cycle makes latent heat transport, and feeds back to monsoon enhancement.

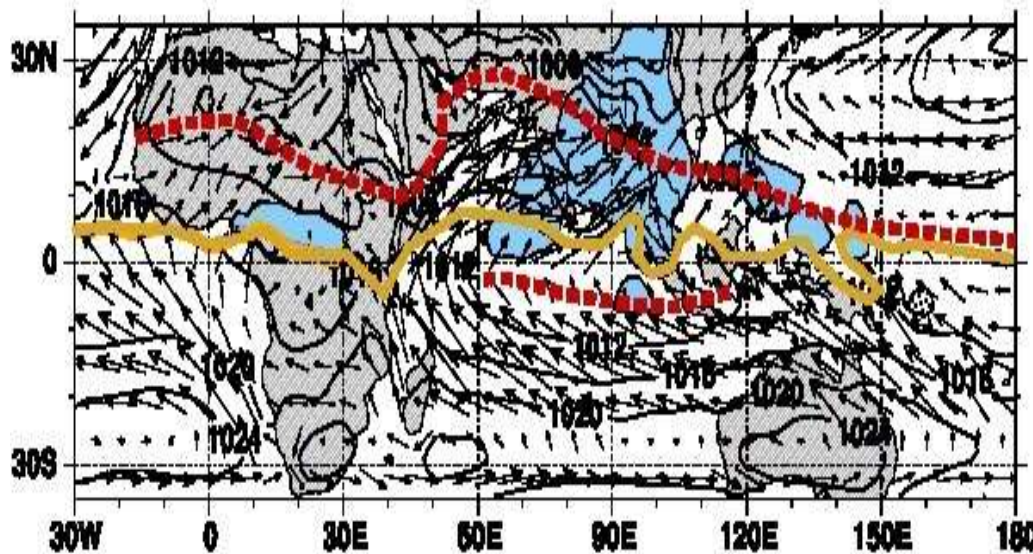
Ocean-Land Contrasts and Monsoon

"Terrestrial" Monsoon Sea-Land Breeze Analogue

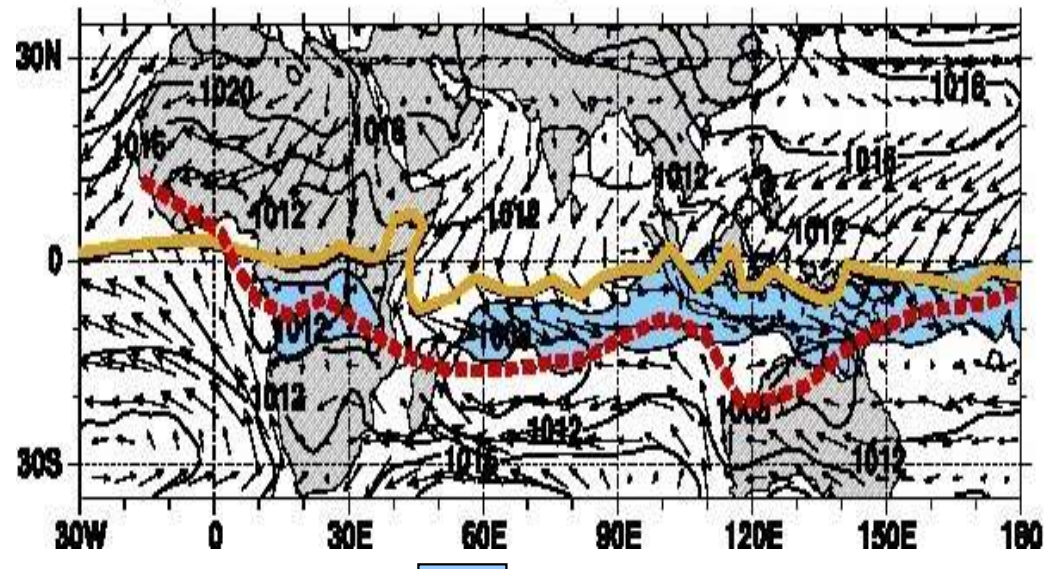


(Wallace & Hobbs, 2006;
originally from UWSSEC)

(a) July 1992 MSL Pressure, 925-mbar wind and OLR



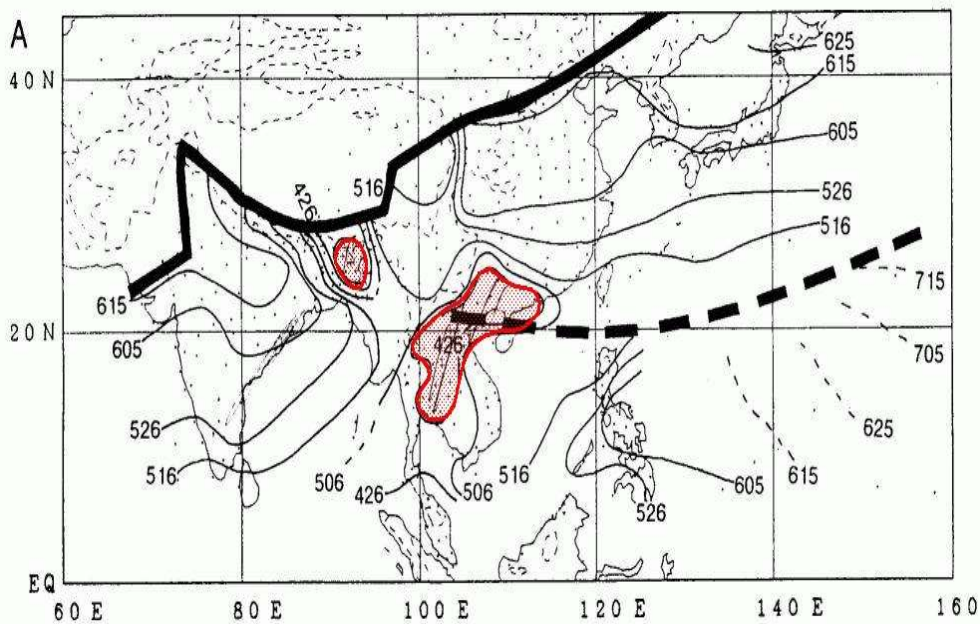
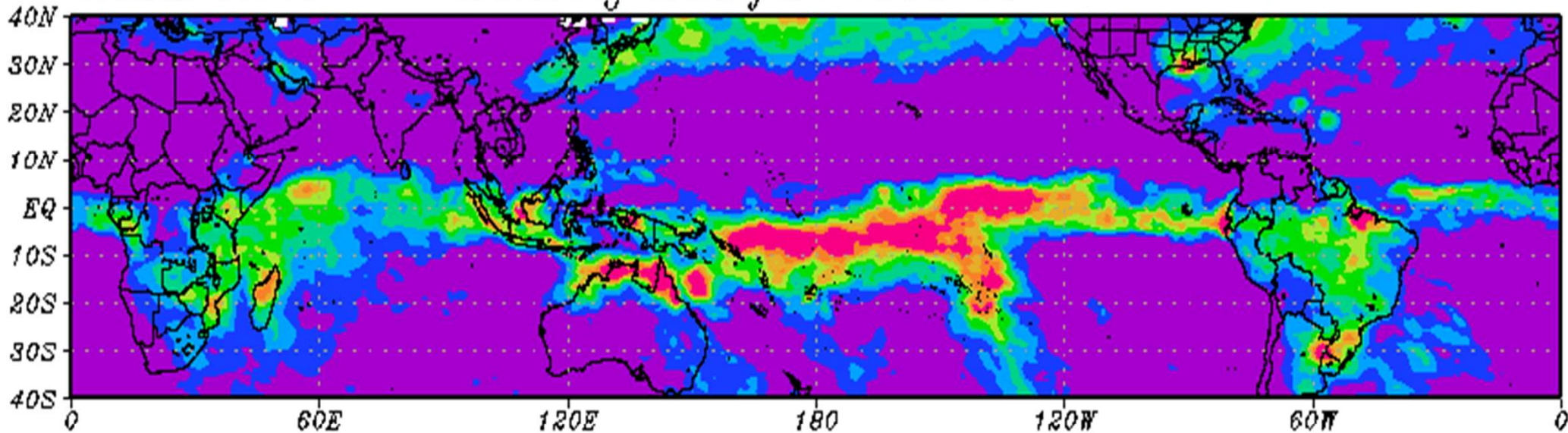
(b) Feb 1992 MSL Pressure, 925-mbar wind and OLR



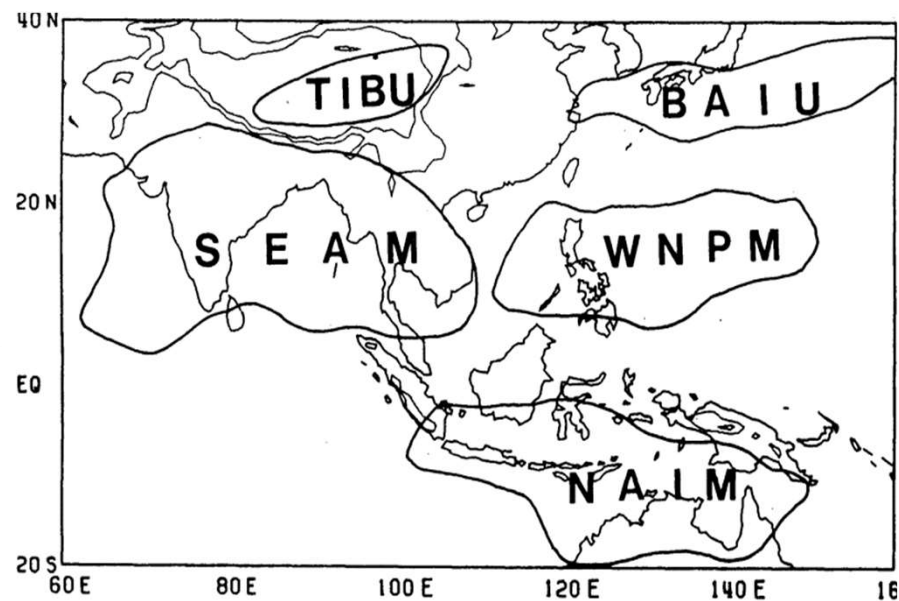
(Webster, 1999)

rain

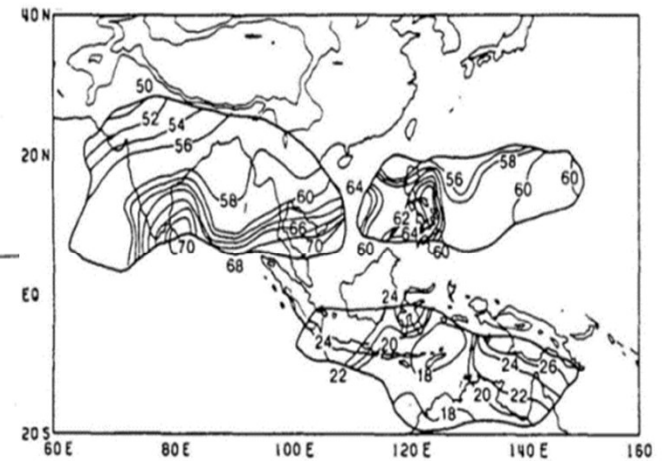
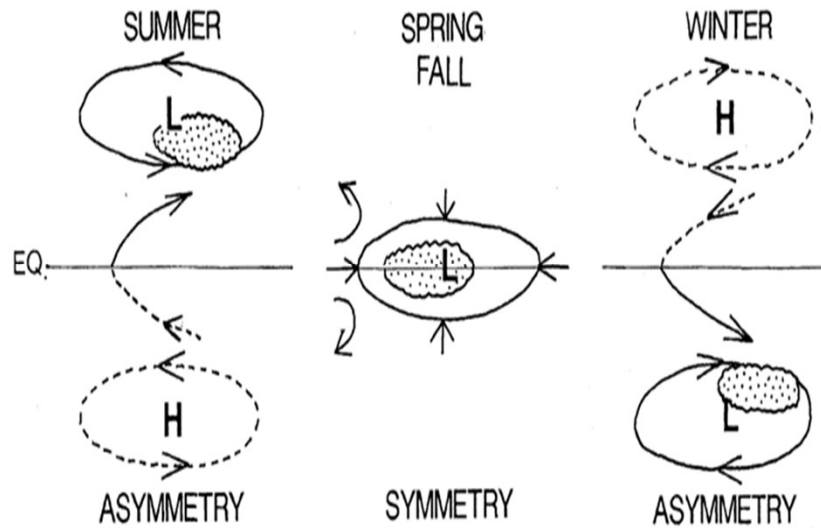
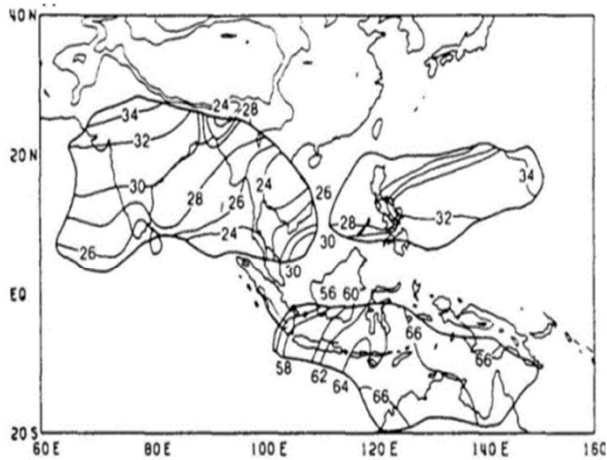
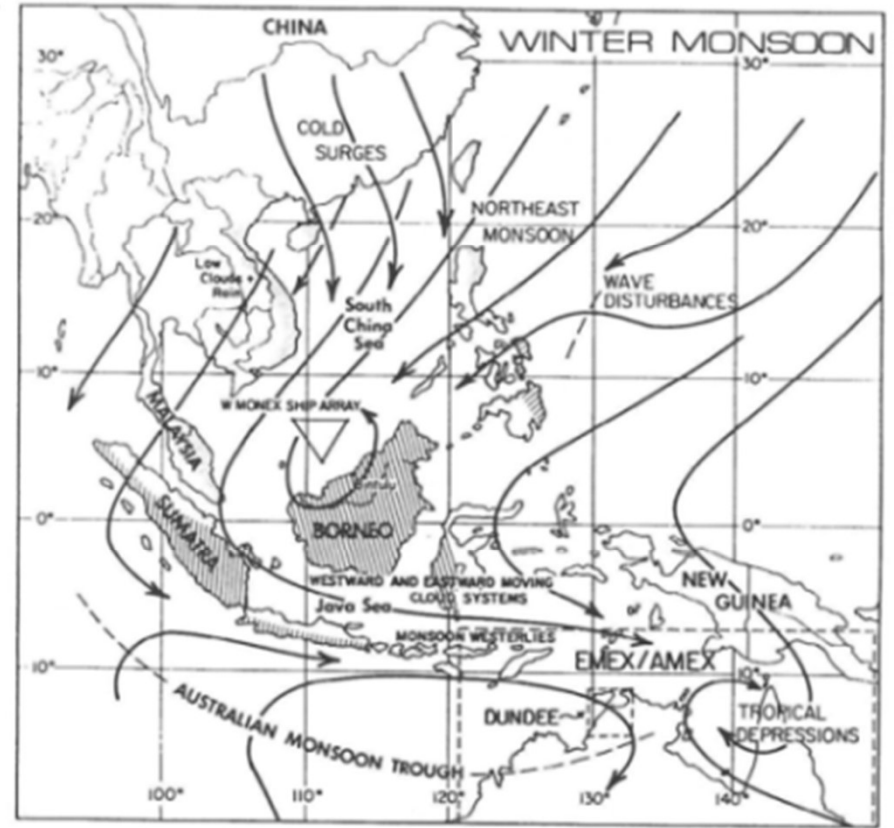
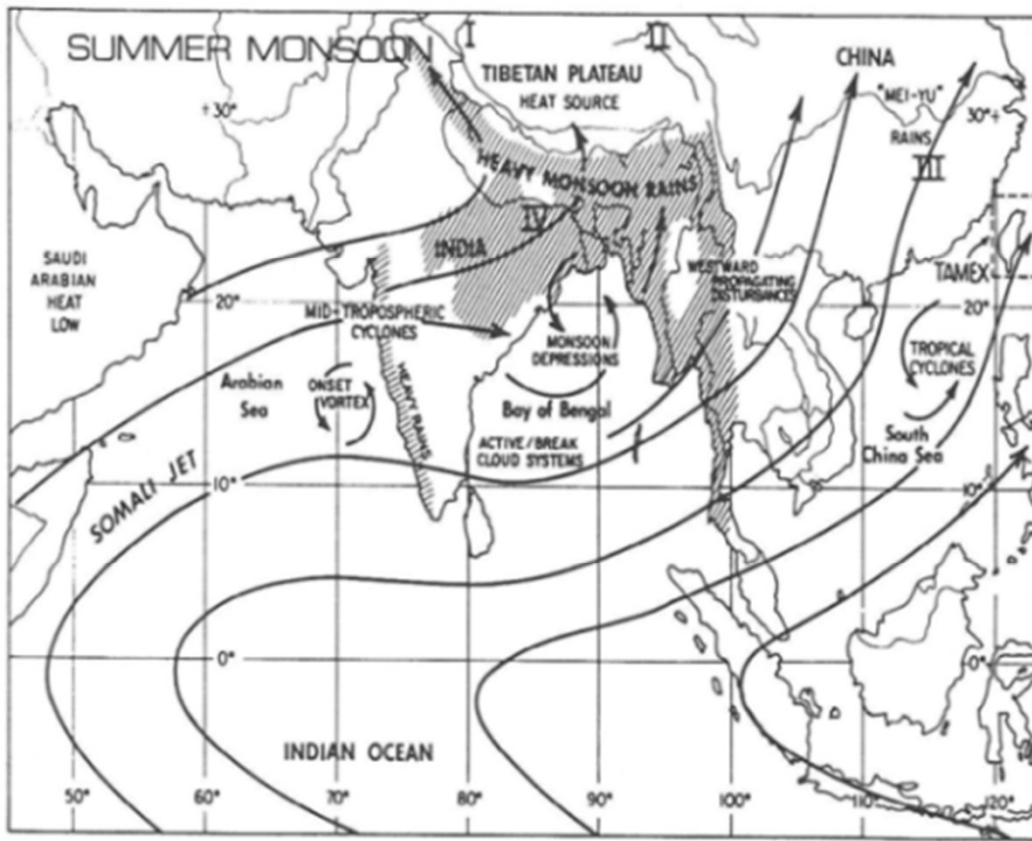
TRMM 3B43 Monthly Rainfall 1998 01



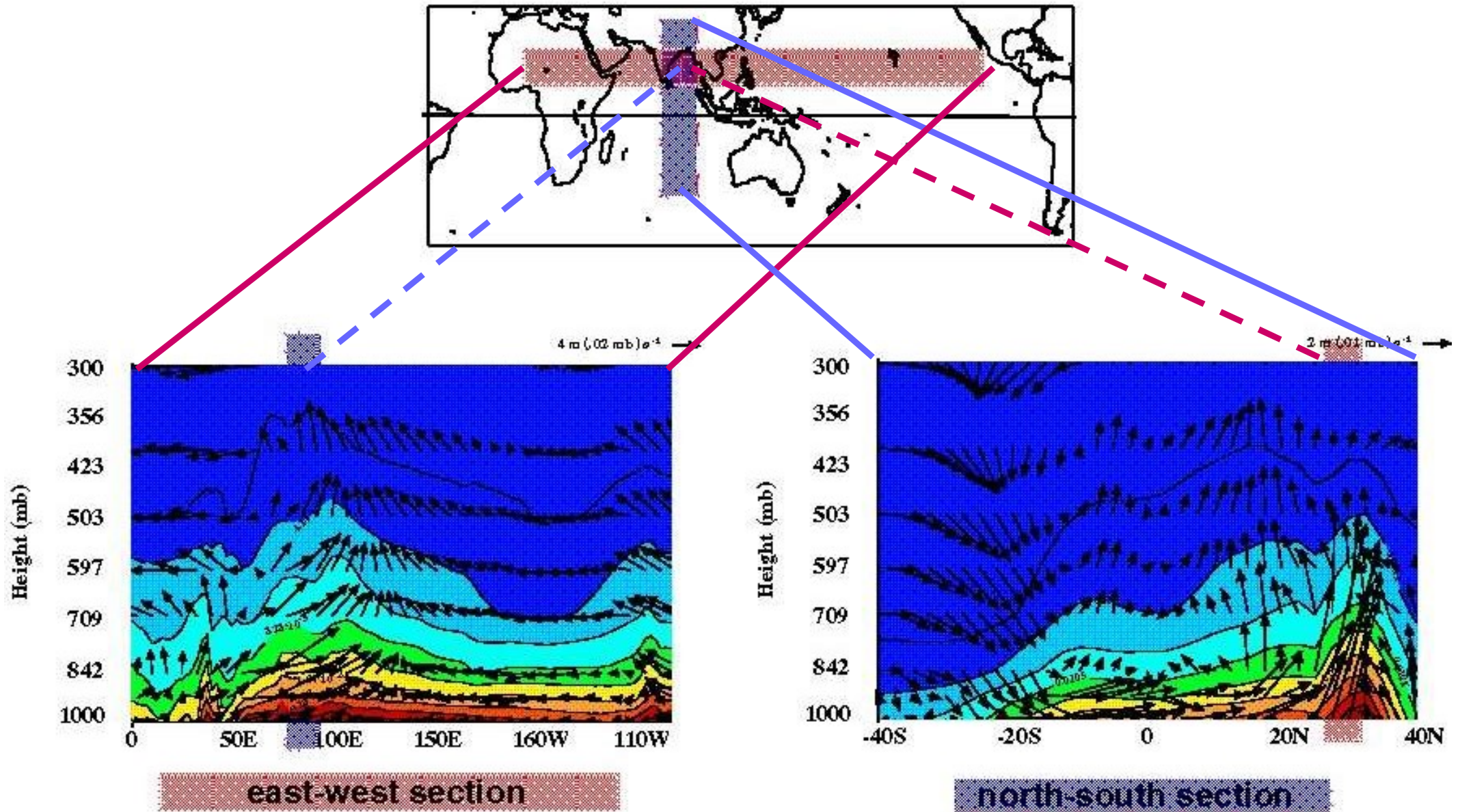
(Matsumoto and Murakami, 1992)



(Murakami and Matsumoto, 1994)

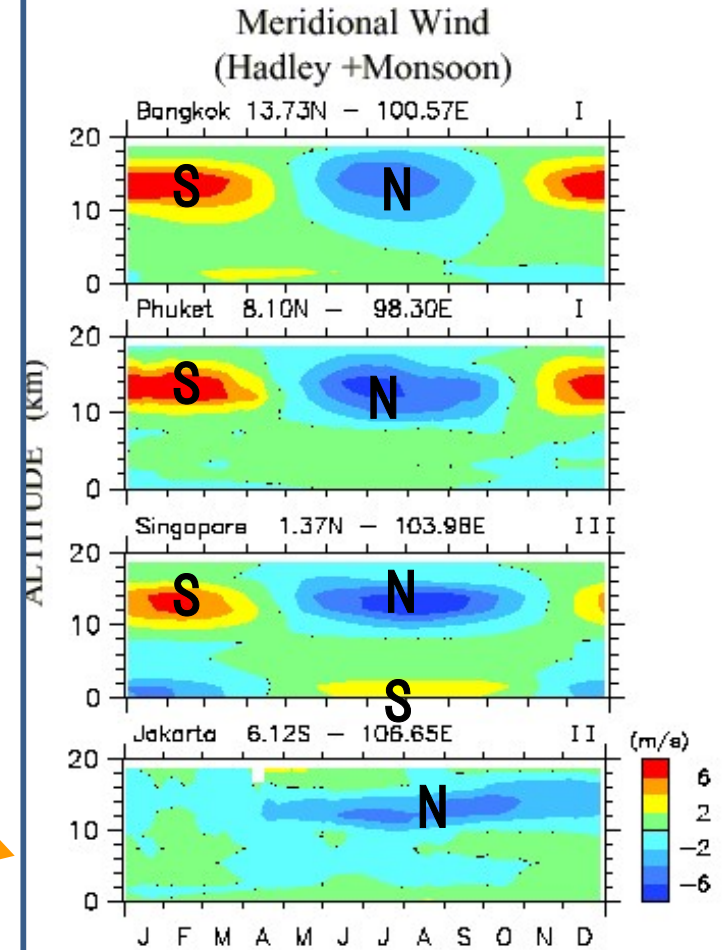
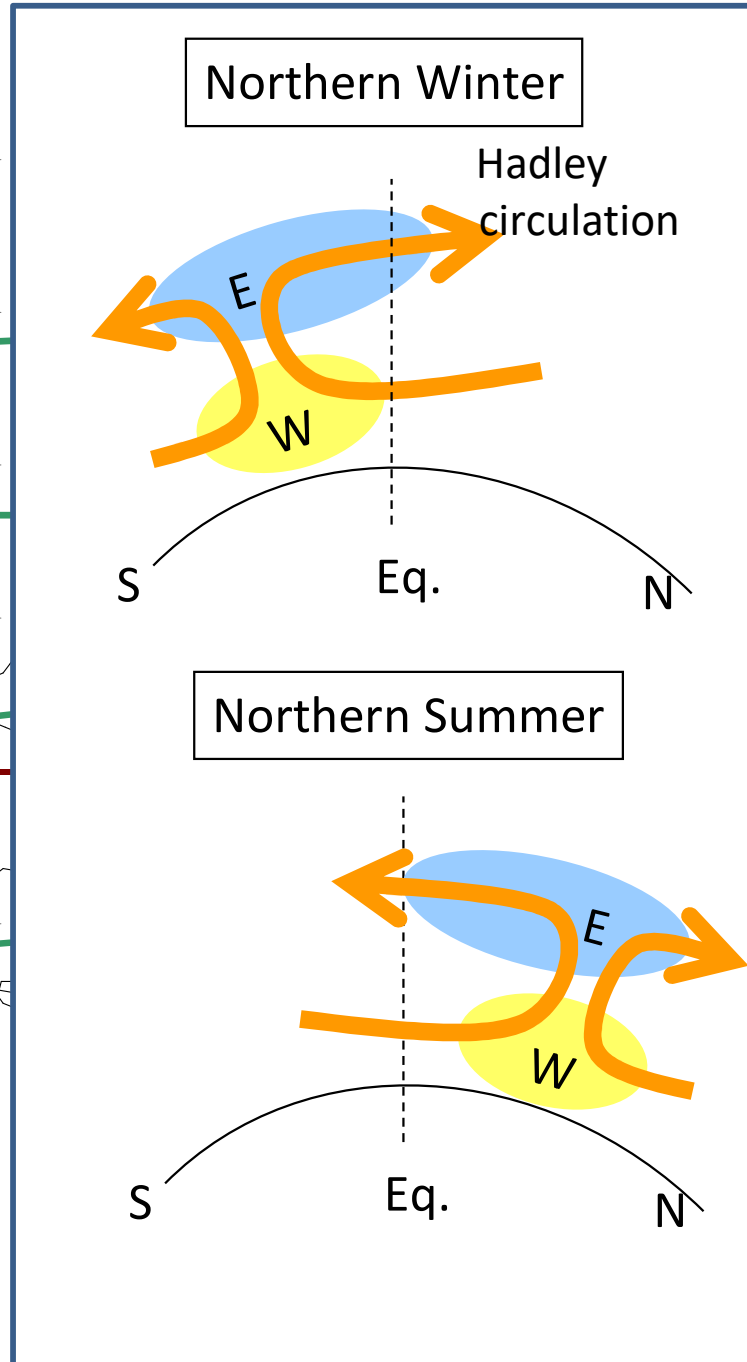
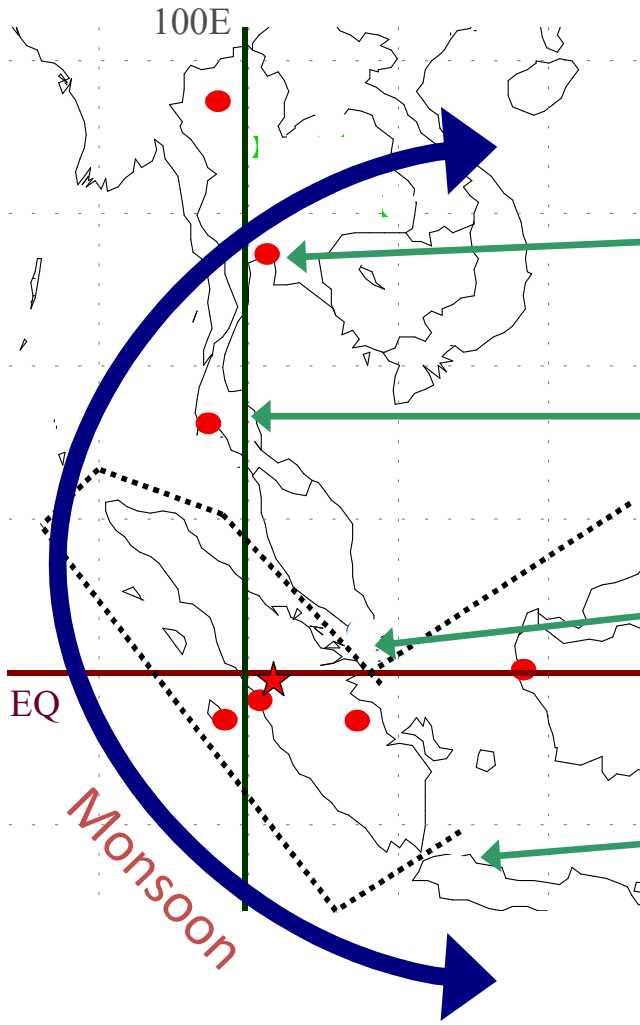


Longitudinal / latitudinal sections in NH Summer

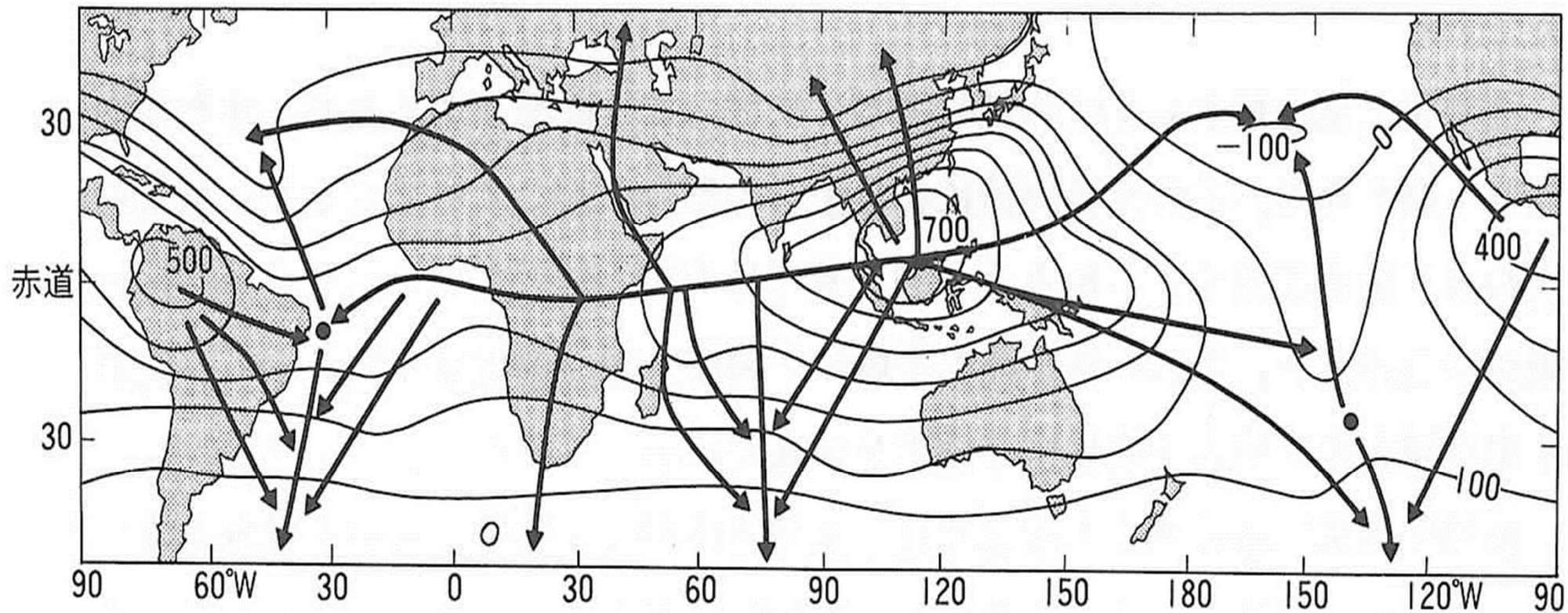


(Webster, 1999)

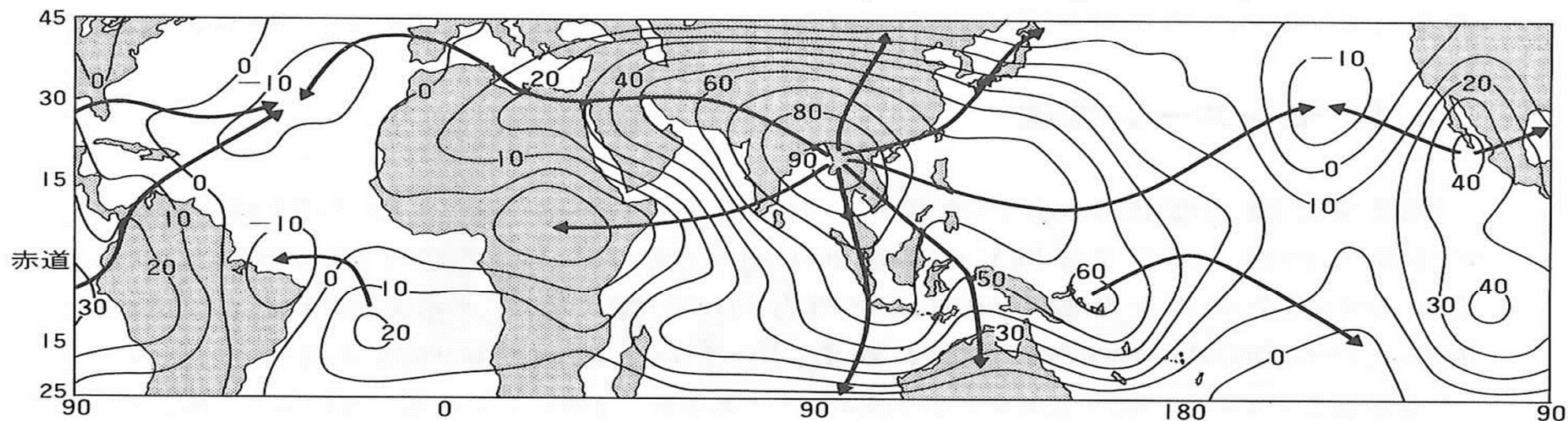
100E meridian (Sumatera-Malay-Thailand) obs.



(Okamoto et al., 2004)



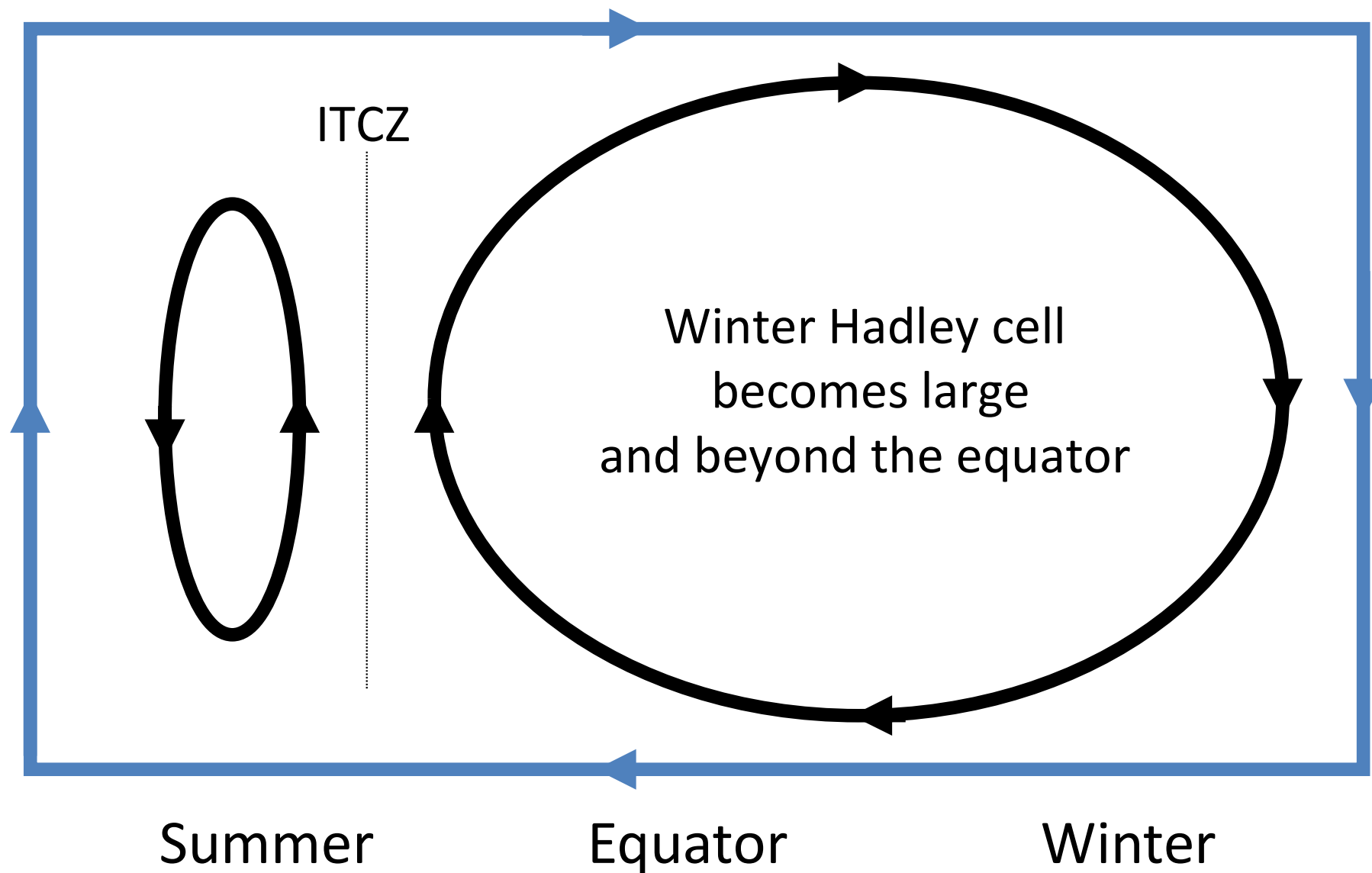
(a)



(b)

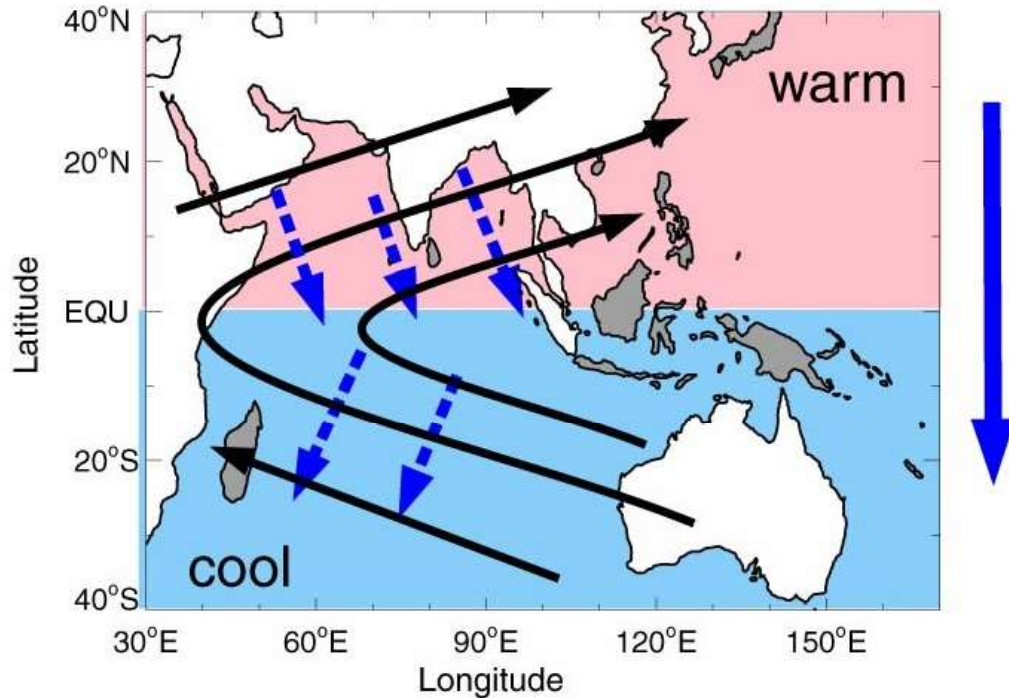
図 5.7 高度 200mb の準定常発散場と速度ポテンシャル χ の分布 (発散風は χ に垂直な方向に吹く). (a)12~2月の図 (Krishnamurti *et al.*, 1973), (b)6~8月の図 (Krishnamurti, 1971).

Superimposing Monsoon and Hadley circulations



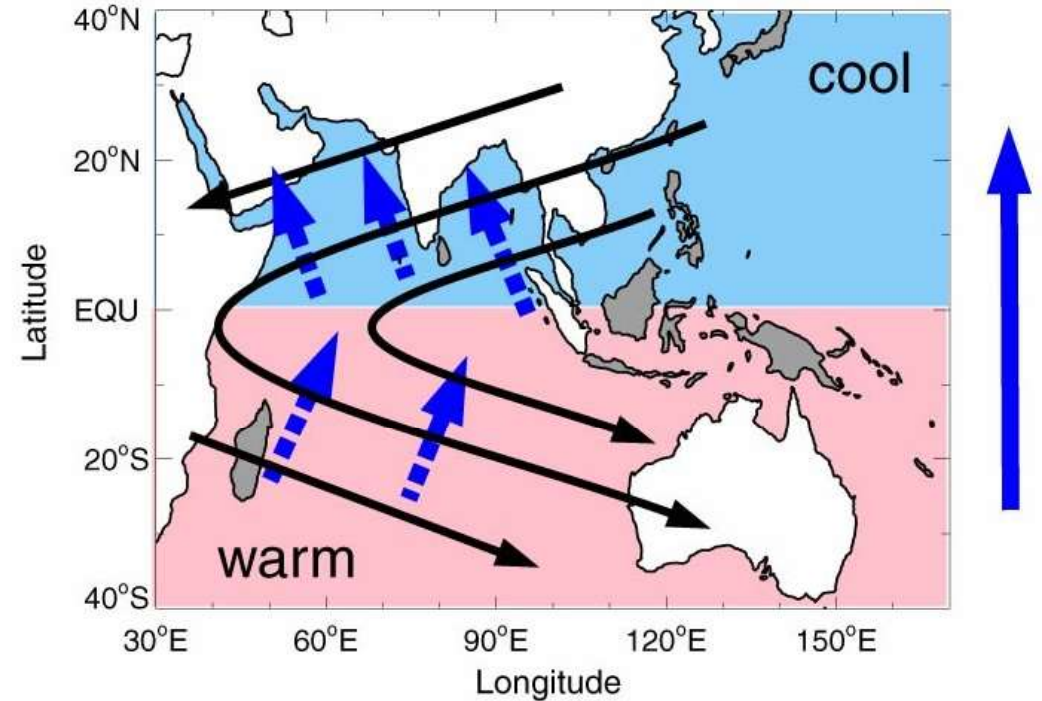
Monsoon-Induced Ocean Heat Transport

Boreal Summer



Southward ocean heat transport of 1.5 PW
(cools NIO while warming SIO)

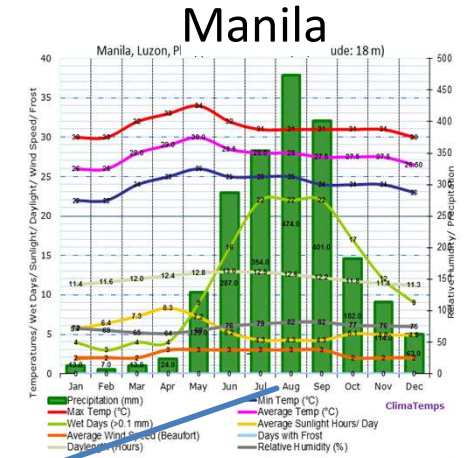
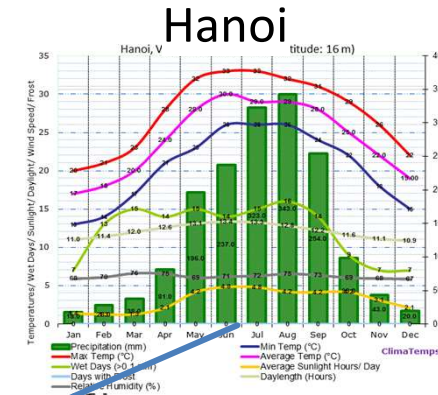
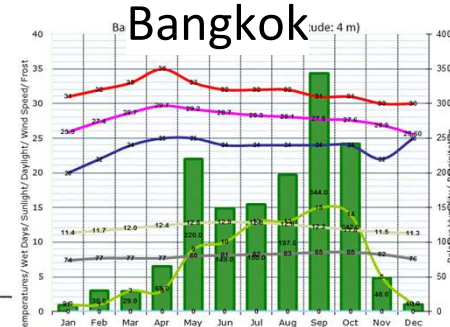
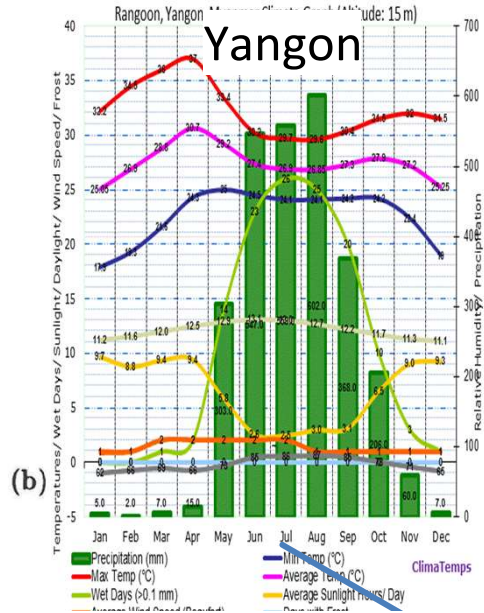
Boreal Winter



Northward ocean heat transport of 1.5 PW
(cools SIO while warming NIO)

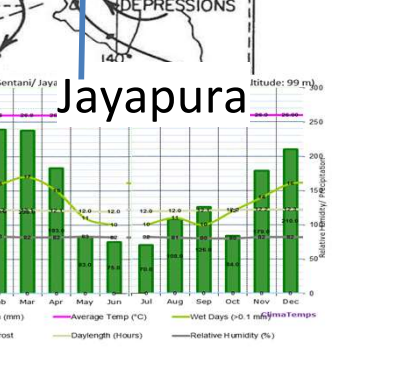
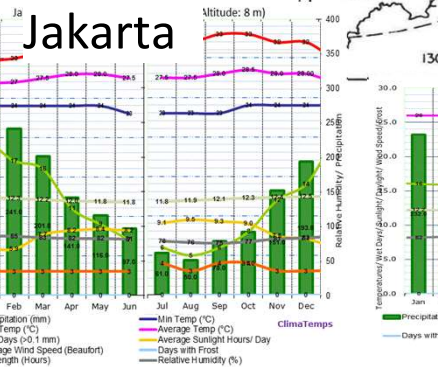
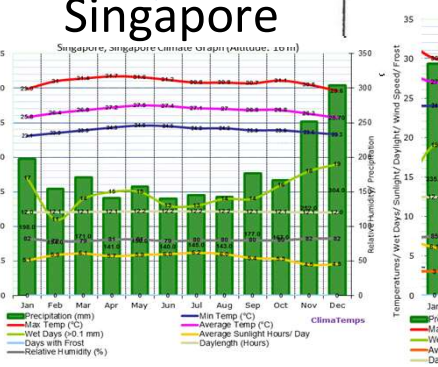
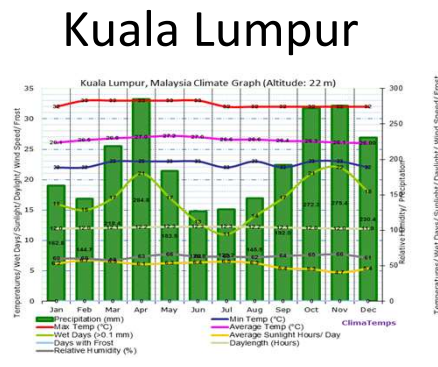
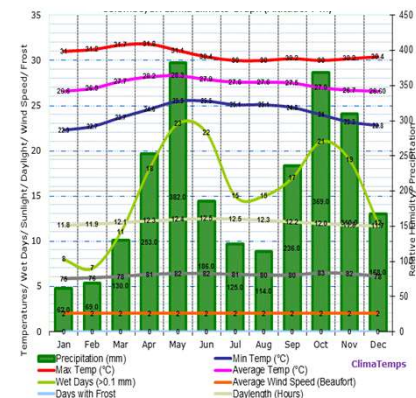
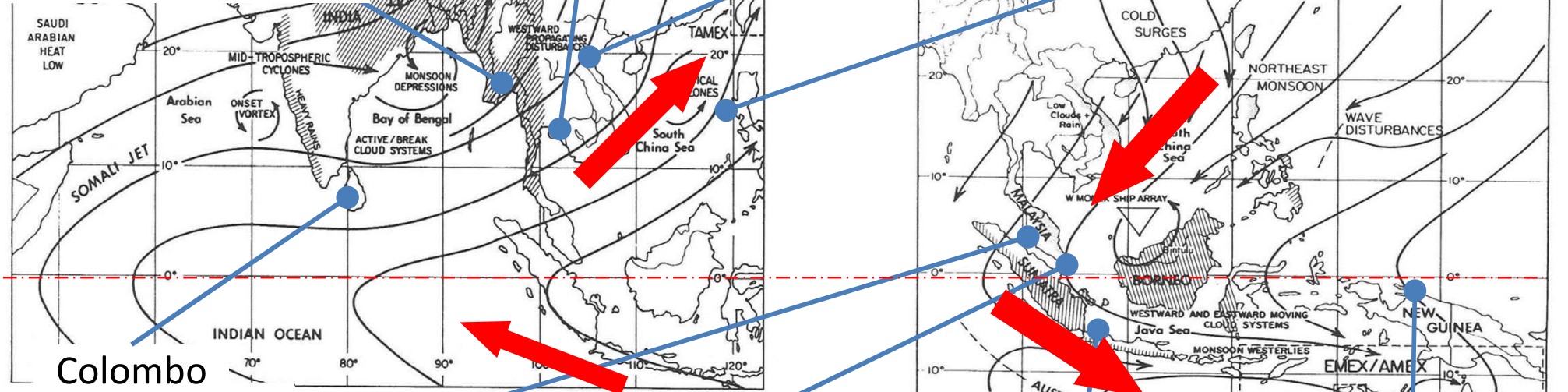
Overall impact of wind-driven Ekman ocean heat transport is to cool the summer hemisphere and warm the winter hemisphere thus reducing the cross-equatorial SST gradient and minimizing seasonal extremes in the monsoon (Webster, 1999)

Monsoons and rainy seasons

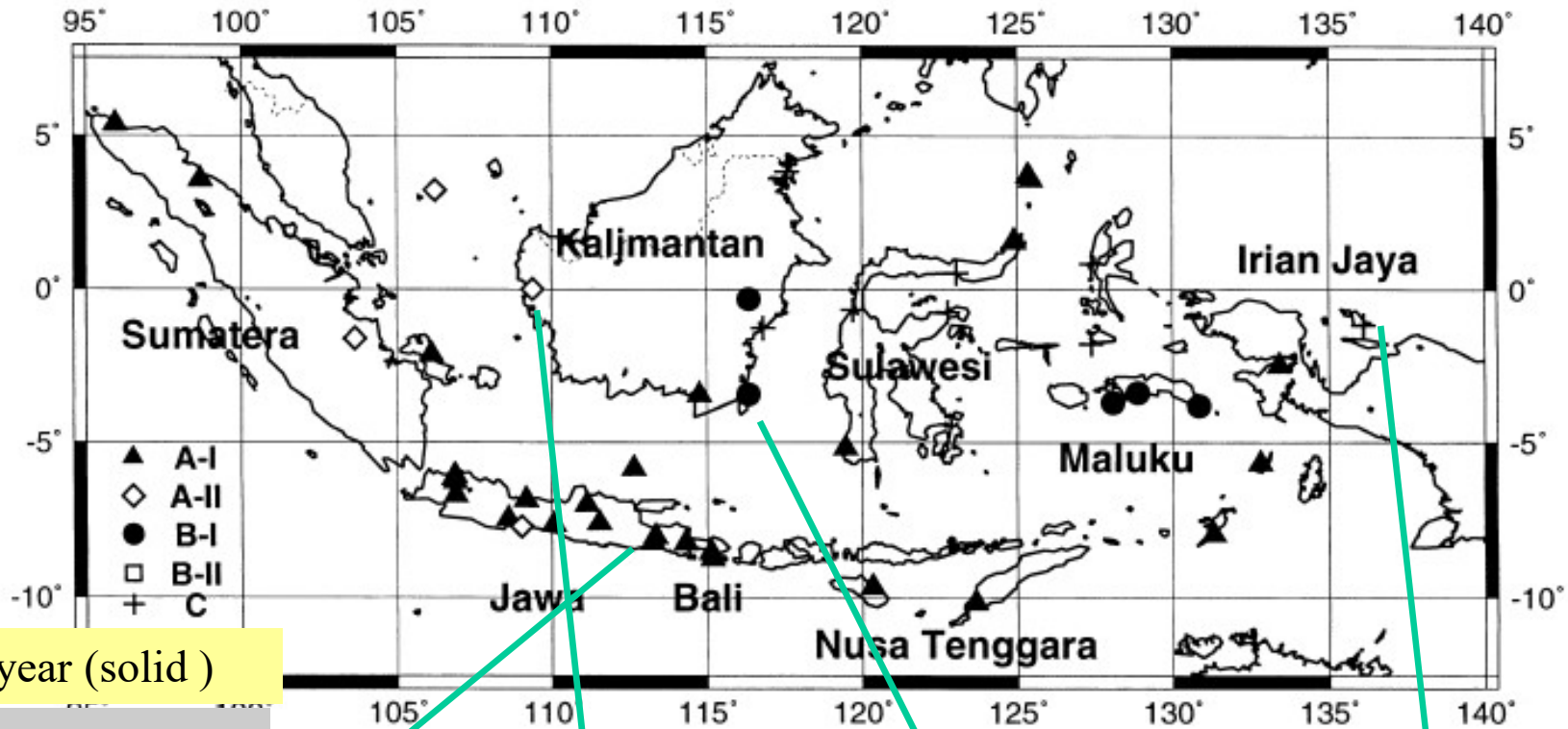


Jun-Jul-Aug monsoon

Dec-Jan-Feb monsoon

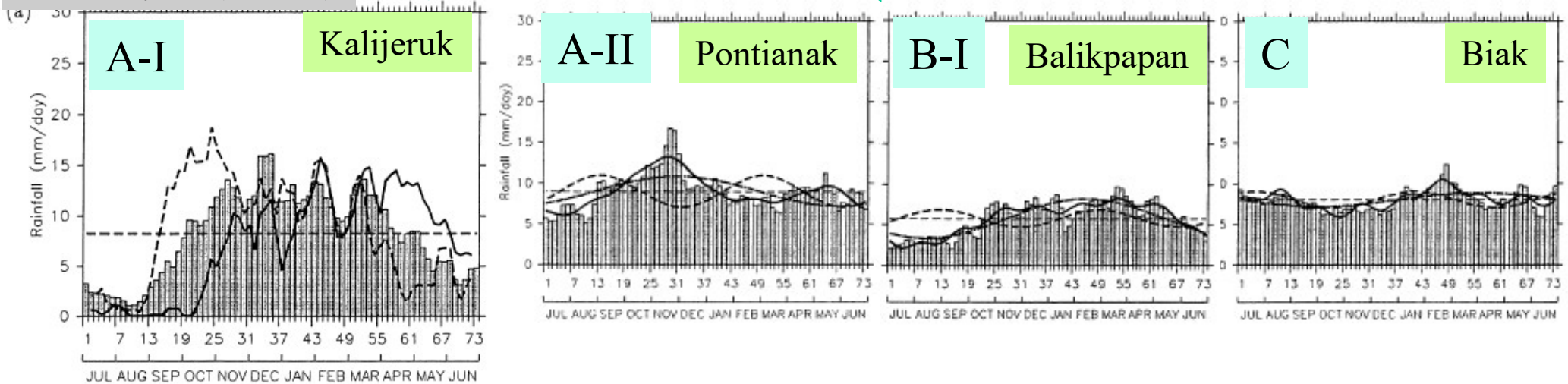


Seasonal and interannual variations



El Nino year (solid)

La Nina year (dashed)

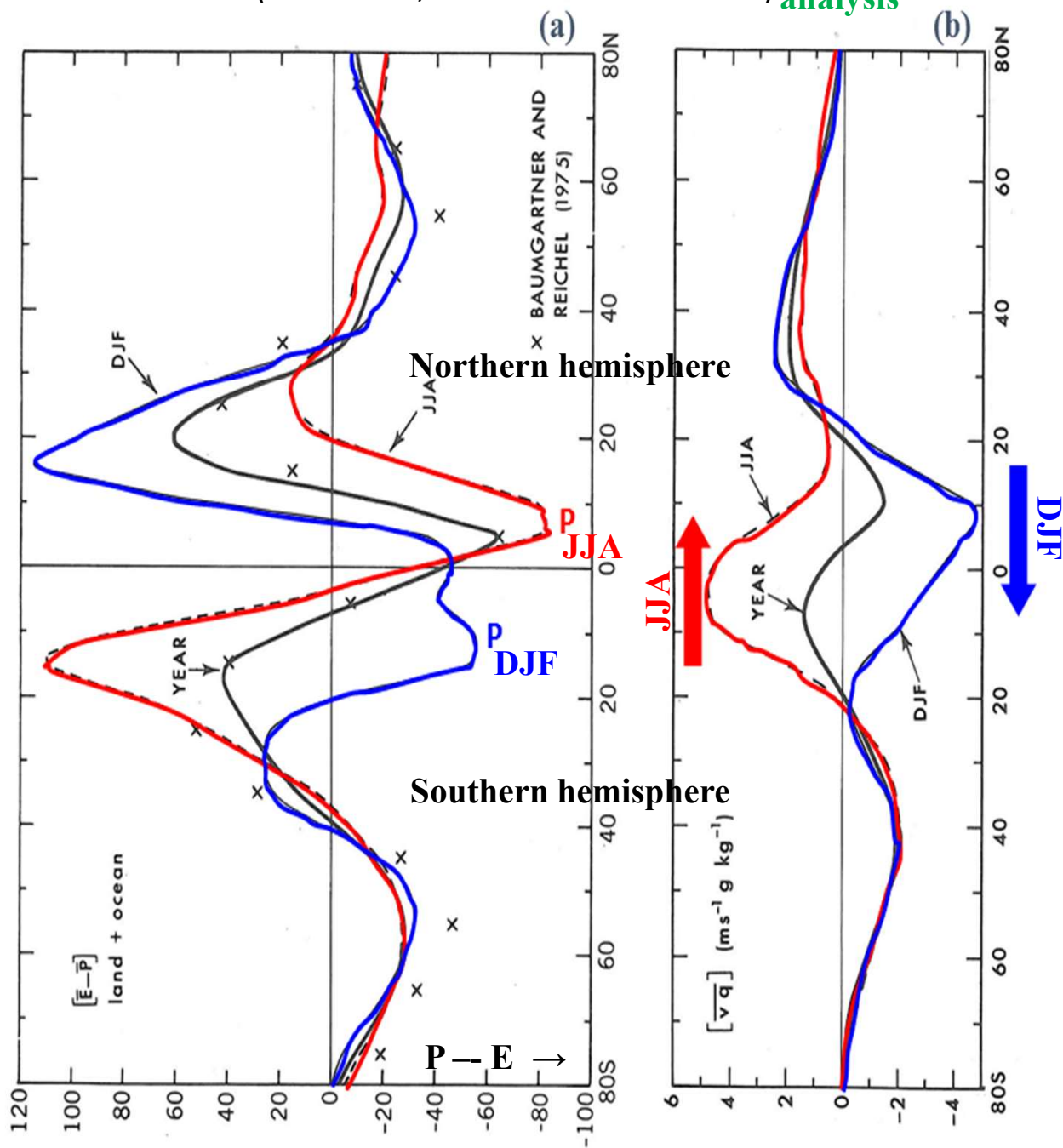


JUL ← → JUN

(Hamada, et al., 2002, *J. Meteorol. Soc. Japan*)

Equatorial annual-cycle dehydrator

(1963-1973, based on 1000 stations) **Subjective analysis**

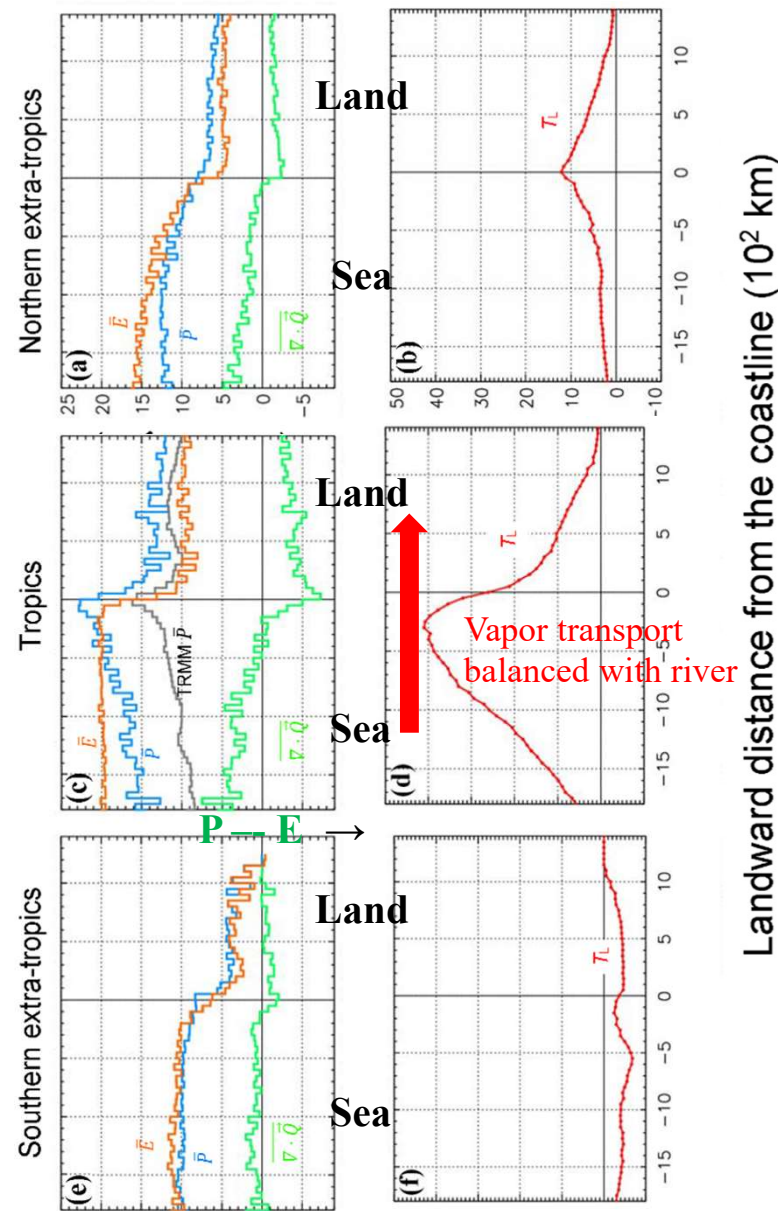


(Peixoto and Oort, 1983, *Variations in the Global Water Budget*)

Coastal diurnal-cycle dehydrator

(1981-2010, from **JRA55** **Operational objective reanalysis**)

$\bar{P}, \bar{E}, \bar{v} \cdot \bar{Q}$ Water Vapor Transport
(10^2 mm yr^{-1}) ($10^3 \text{ km}^3 \text{ yr}^{-1}$)



(Ogino et al., 2017, *GRL*)

Contribution of **tropical** rainfall to global climate balance

Global mean rainfall $\approx 1,100 \text{ mm/year} \times 0.7 \text{ (ocean)} + 700 \text{ mm/year} \times 0.3 \text{ (land)}$
 $\approx 1,000 \text{ mm/year}$

(Multiplying water density $1,000 \text{ kg/m}^3$)

→ Water mass precipitating per unit area $\approx 1 \times 10^3 \text{ kg/m}^2/\text{year}$

(Multiplying latent heat $\approx 600 \text{ cal/g} \approx 2.5 \times 10^6 \text{ J/kg}$)

(Dividing the result by 1 year $\approx 3.15 \times 10^7 \text{ s}$)

→ Heating atmosphere per unit mass and unit time $\approx 80 \text{ W/m}^2$
(about 1/3 of the green house (IR heating) effect)

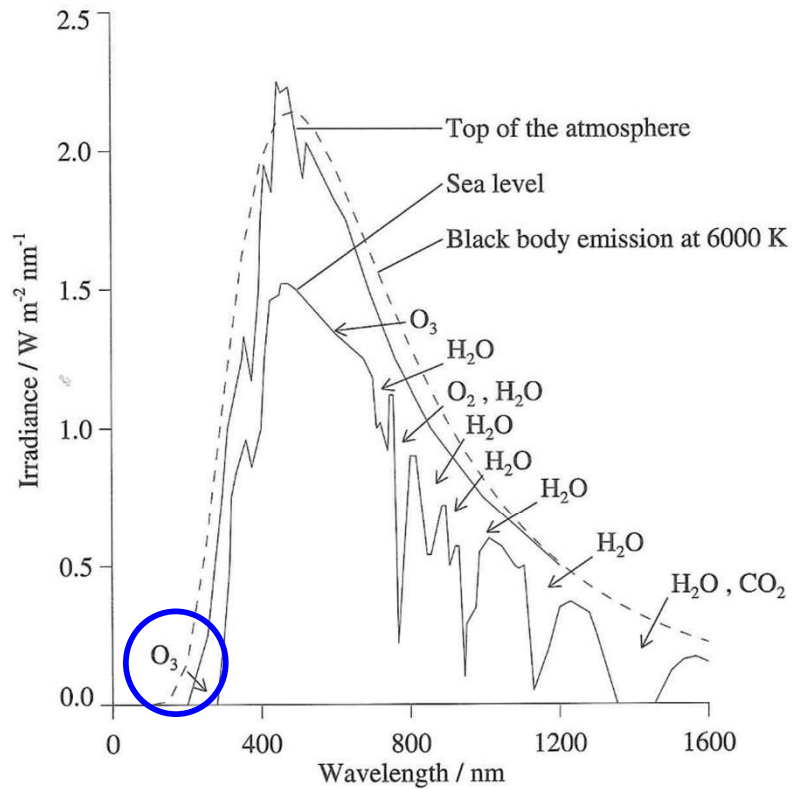
Tropical rainfall: largest over the globe (2,000 mm/year near the equator)

Southeast Asian rainfall: largest in the tropics (2,500 mm/year in average)

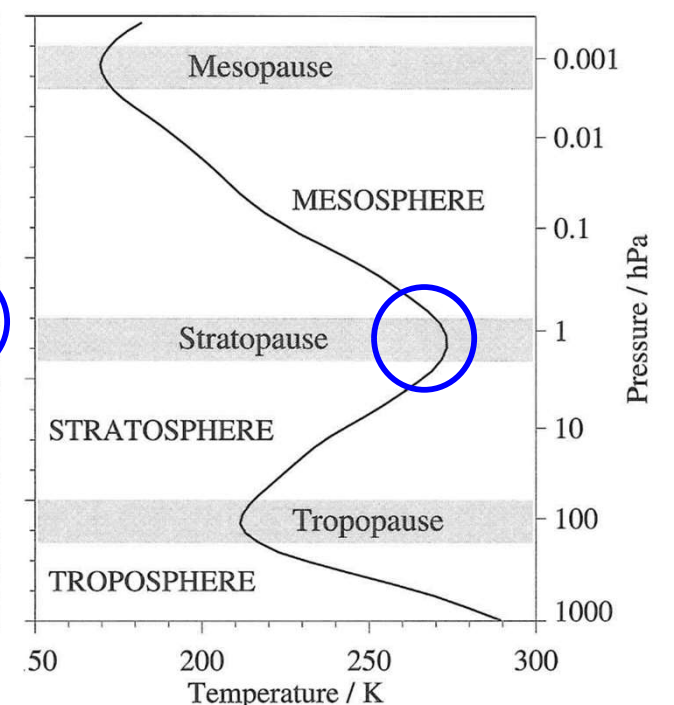
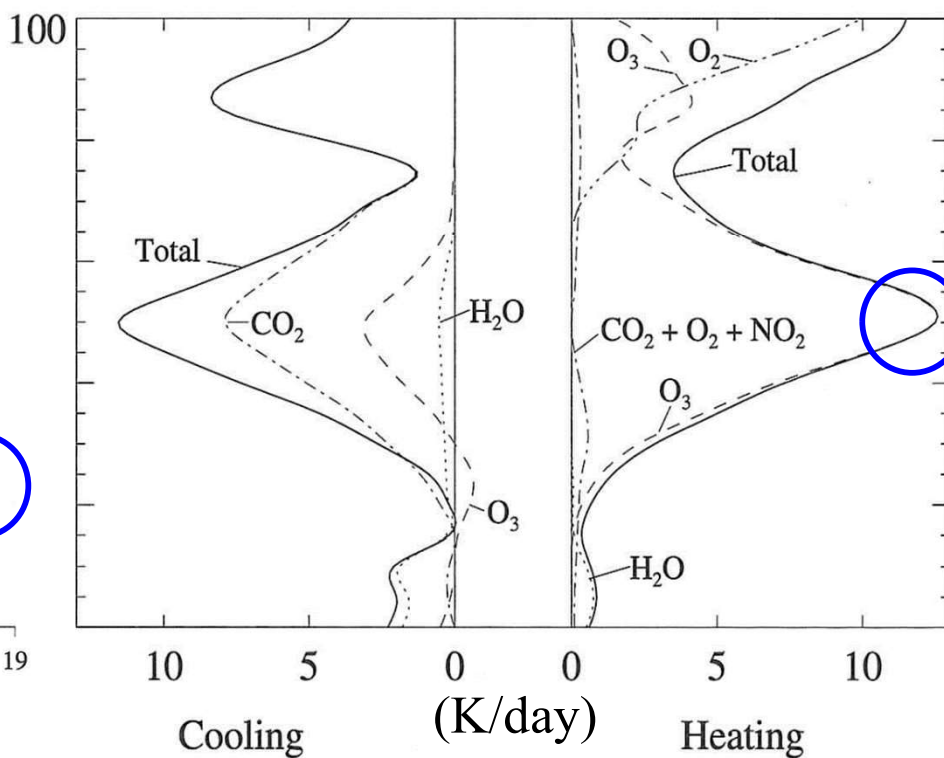
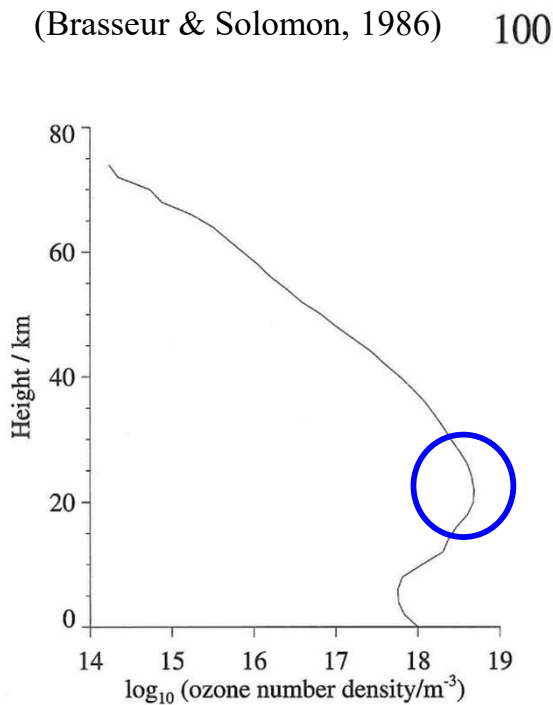
→ about 2.5 times of the global mean

→ Maritime-continent rainfall variations affect the global climate greatly.

Ozone UV heating and Middle Atmosphere (Mesosphere & Stratosphere)



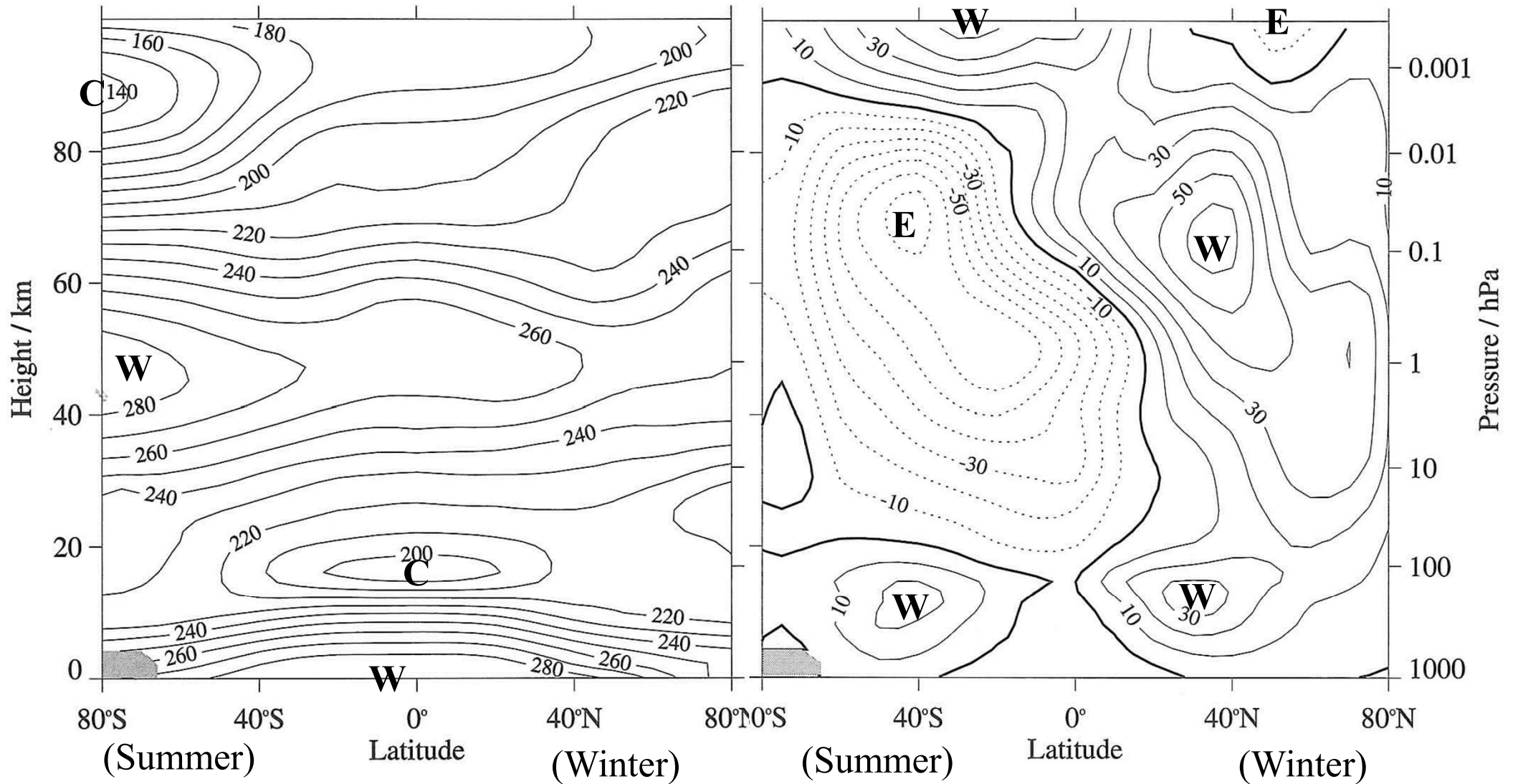
(Brasseur & Solomon, 1986)

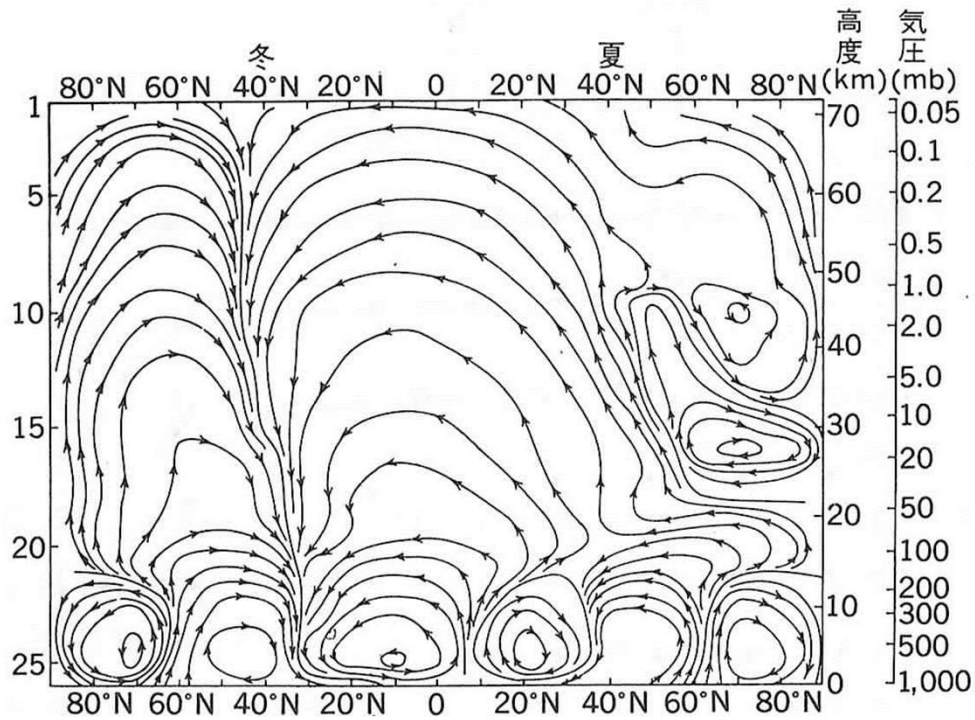


(CIRA1986)

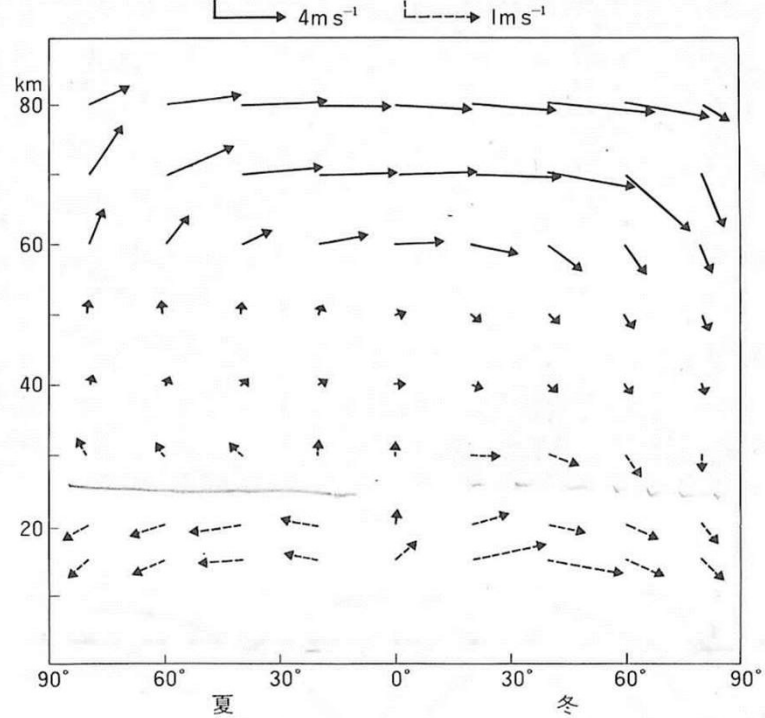
Meridional distribution of Temperature and Zonal Wind

(January, CIRA1986)

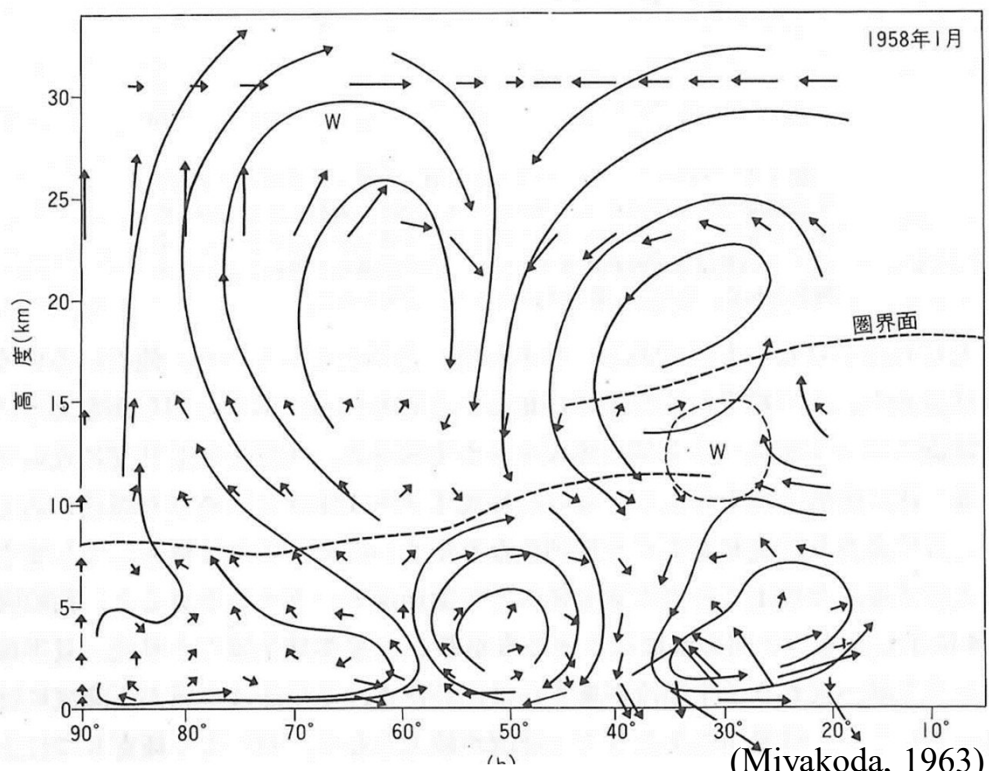




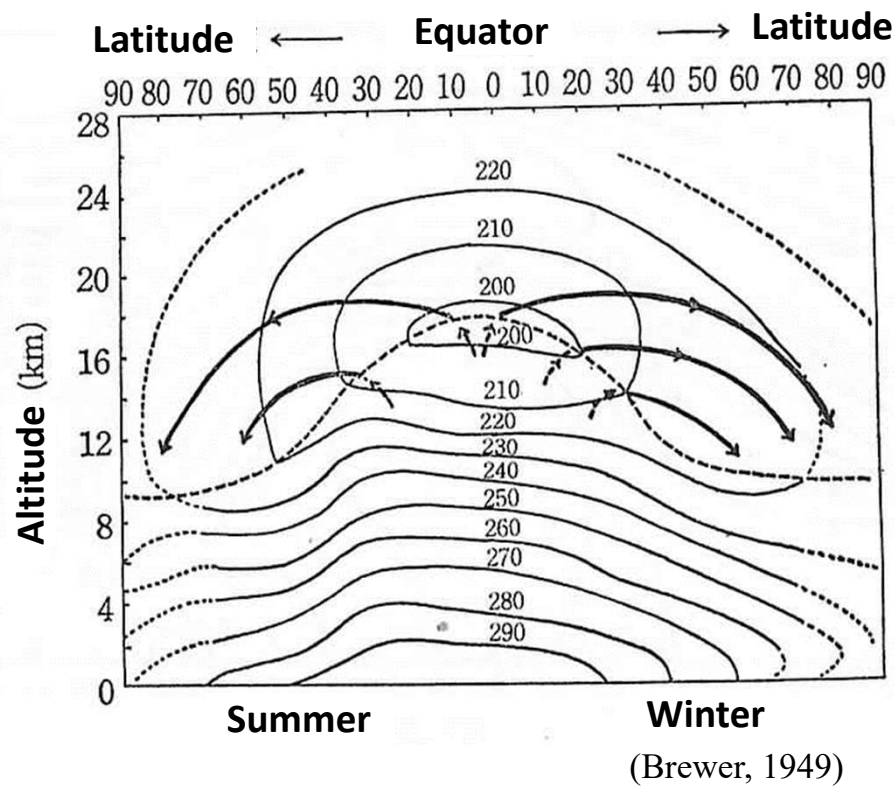
(Cunnold et al., 1975)



(Murgatroyd & Singleton, 1961)

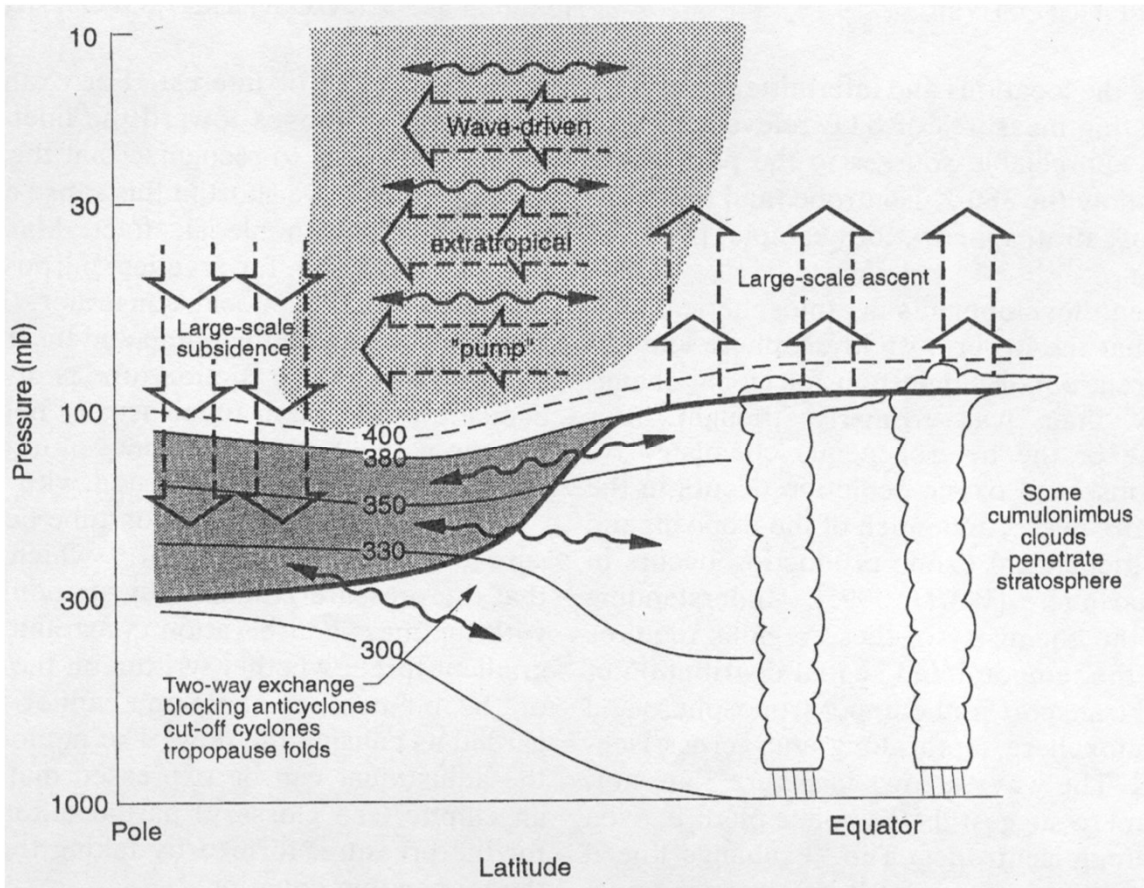


(Miyakoda, 1963)



(Brewer, 1949)

Stratospheric response to tropospheric convection



(Holton, 1995)

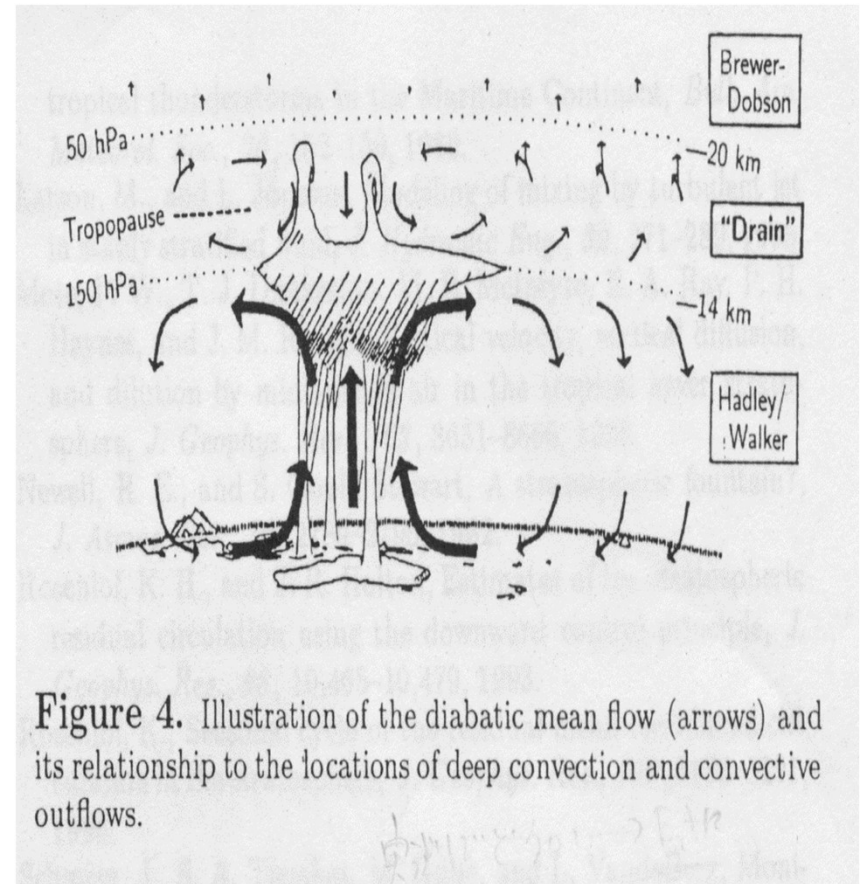


Figure 4. Illustration of the diabatic mean flow (arrows) and its relationship to the locations of deep convection and convective outflows.

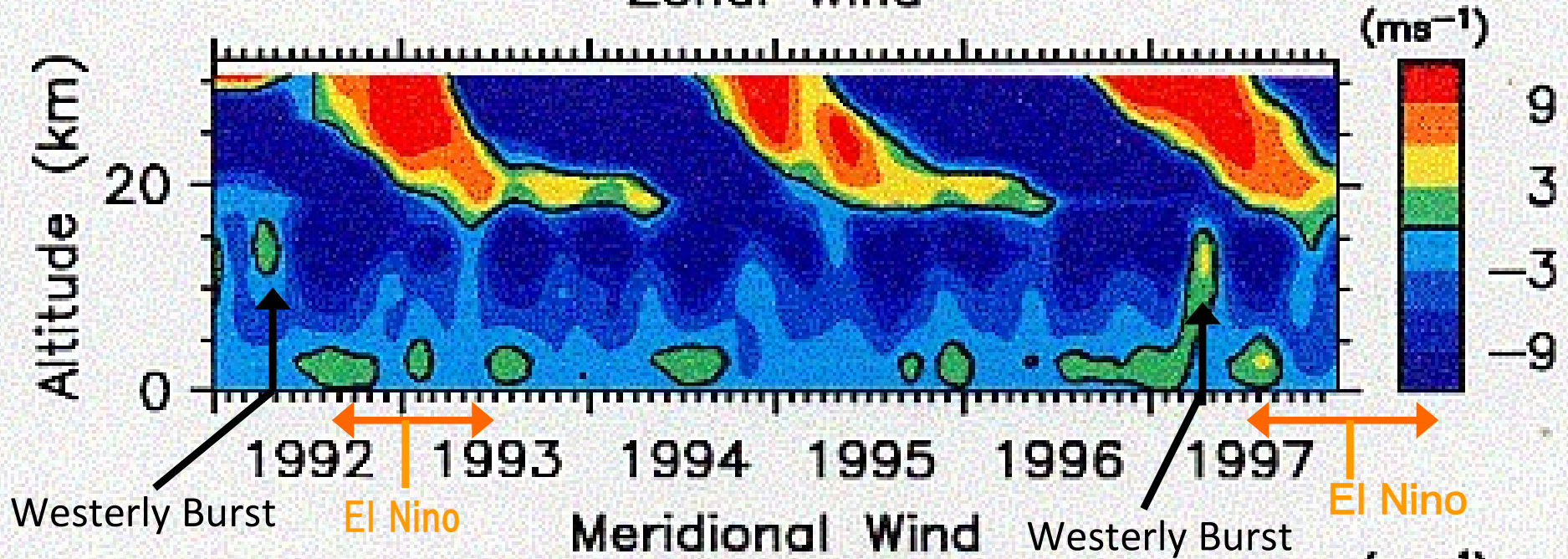
(Sherwood, 2000)

cf. Ogino's ozone observations at Hanoi

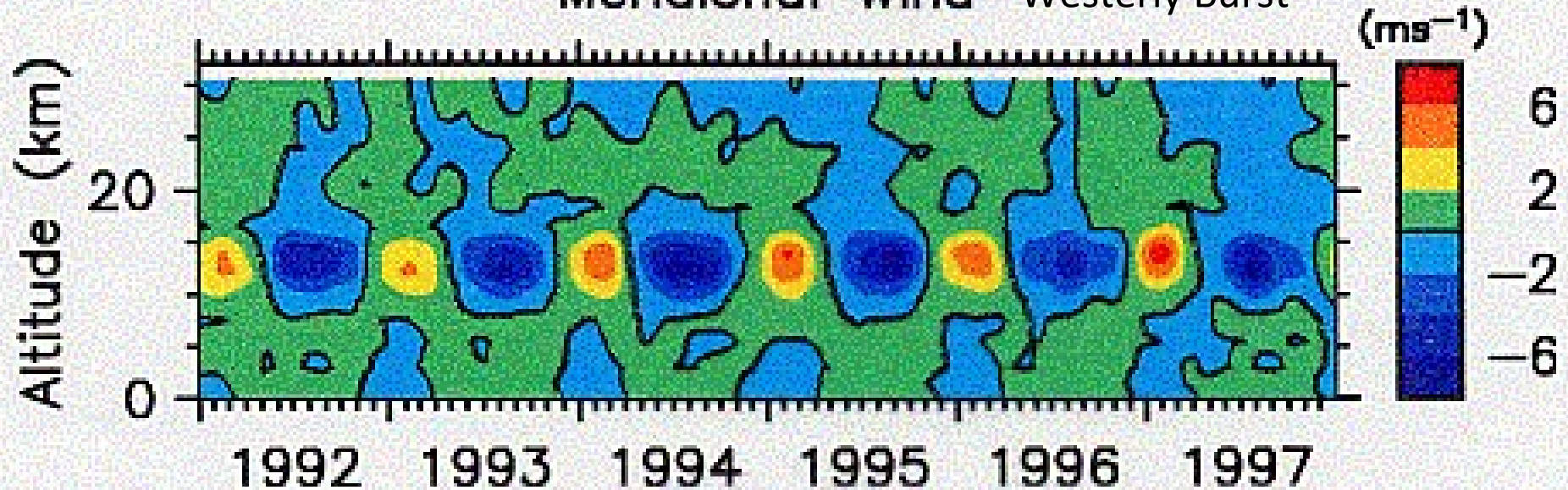
Interannual Variations of Wind over Indonesia

(Okamoto, Yamanaka et al., 2003b)

Zonal Wind



Meridional Wind



Variabilities of climate and weather

- Annual mean – Vertical/Meridional dependencies

- Shorter/smaller-scale variabilities
 - Waves and convections/clouds

- Longer/larger-scale variabilities
 - Interannual variations