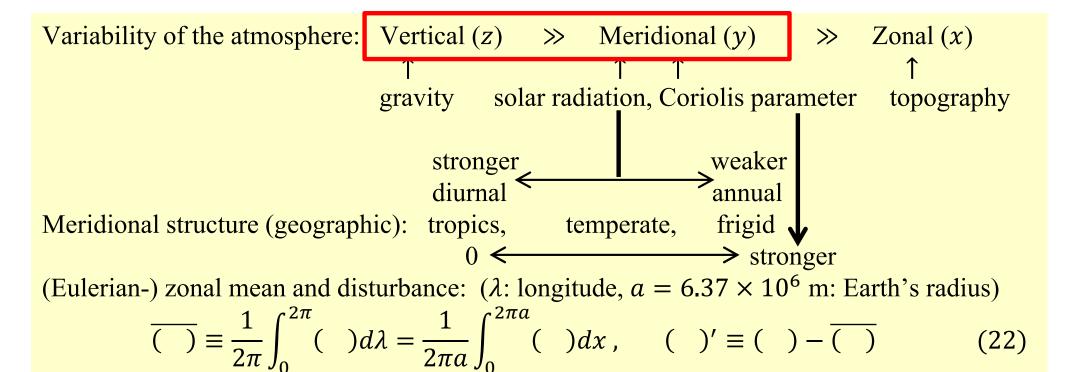
#### 4. Mean zonal-meridional circulations: 4.1-a. Differential heating



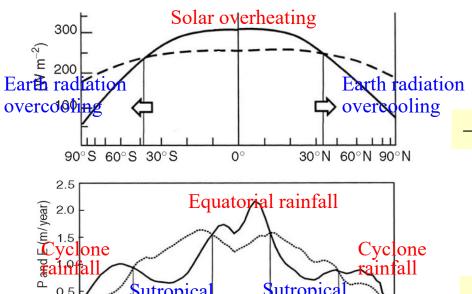
Cyclone

rainfall

60°N 90°N

Sutropica

30°N



**Sutropical** 

90°S 60°S

Figure 7 Annual-mean net incoming solar radiation (solid line) and outgoing terrestrial radiation (dashed line) as a function of latitude, expressed in units of W m<sup>-2</sup>. Distance on the latitude scale is proportional to area on the Earth's surface. (Diagram provided by Socorro Medina.)

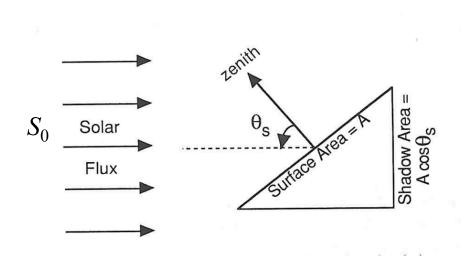
→ Poleward atmospheric and oceanic heat transport

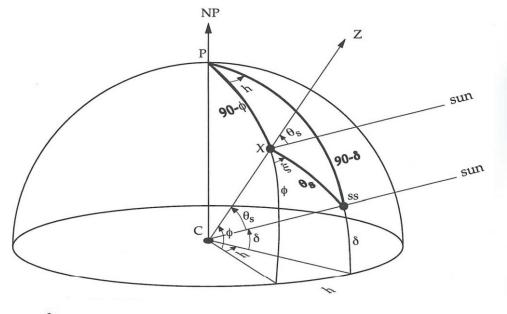
Figure 10 Annual-mean precipitation (solid) and evaporation (dashed) as a function of latitude, expressed in units of meters per year. Distance on the latitude scale is proportional to area on the Earth's surface. Based on NCEP/NCAR Reanalyses for the period 1958–1997. (Diagram provided by Socorro Medina.)

→ Equatorward water vepor transport

(Wallace, 2005)

#### 2-dimensional climate: Meridional-seasonal





$$Q = S_0 \left(\frac{\overline{d}}{d}\right)^2 \cos \theta_s$$

 $\cos \theta_{s} = \cos(90 - \phi)\cos(90 - \delta) + \sin(90 - \phi)\sin(90 - \delta)\cos h$ 

(d: Solar distance, given by Kepler's eqation)

 $\cos \theta_s = \sin \phi \sin \delta + \cos \phi \cos \delta \cos h$ 

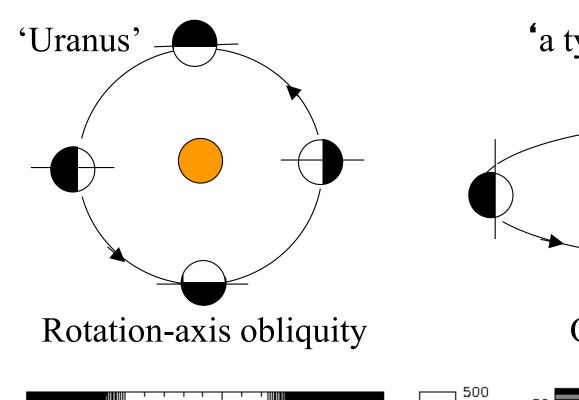
$$\frac{\sin(180 - \xi)}{\sin(90 - \delta)} = \frac{\sin h}{\sin \theta_s}$$

$$\cos h_0 = -\tan \phi \tan \delta$$

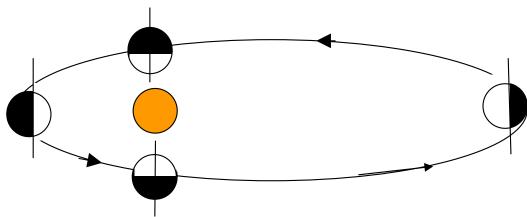
$$\sin \xi = \frac{\cos \delta \sin h}{\sin \theta_s}$$

$$\overline{Q}^{\text{day}} = \frac{S_0}{\pi} \left(\frac{\overline{d}}{d}\right)^2 \left[h_0 \sin \phi \sin \delta + \cos \phi \cos \delta \sin h_0\right]$$

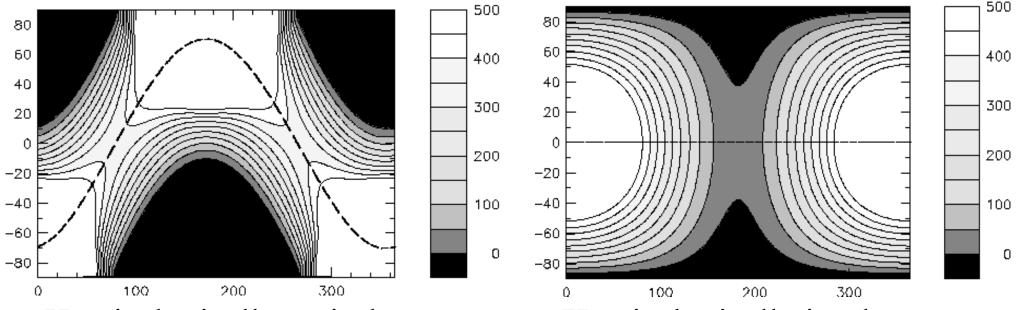
#### Two extreme cases of seasonal cycle forcing



'a type of extra-solar planet'



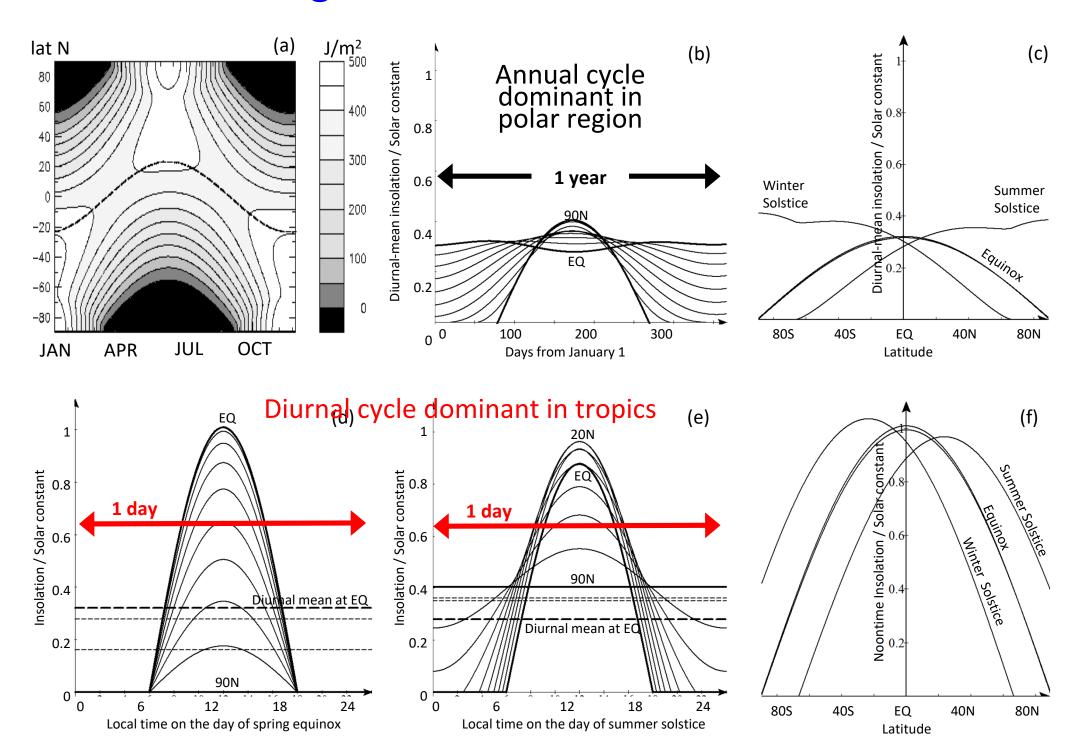
Orbital eccentricity



Hemispherically anti-phase

Hemispherically in-phase

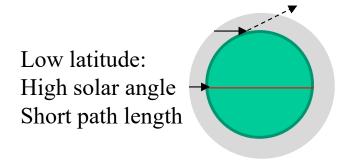
#### Solar heating on earth with revolution and rotation



## Why tropics is tropics?

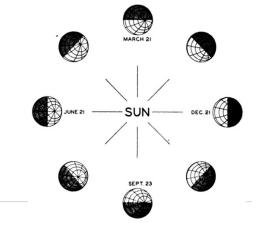
- High latitudes: Low solar angle
  - Large reflection (strong albedo)
  - Large refraction (long optical depth)
- Low temperature → Cryosphere/Clouds —

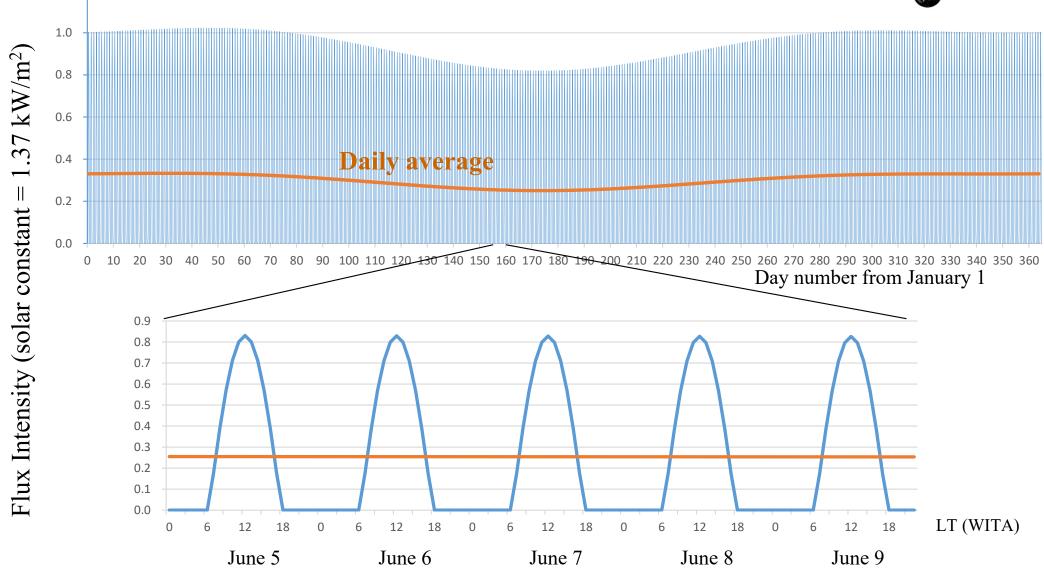
- Rapid rotation &Slow circular revolution
  - Zonal homogeneity
  - Weak hemispheric anti-symmetry



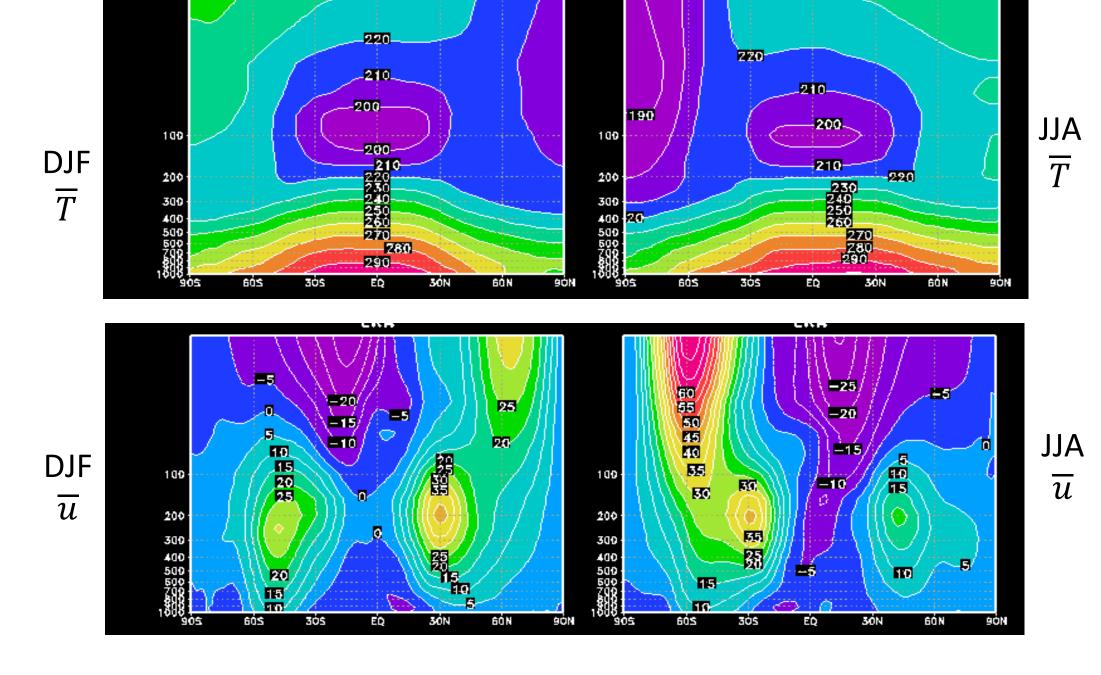
# Solar radiation heating at the atmosphere top over Denpasar (8°39'S)

1.2

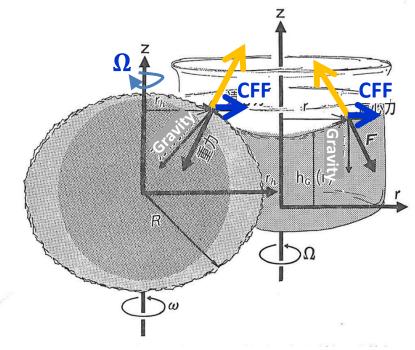


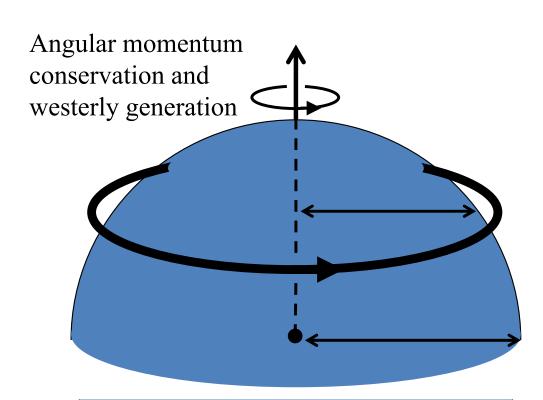


#### 4.1-c Trade wind (Equatorial easterly & mid-latitude westerly)



#### Rotating fluid and centrifugal/Coriolis forces





#### "Thermal wind" equilibrium

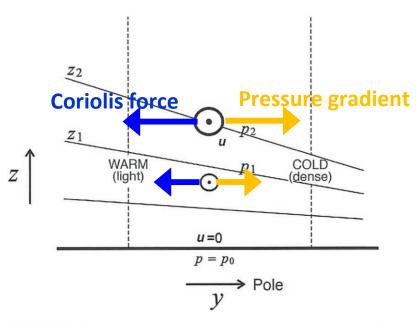
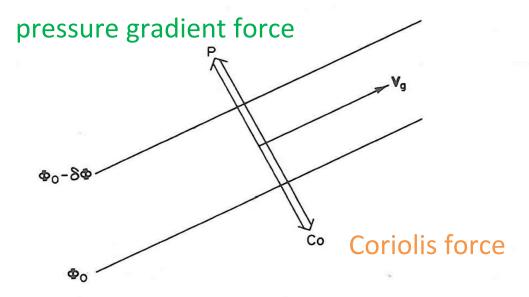


FIGURE 5.14. Warm columns of air expand, cold columns contract, leading to a tilt of pressure surfaces, a tilt which typically increases with height in the troposphere. In Section 7.3, we will see that the corresponding winds are out of the paper, as marked by  $\odot$  in the figure. (Marshall & Plumb, 2009)

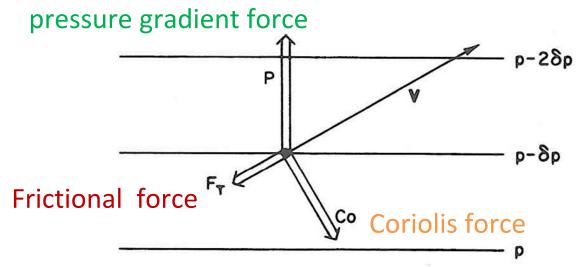


"Keirin" circuit



**Fig. 3.2** Balance of forces for geostrophic equilibrium. The pressure gradient force is designated by P and the Coriolis force by Co.

(Holton, 1972)



**Fig. 5.3** Balance of forces in the well-mixed planetary boundary layer: P designat gradient force, Co the Coriolis force, and  $F_T$  the turbulent drag.



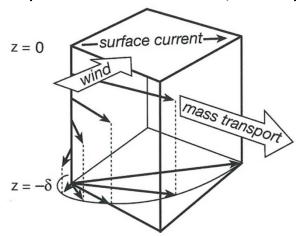
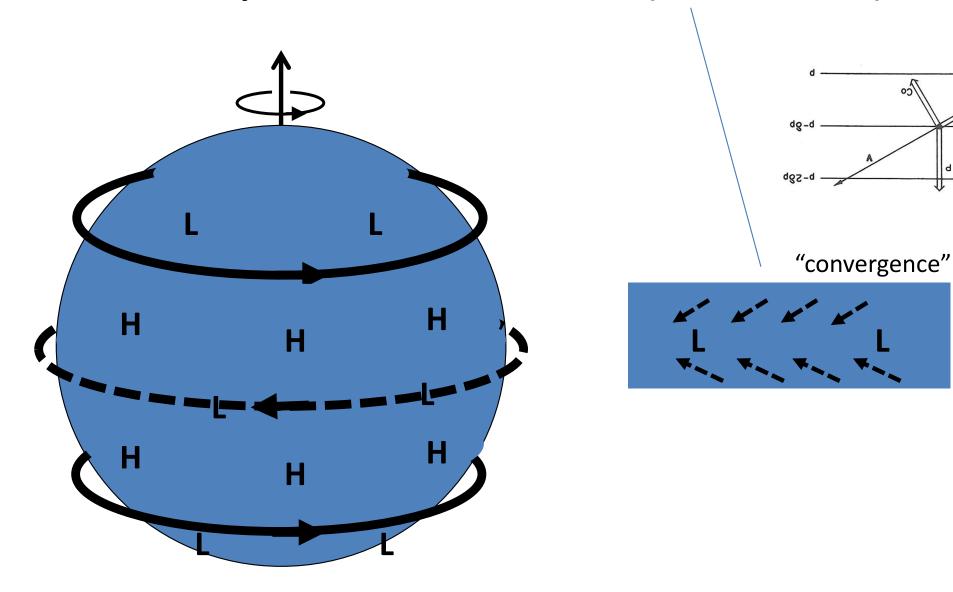
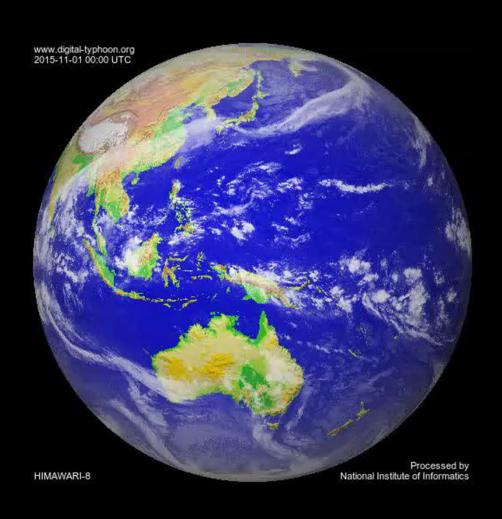


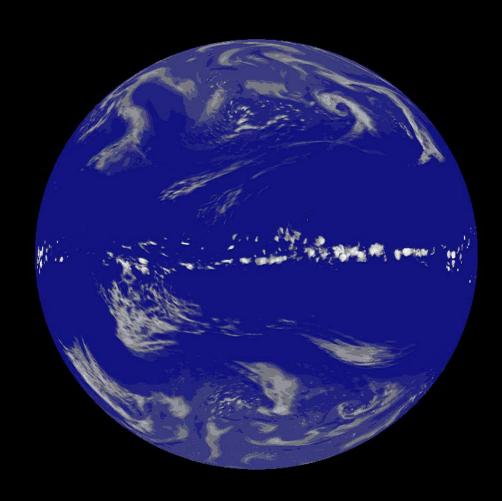
FIGURE 10.5. The mass transport of the Ekman layer is directed to the right of the wind in the northern hemisphere (see Eq. 10-5). Theory suggests that horizontal currents,  $u_{ag}$ , within the Ekman layer spiral with depth as shown.

## Geostrophic and surface (frictional) flow



## Earth \( \neq \text{"Aqua-Planet"} \)





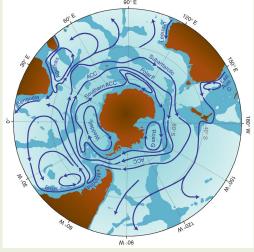
Himawari-8/JMA (1-30 Nov 2015)

NICAM/JAMSTEC (Satoh et al., 2008)

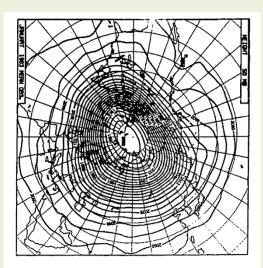
"stripes" of deep atmosphere of Jupiter



circumpolar ocean current around Antarctica

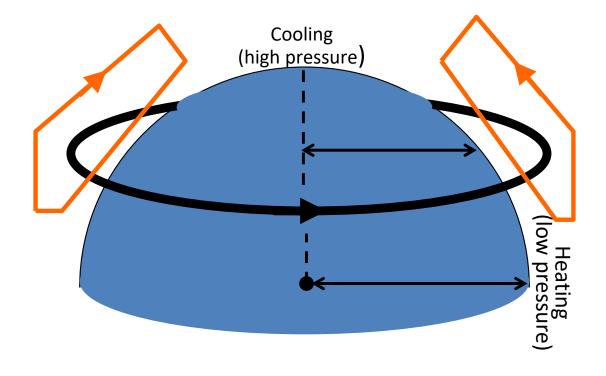


Arctic anticyclone in summer stratosphere



#### "Aqua-planet" zonal vortex

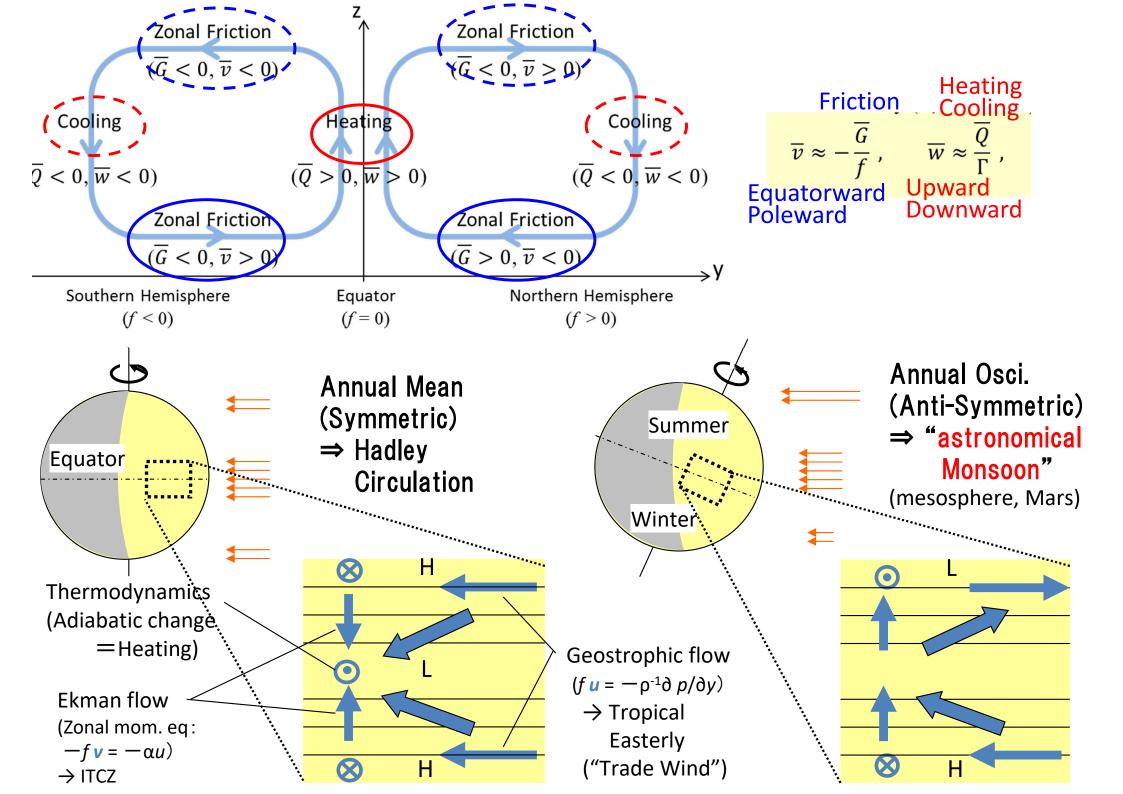
with differential heating



Zonal flow  $(u) \Leftrightarrow \text{Meridional } (y) \text{ pressure gradient}$  $(y\text{-momentum eq. (geostrophic}): f u = -\rho^{-1}\partial \rho/\partial y)$ 

Meridional flow  $(v) \Leftrightarrow \text{Zonal } (x) \text{ friction/drag}$  $(x\text{-momentum eq. (Ekman)} : -f v = -\alpha u)$ 

Vertical flow  $(w) \Leftrightarrow \text{Radiative/Latent heating}$  (thermodynamic eq. (pseudoadiabatic) :  $\Gamma w = \alpha' \Delta T$ )



## **Hermann Ludwig Ferdinand von Helmholtz (1821 – 1894)**



Quasi-2D (horizontal) flow velocity:

$$\boldsymbol{u} = \nabla \times (0, 0, \phi) - \nabla \chi$$

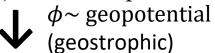
 $(\phi: stream function; \chi: velocity potential)$ 

$$\nabla \cdot (\phi_y, -\phi_x, 0) = 0$$
  
non-divergent  
(solenoidal)

 $\nabla \times (-\nabla \chi) = 0$  irrotational

#### "vortex"

on (quasi-)horizontal plane



Weather maps

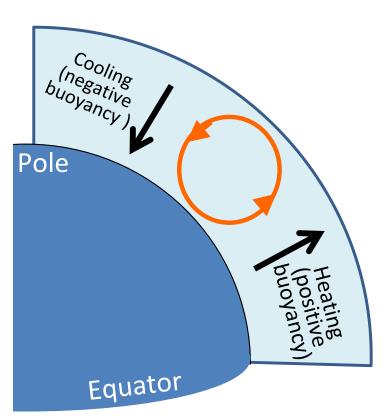
"convection" on vertical plane

Rotation makes narrower equatorial region and baroclinic migra-latitude zone.

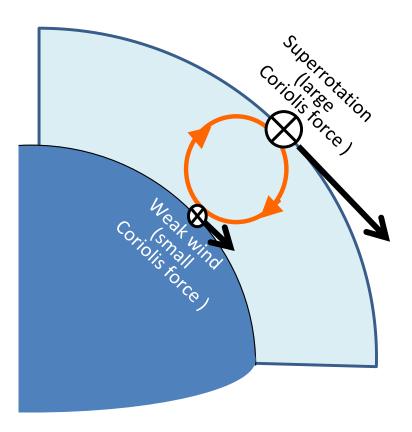
(A) Cooling Heating Cooling Cooling High pressure (C) Hadley cells (Wallace, 2004)

### Direct / indirect circulations due to differential heating / shear satisfying by meridional temperature gradient

(a) Buoyancy torque



(b) Centrifugal/Corioris torque



(adapted from Matsuda & Yoden, 1985)

#### Solar differential heating and monsoon circulation

	Global differential heating	Local heat capacity contrast
Rotation (Diurnal)	Day-Night circulation (Tides)	Sea-Land (Mountain-valley) breeze
Revolution (Annual)	Summer-winter circulation (hemispheric anti phase) Perihelion-Aphelion (in phase)	Ocean-Continent (Monsoon)

Monsoon is a seasonal cycle of wind, generated by the solar radiation through the following two reasons:

- (i) Astronomical (planetological): Hemispheric differential heating, driving a global meridional circulation. The dominat component of solar heating is equatorially symmetric, which drives the Hadley circulation. Seasonal variation of solar radiation on a planet is generated by the following two reasons:
  - Eccentricity if the planetary orbit: globally in-phase, which does not contribute to monsoon.
  - Inclination of the planetary rotation axis: anti-phase between the hemispheres, contributing to monsoon. Diurnal cycle of (day-night) hemispheric differential heating of solar radiation may generate thermal tides.
- (ii) Geographical (terrestrial): Continent-ocean differential heating, driving a flow crossing the coastline.

  Land (solid) sea (liquid) heat capacity contrast, generating a diurnal cycle of sea-land breeze circulation.

  Integration (residue) of diurnal cycle generates the seasonal cycle and thus the monsoon.

  Wind from ocean to continent (in summer hemisphere) brings moisture, and therefore generates rainy season.

  The cloud-precipitation water cycle makes latent heat transport, and feedbacks to monsoon enhancement.

#### **Edmond Halley (1656 –1742)**

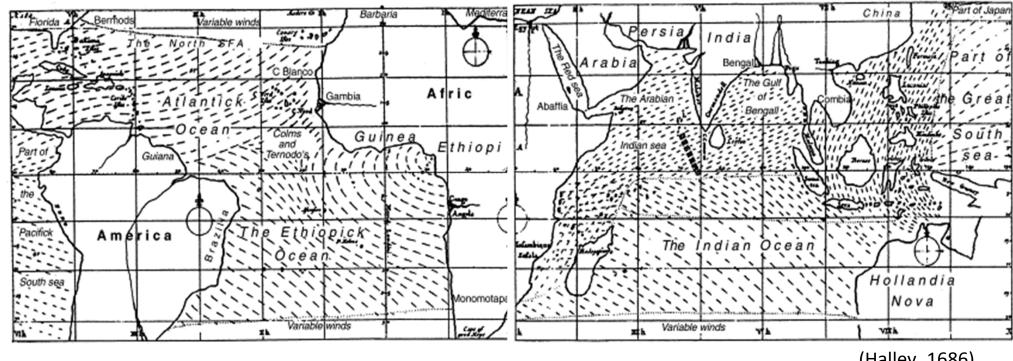


(http://www.staff.science.uu.nl/~gent0113/astrology/newton.htm)

#### **George Hadley (1685 – 1768)**

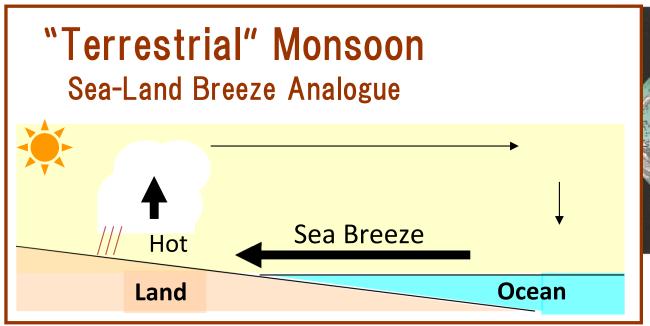


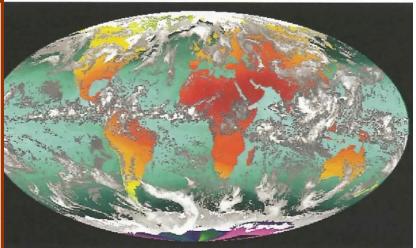
(Hadley, 1735; reproduced by Lorenz, 1967)



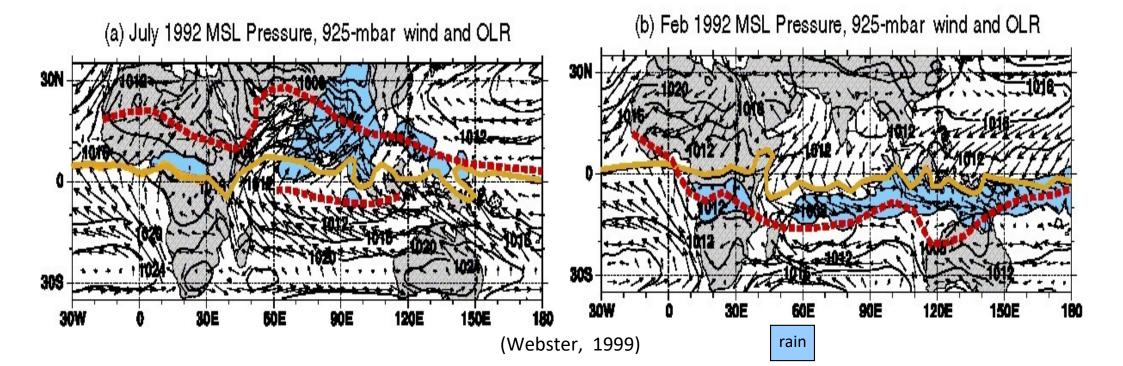
(Halley, 1686)

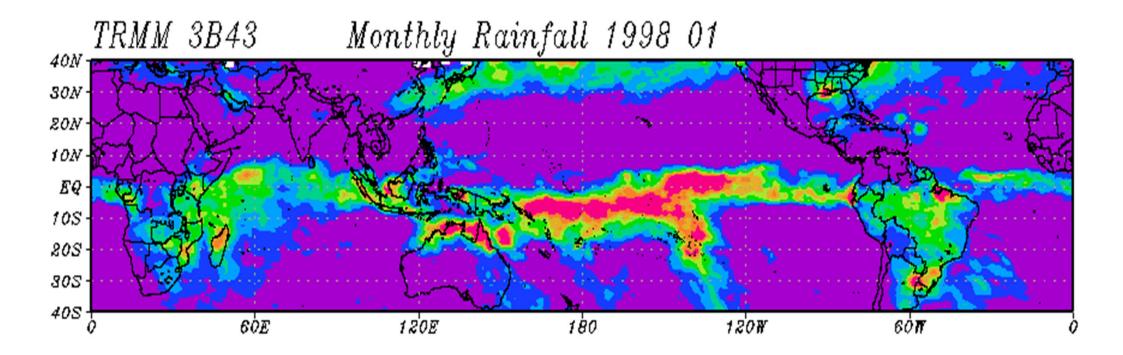
#### **Ocean-Land Contrasts and Monsoon**

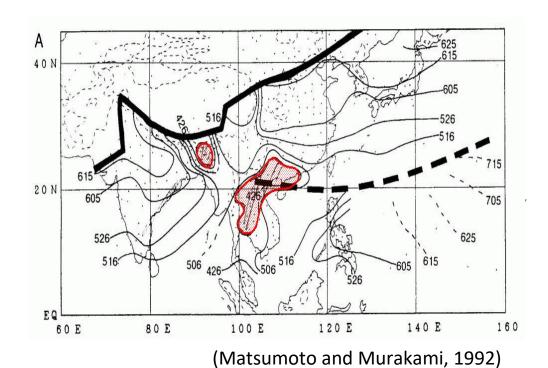


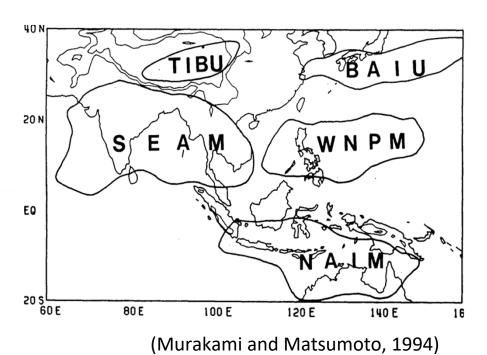


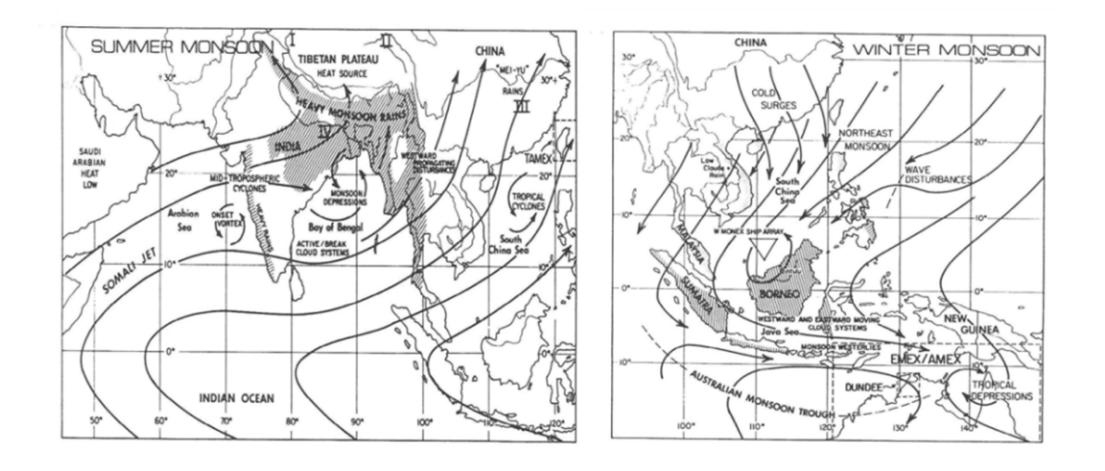
(Wallace & Hobbs, 2006; originally from UWSSEC)

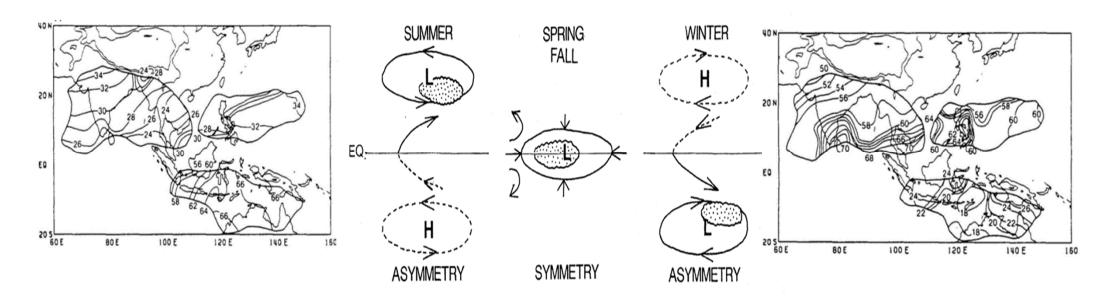




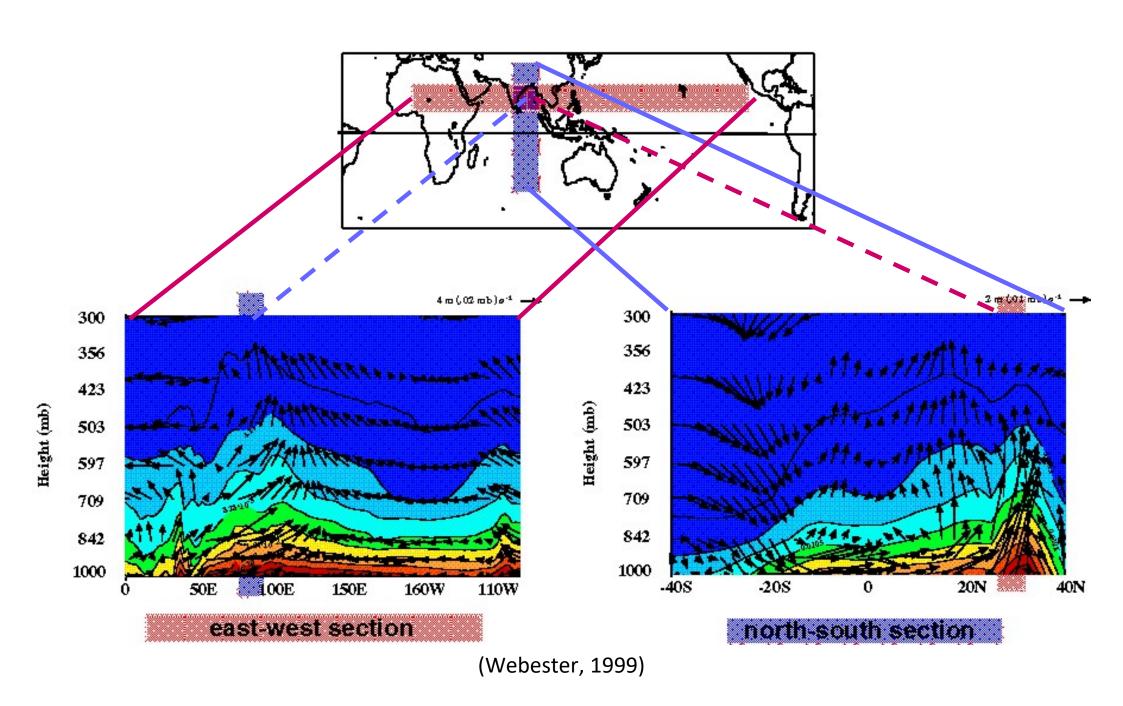




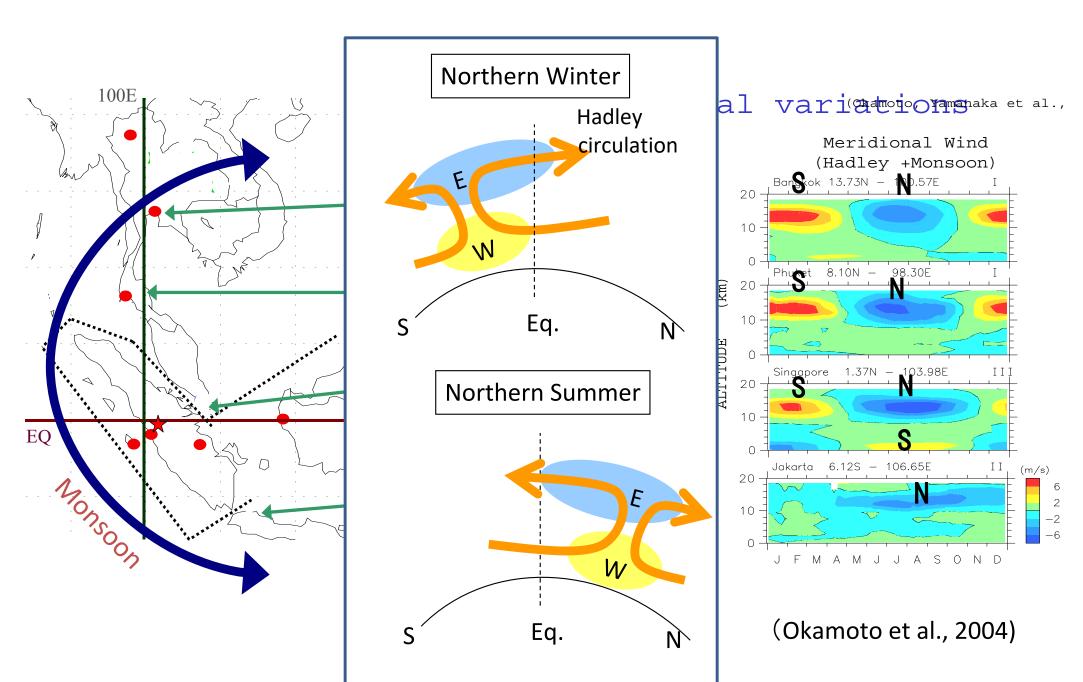




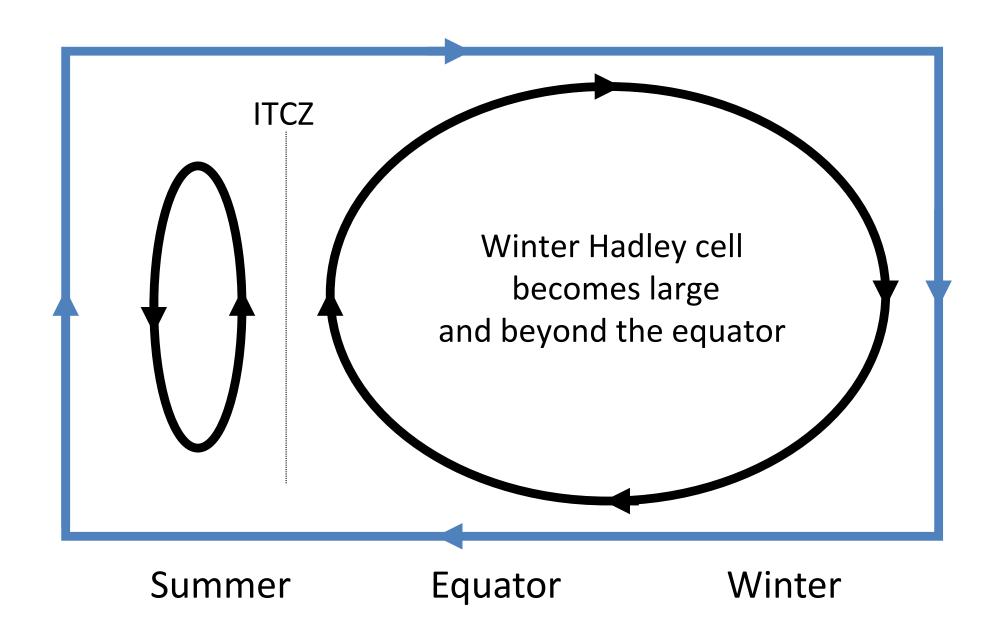
#### Longitudinal / latitudinal sections in NH Summer



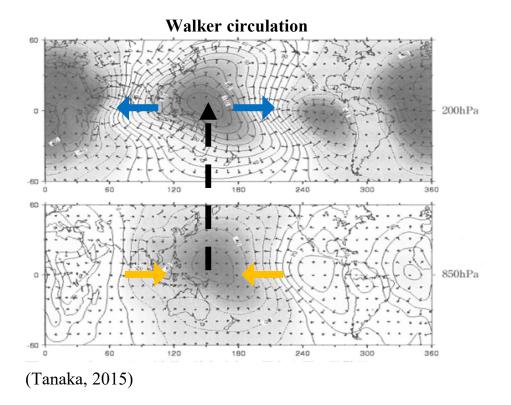
#### 100E meridian (Sumatera-Malay-Thailand) obs.

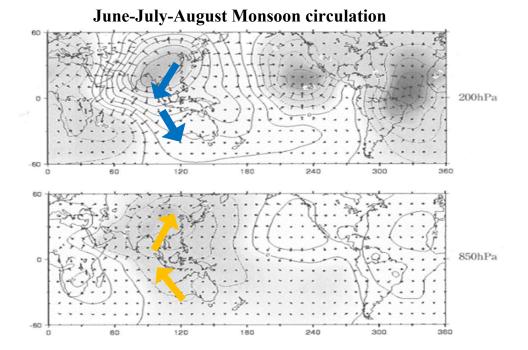


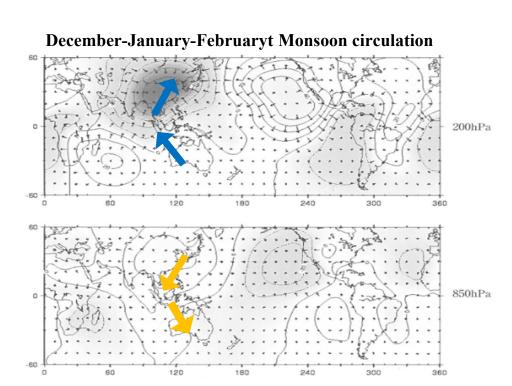
#### **Superimposing Monsoon and Hadley circulations**



#### Heating-induced zonal (Walker) and Summer-Winter (Monsoon) circulations







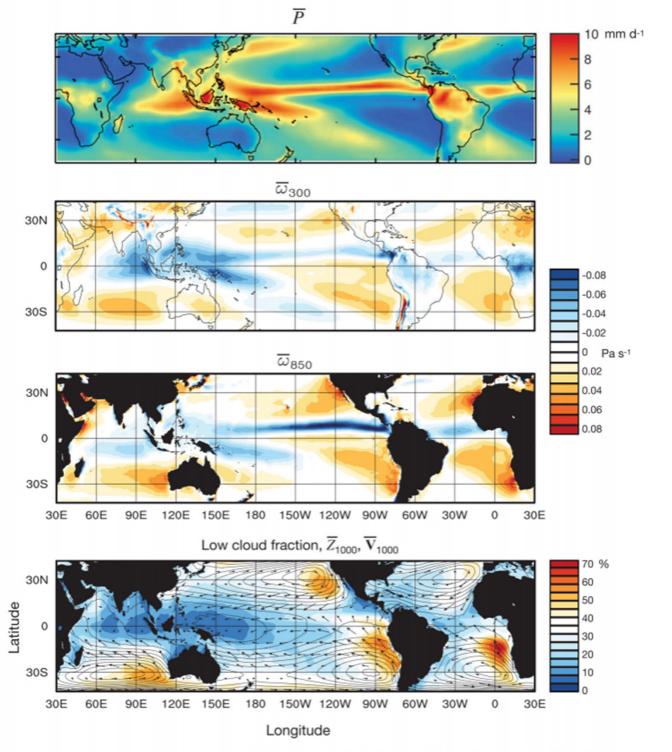
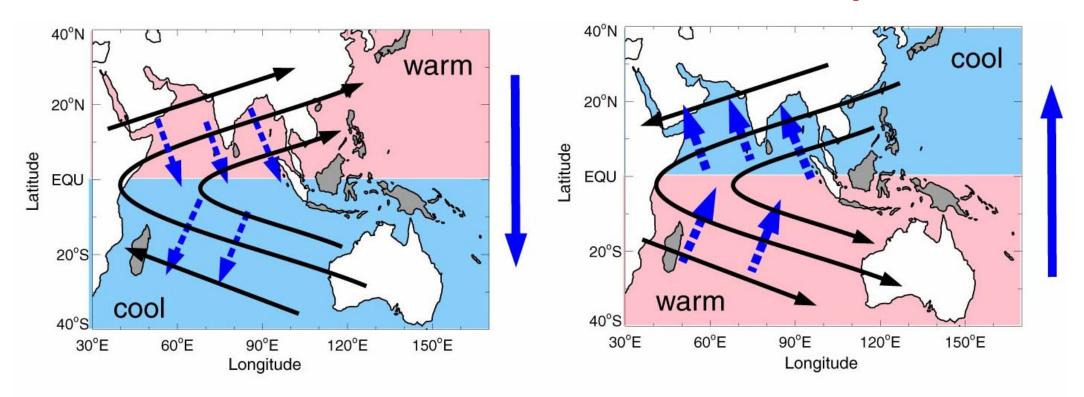


Figure 14.1 Annual mean maps: (top) rain rate; (second) vertical velocity at the 300 hPa level; (third) vertical velocity at the 850 hPa level; and (bottom) the percentage of the sky covered by low clouds, superposed on the 1000 hPa wind vectors and geopotential height (contour interval 10 m). Note that positive values of ω indicate descent.

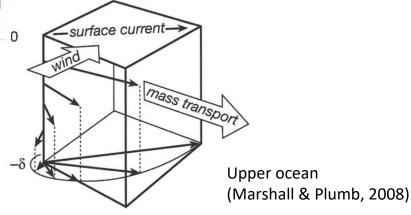
#### **Monsoon-Induced Ocean Heat Transport**

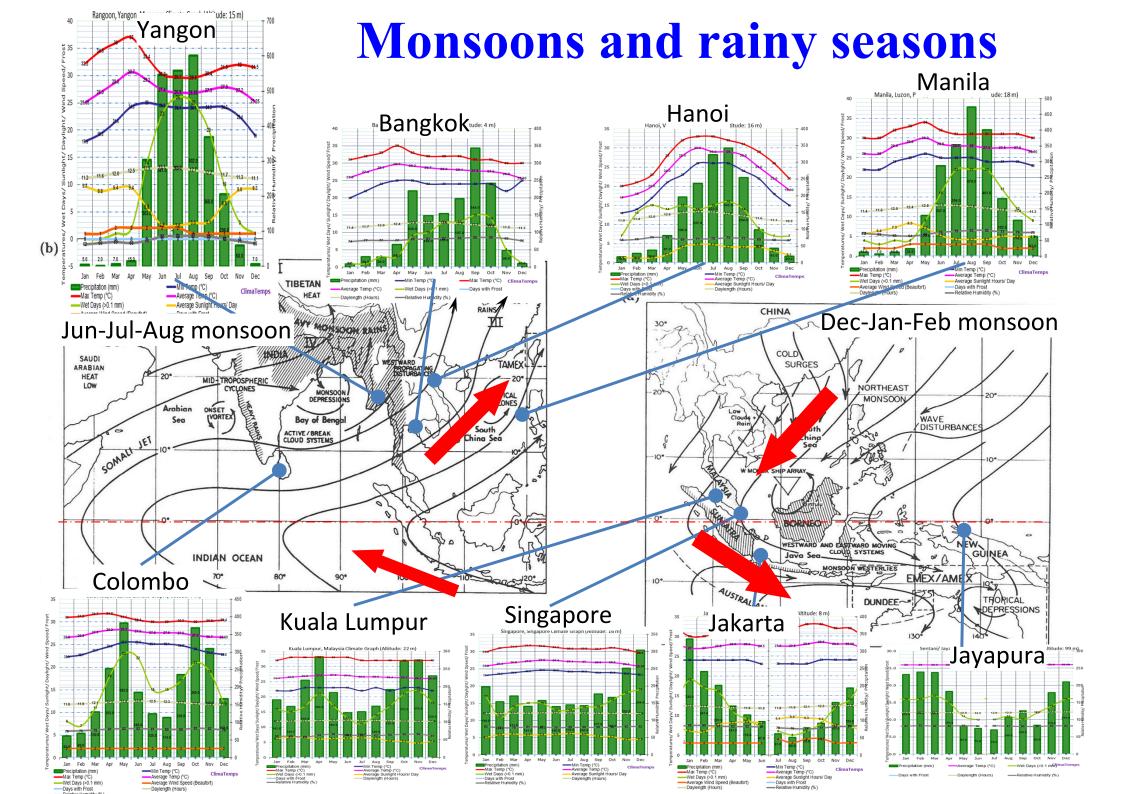


Southward ocean heat transport of 1.5 PW Northward ocean heat transport of 1.5 PW (cools NIO while warming SIO) (Webster, 1999) (cools SIO while warming NIO)

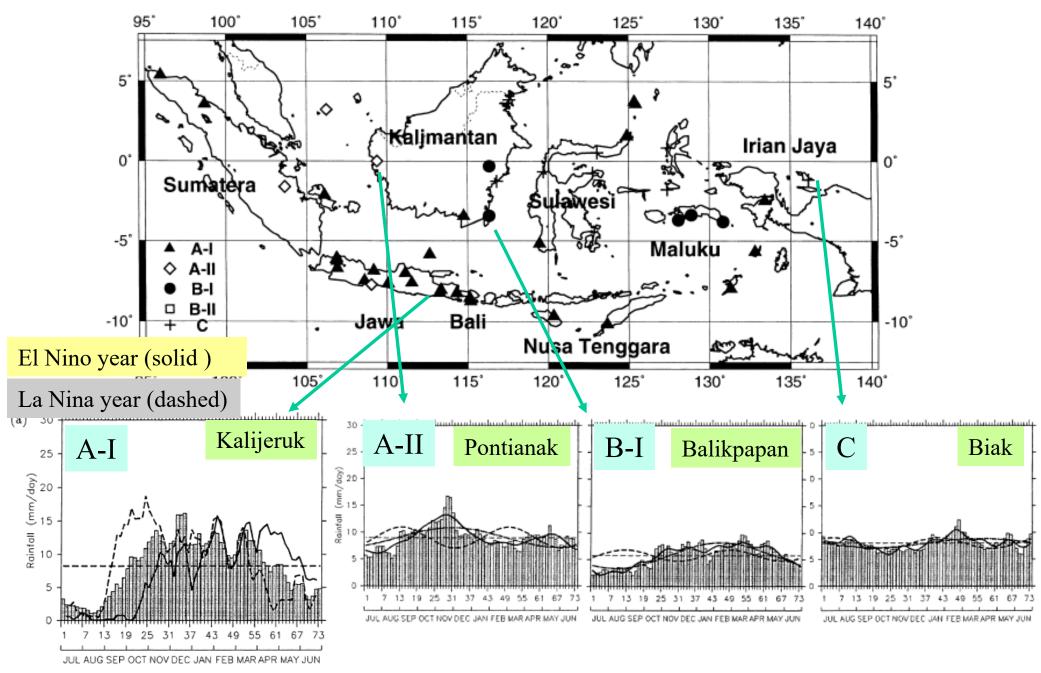
Overall impact of wind-driven Ekman ocean heat transport is to cool the summer hemisphere and warm the winter hemisphere thus reducing the cross-equatorial SST gradient and minimizing seasonal extremes in the monsoon

FIGURE 10.5. The mass transport of the Ekman layer is directed to the right of the wind in the northern hemisphere (see Eq. 10-5). Theory suggests that horizontal currents,  $u_{ag}$ , within the Ekman layer spiral with depth as shown.



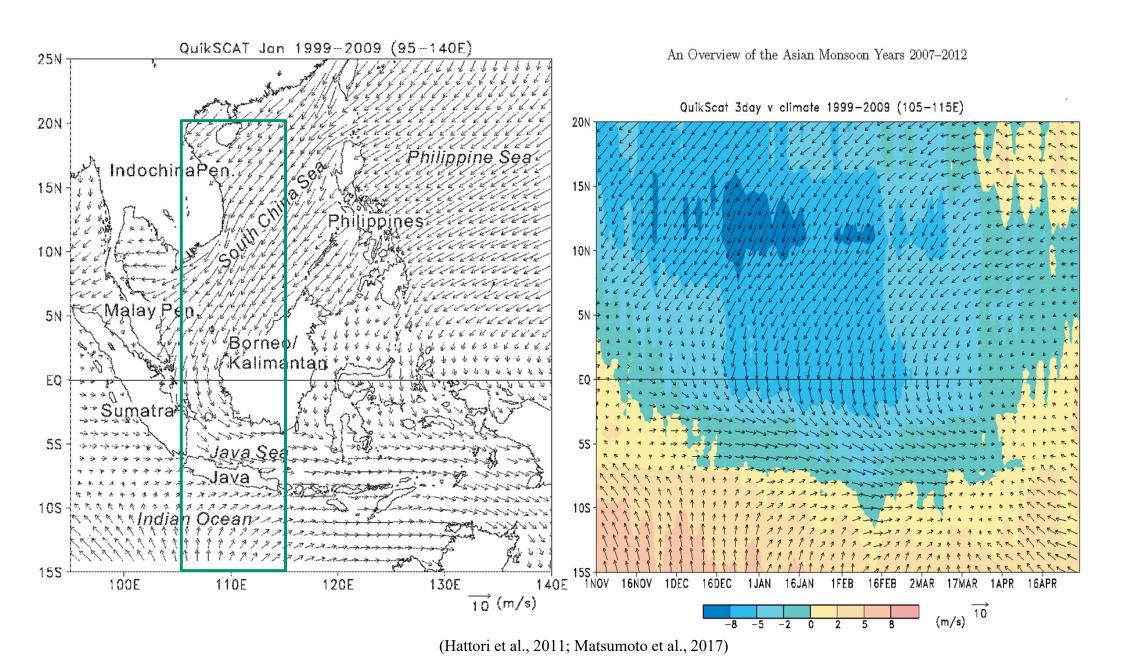


#### Seasonal and interannual variations

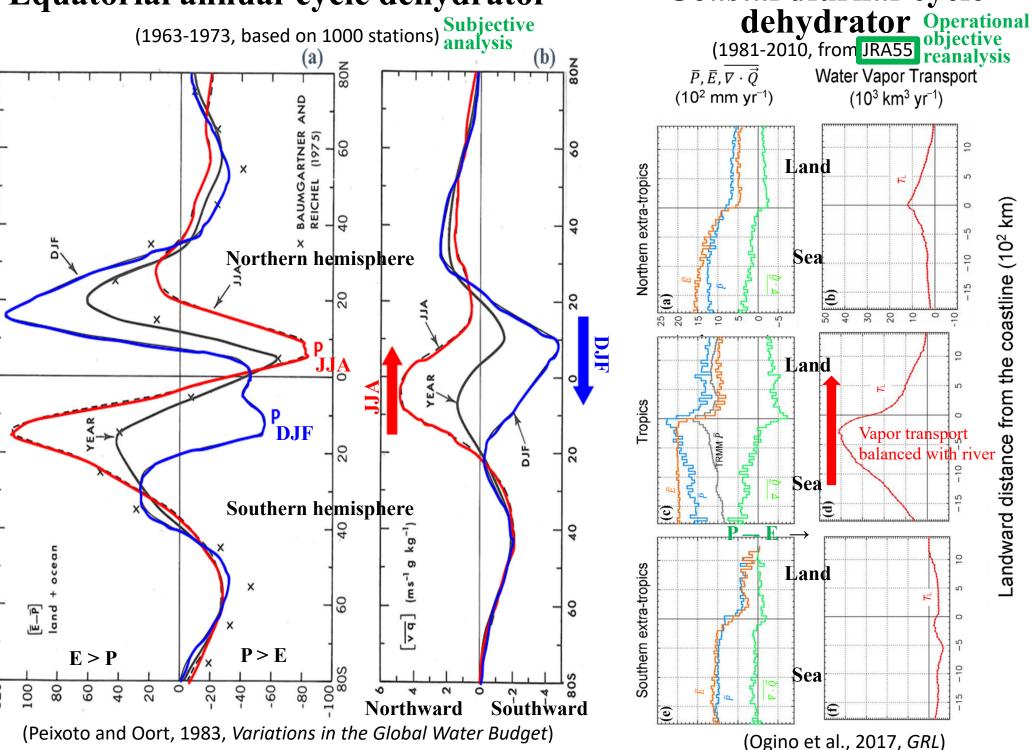


(Hamada, et al., 2002, J. Meteorol. Soc. Japan)

## Cold surge climatology

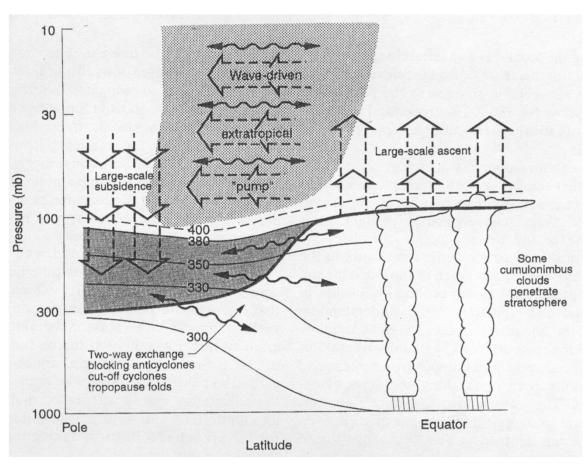


#### **Equatorial annual-cycle dehydrator**



Coastal diurnal-cycle

#### Stratospheric response to tropospheric convection



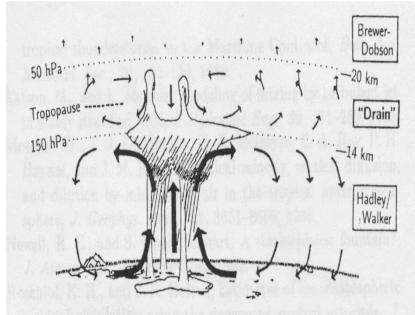
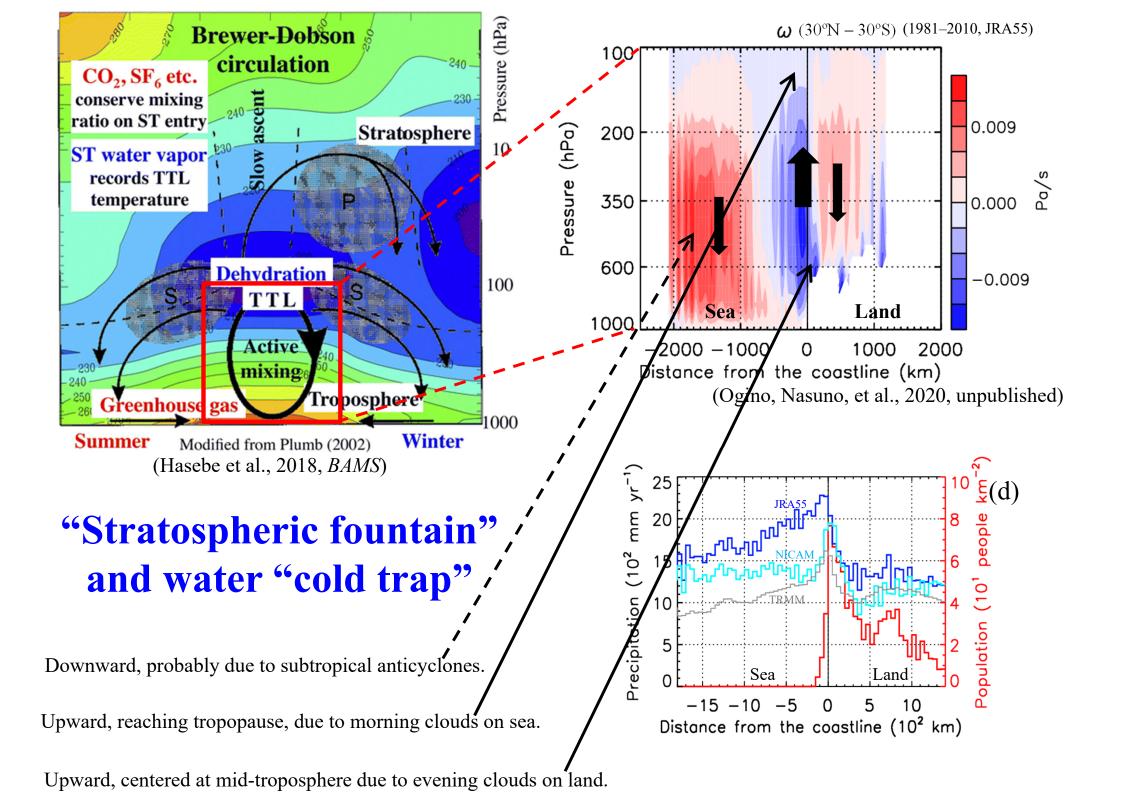


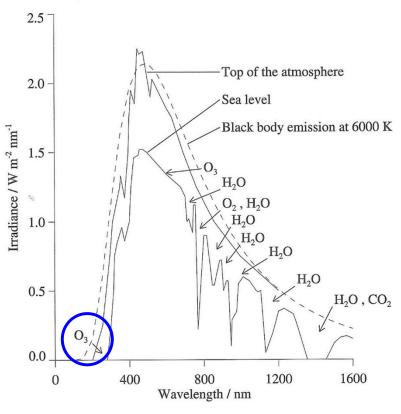
Figure 4. Illustration of the diabatic mean flow (arrows) and its relationship to the locations of deep convection and convective outflows.

(Holton, 1995)

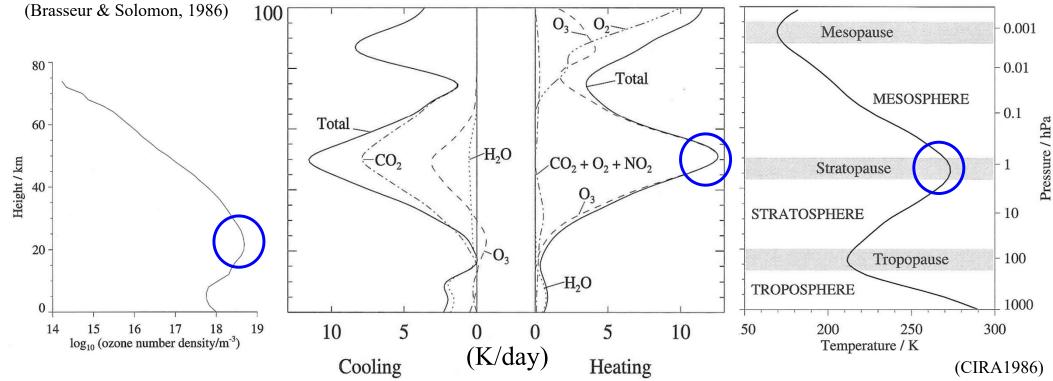
(Sherwood, 2000)

cf. Ogino's ozone observations at Hanoi





# Ozone UV heating and Middle Atmosphere (Mesosphere & Stratosphere)



## Meridional distribution of Temperature and Zonal Wind

(January, CIRA1986)

